

# Bekenstein-Hawking entropy and strange metals

Quantum Information in Quantum Gravity II  
Perimeter Institute, Waterloo  
August 20, 2015

Subir Sachdev

Talk online: [sachdev.physics.harvard.edu](http://sachdev.physics.harvard.edu)



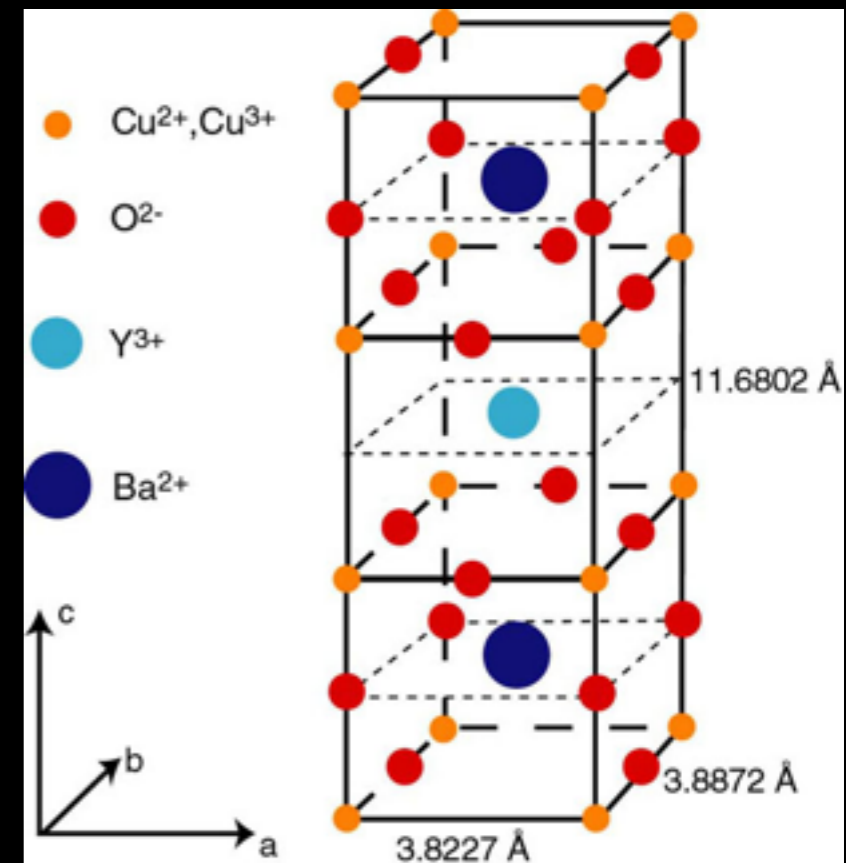
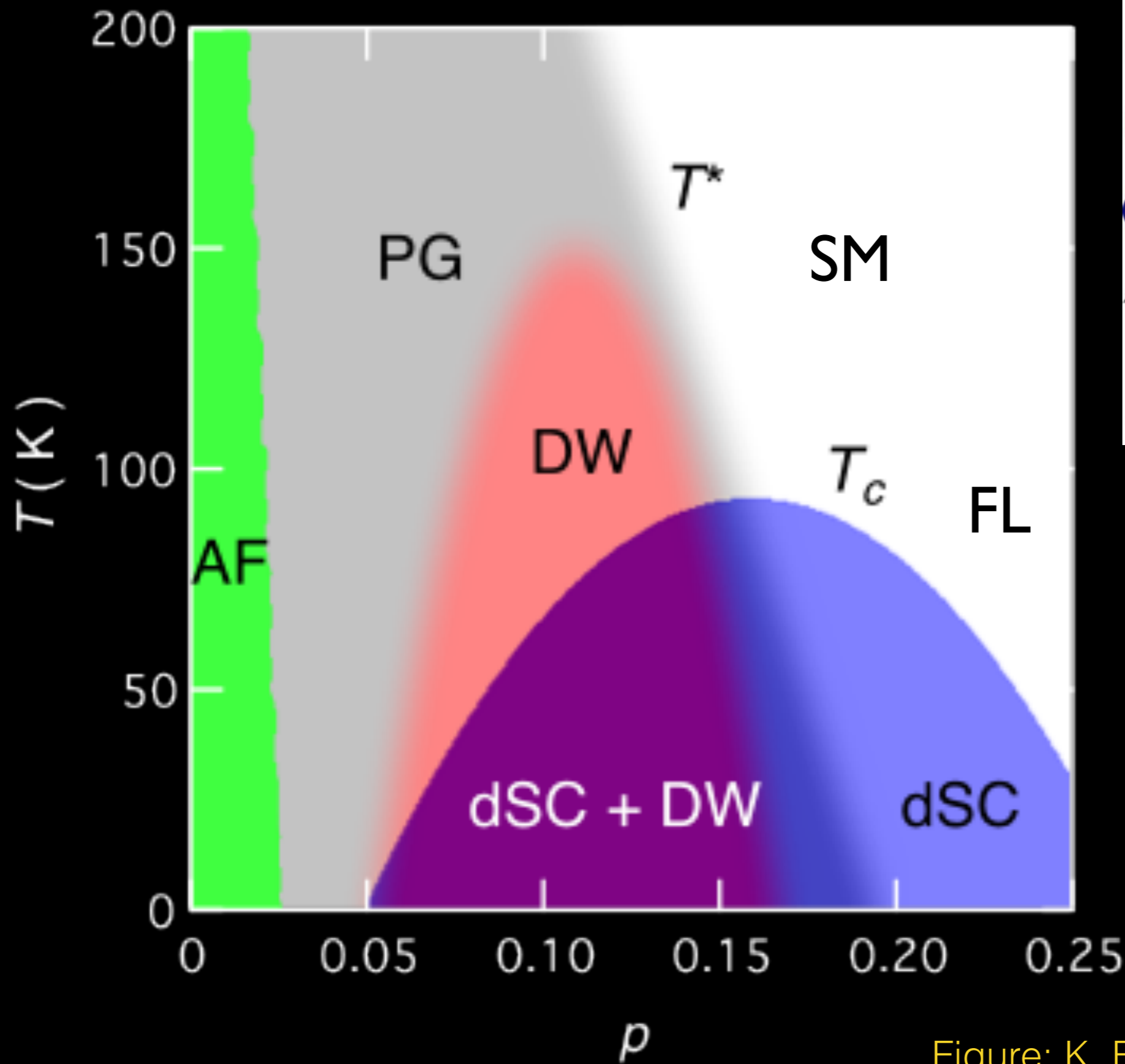
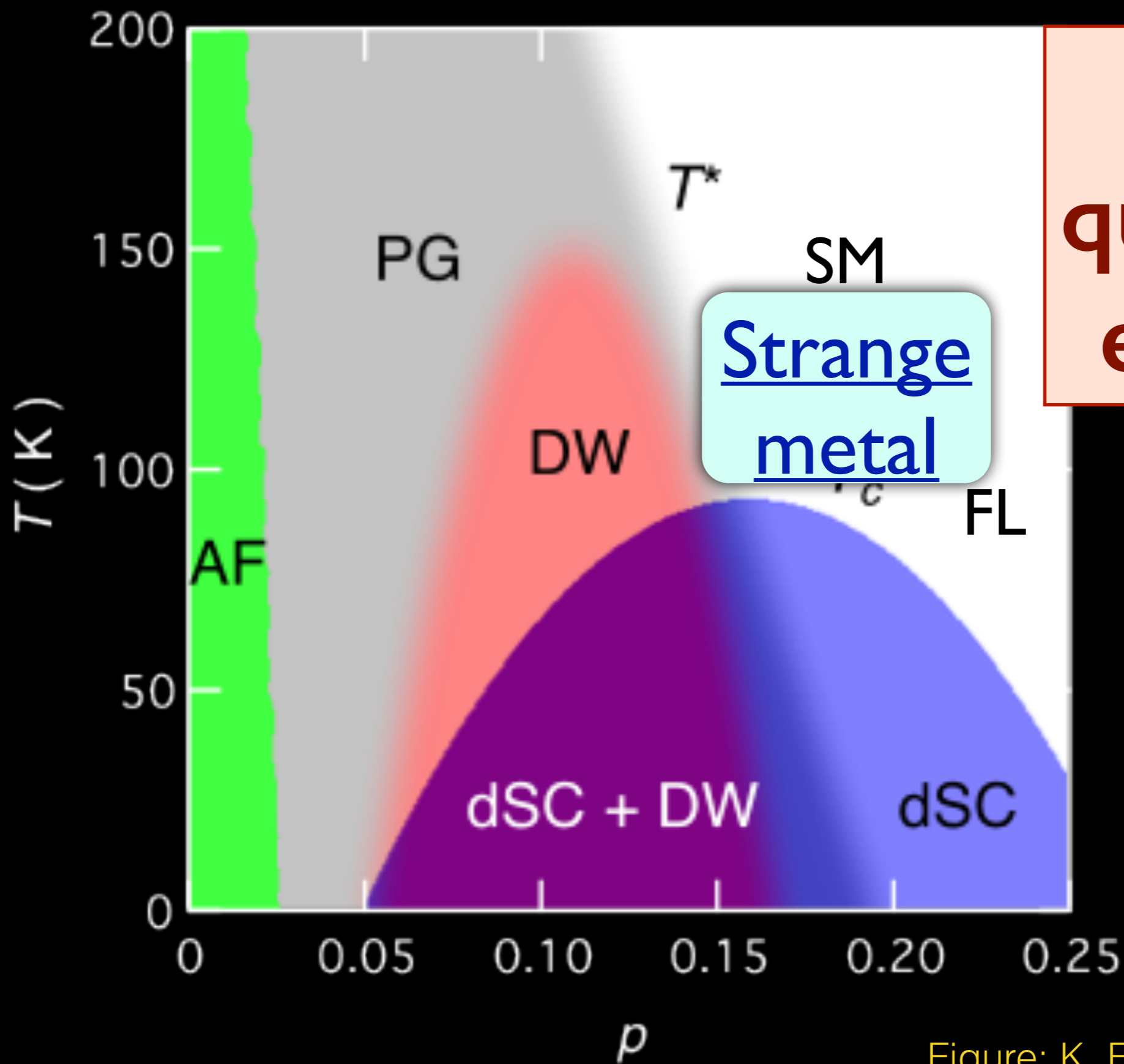


Figure: K. Fujita and J. C. Seamus Davis

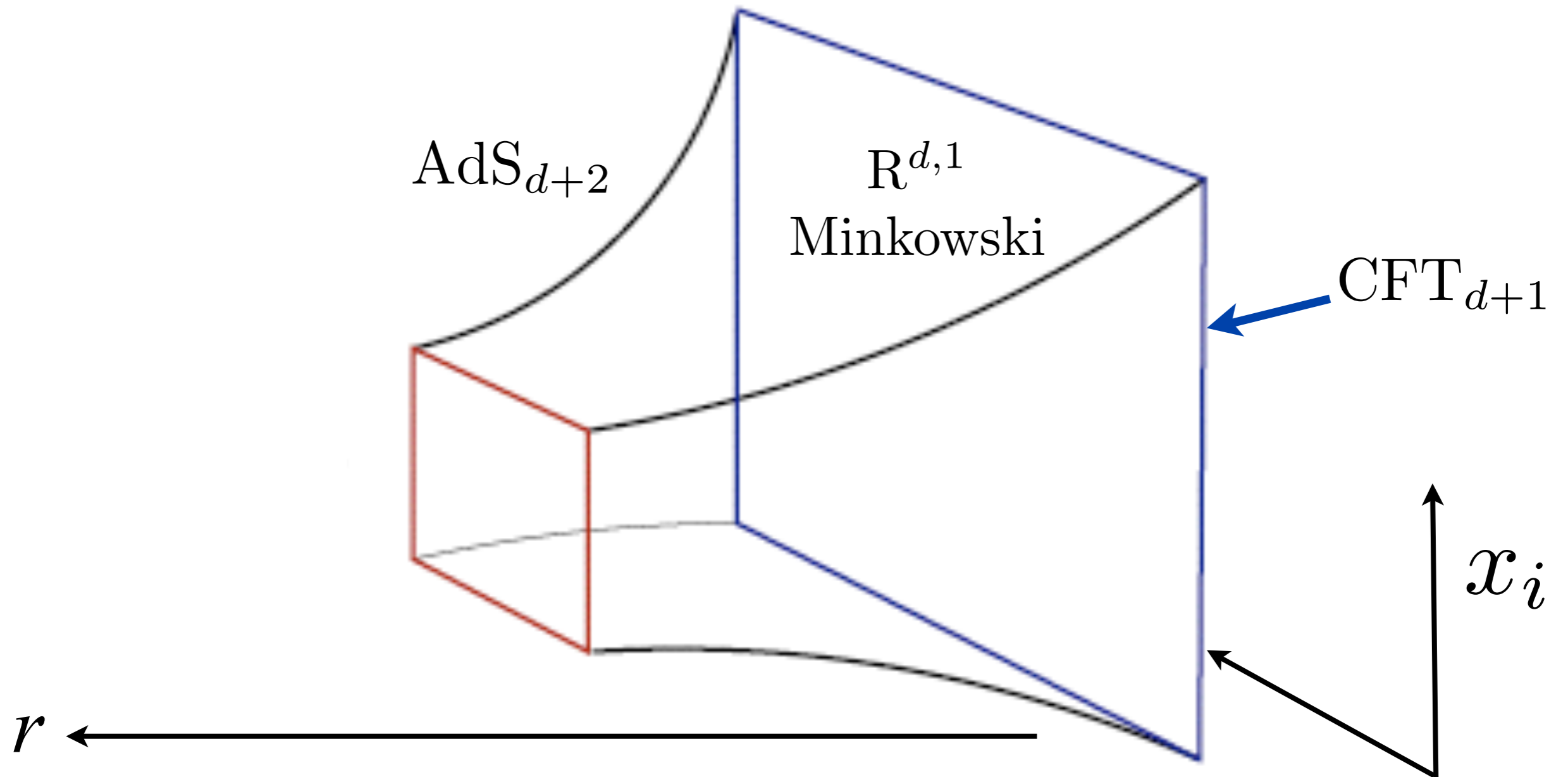


**No  
quasiparticle  
excitations**

Figure: K. Fujita and J. C. Seamus Davis

# AdS/CFT correspondence at zero temperature

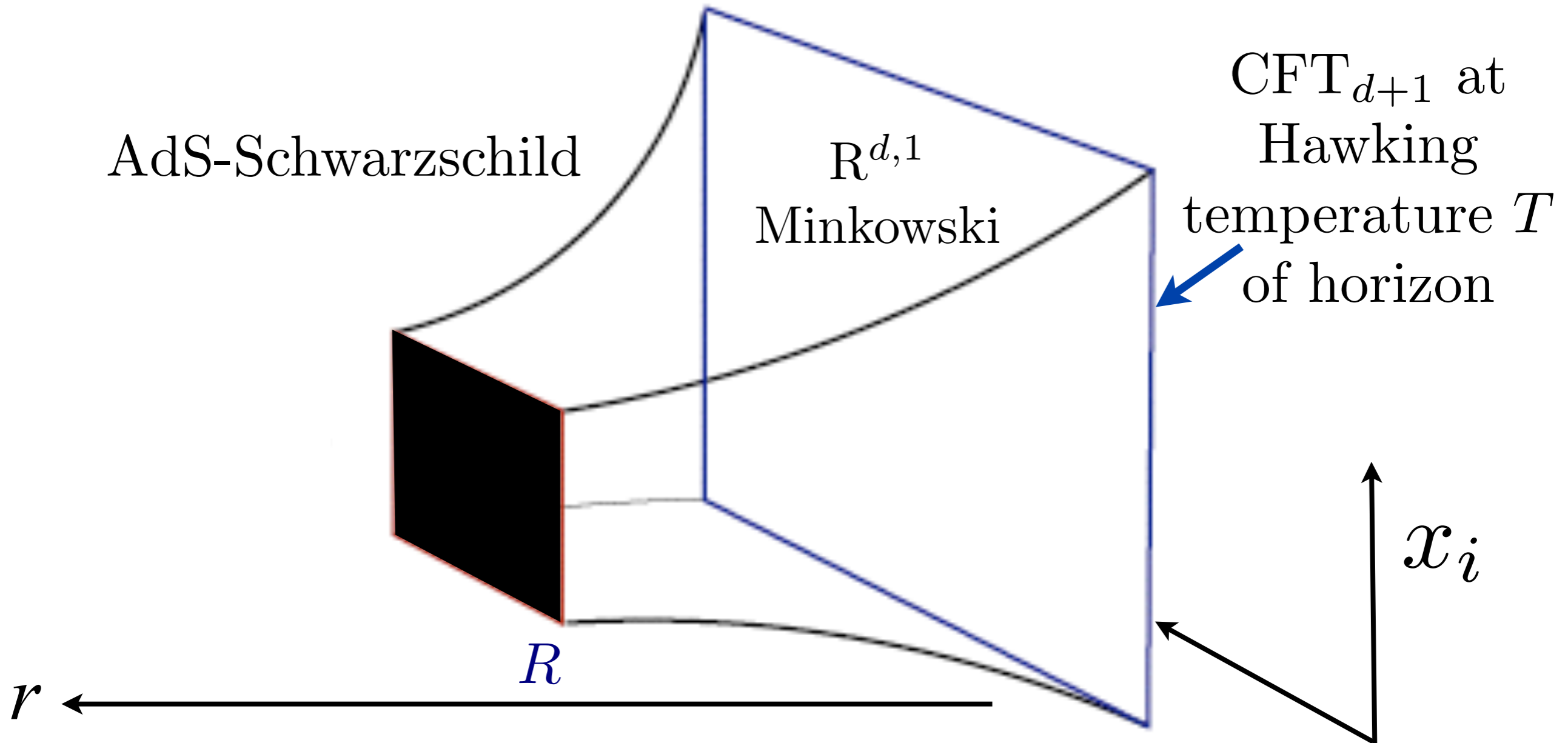
Einstein gravity  $\mathcal{S}_E = \int d^{d+2}x \sqrt{-g} \left[ \frac{1}{2\kappa^2} \left( \mathcal{R} + \frac{d(d+1)}{L^2} \right) \right]$



$$ds^2 = \left( \frac{L}{r} \right)^2 [dr^2 - dt^2 + d\vec{x}^2]$$

# AdS/CFT correspondence at non-zero temperature

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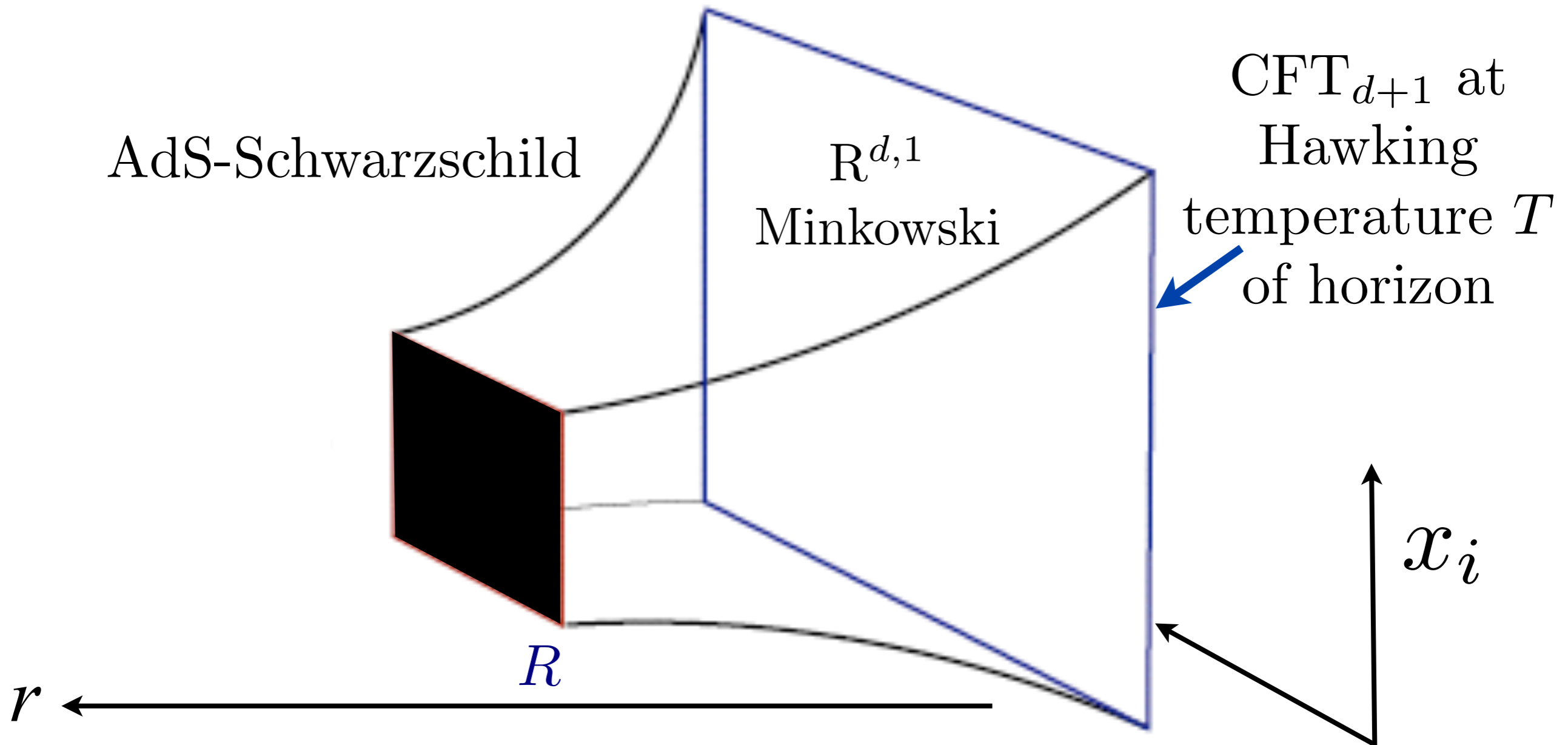


$$ds^2 = \left( \frac{L}{r} \right)^2 \left[ \frac{dr^2}{f(r)} - f(r) dt^2 + d\vec{x}^2 \right]$$

with  $f(r) = 1 - (r/R)^{d+1}$  and  $T = (d+1)/(4\pi R)$ .

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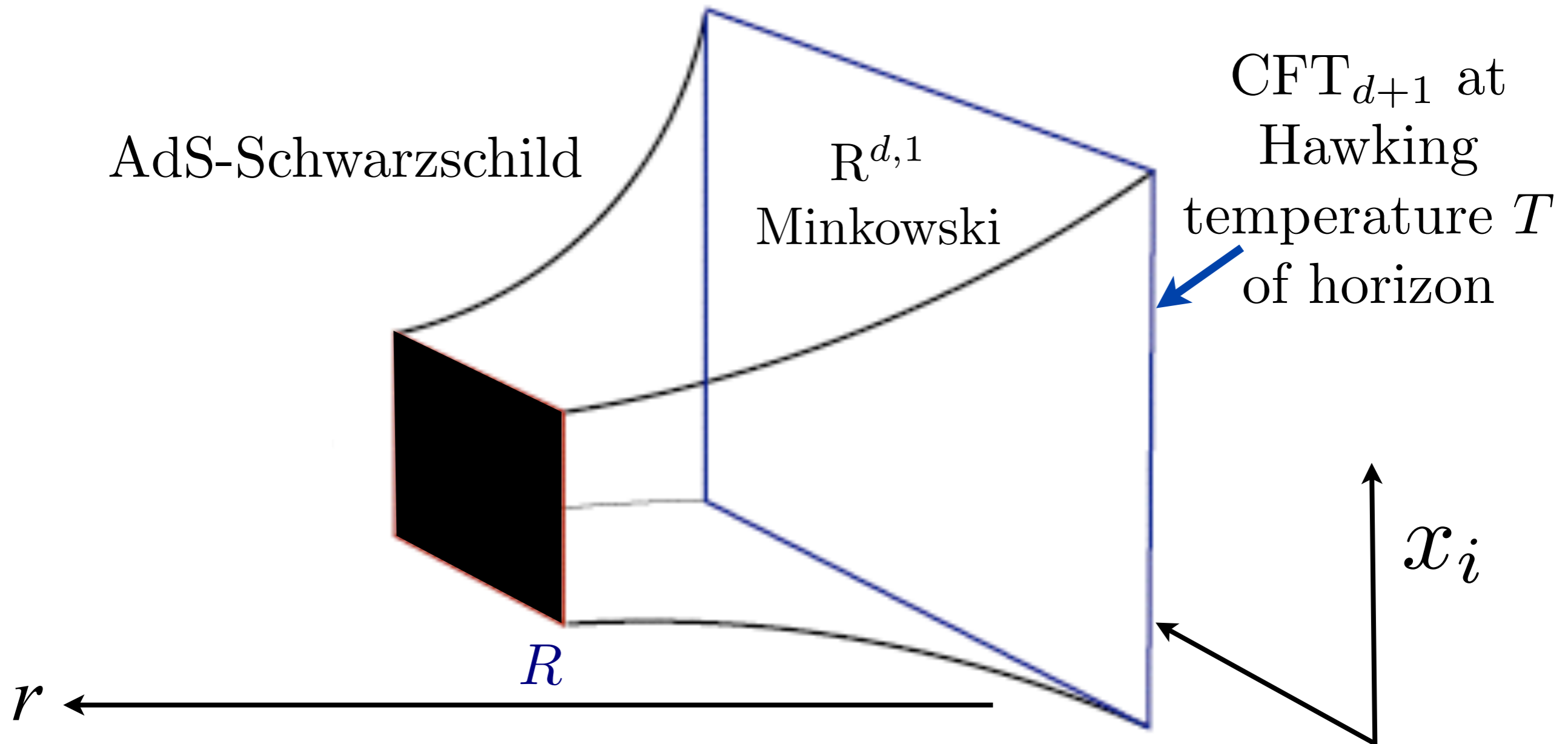


Entropy density of  $\text{CFT}_{d+1}$ ,  $\mathcal{S} \sim T^d$

Bekenstein-Hawking entropy density,  $\mathcal{S}_{\text{BH}} \sim T^d$

# AdS/CFT correspondence at non-zero temperature

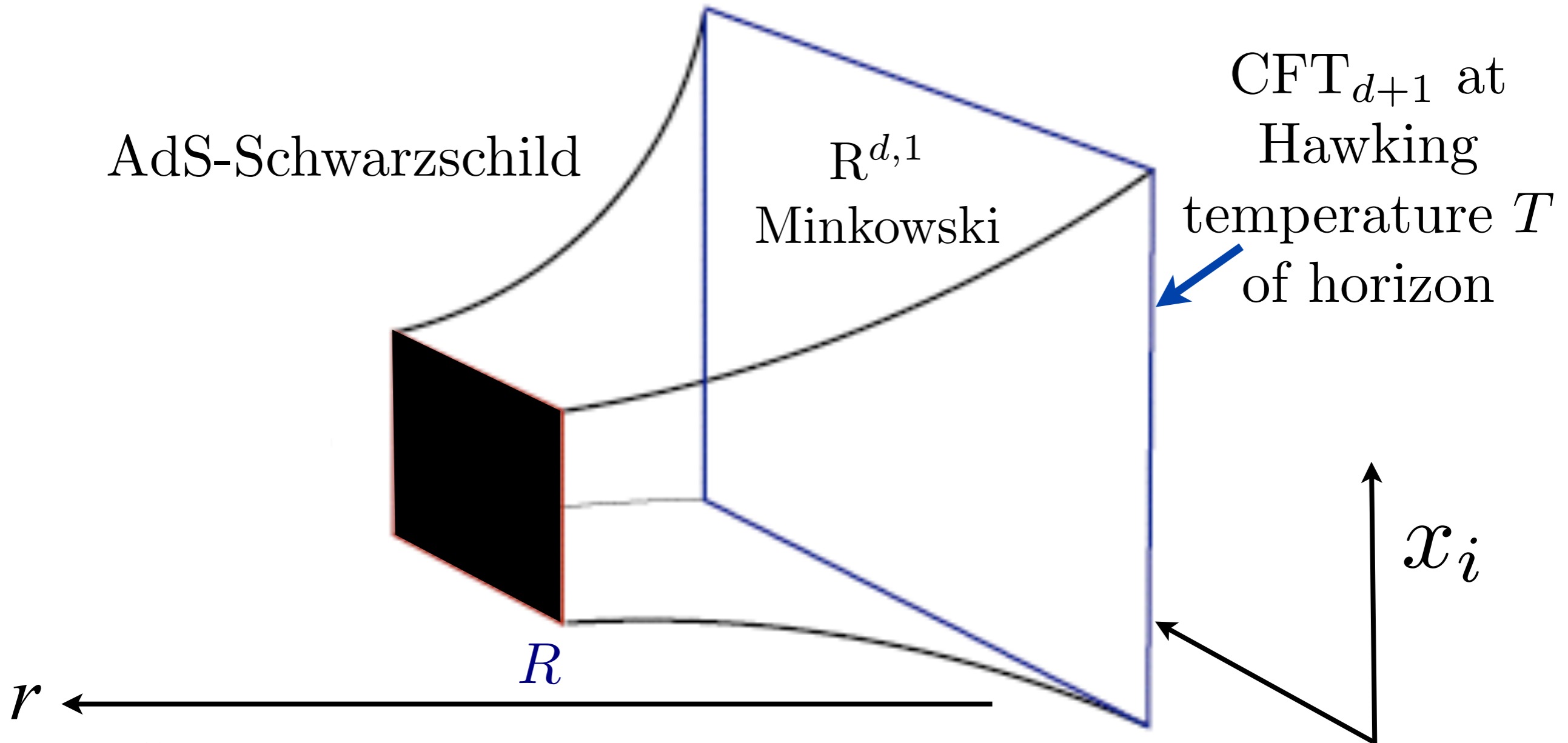
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For  $\text{SU}(N)$  SYM in  $d = 3$ ,  $\mathcal{S}_{\text{BH}} = (\pi^2/2)N^2T^3$ . But there is (still) no confirmation of this from a field-theory computation on SYM.

# AdS/CFT correspondence at non-zero temperature

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Correspondence in  $d = 1$ :

$$\mathcal{S} = \mathcal{S}_{\text{BH}} = \frac{\pi}{3} c T,$$

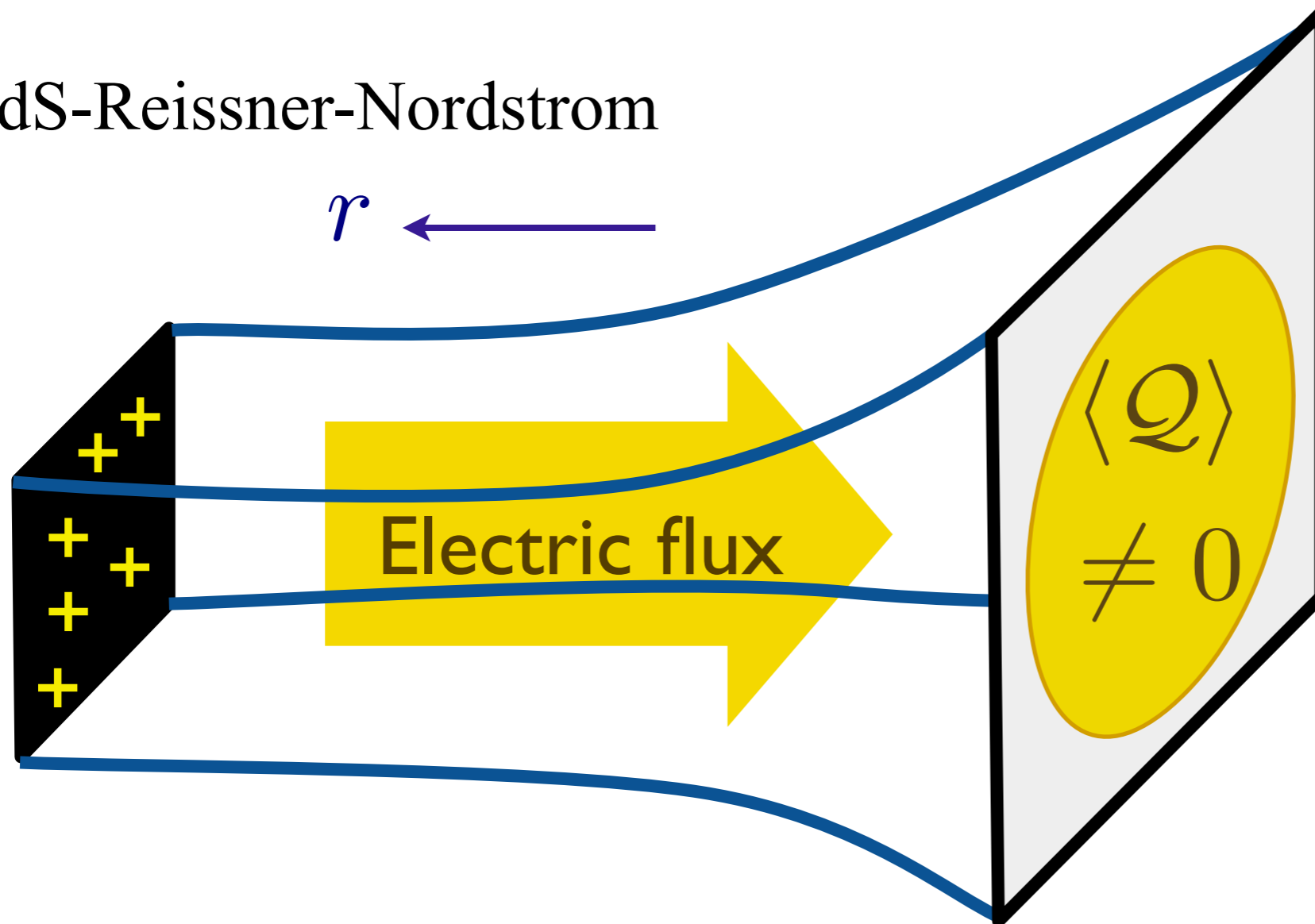
where  $c = 12\pi L/\kappa^2$  is the central charge of the  $CFT_2$ .



# Charged black branes

Einstein-Maxwell theory  $\mathcal{S}_{EM} = \int d^{d+2}x \sqrt{-g} \left[ \frac{1}{2\kappa^2} \left( \mathcal{R} + \frac{d(d+1)}{L^2} - \frac{R^2}{g_F^2} F^2 \right) \right]$

AdS-Reissner-Nordstrom

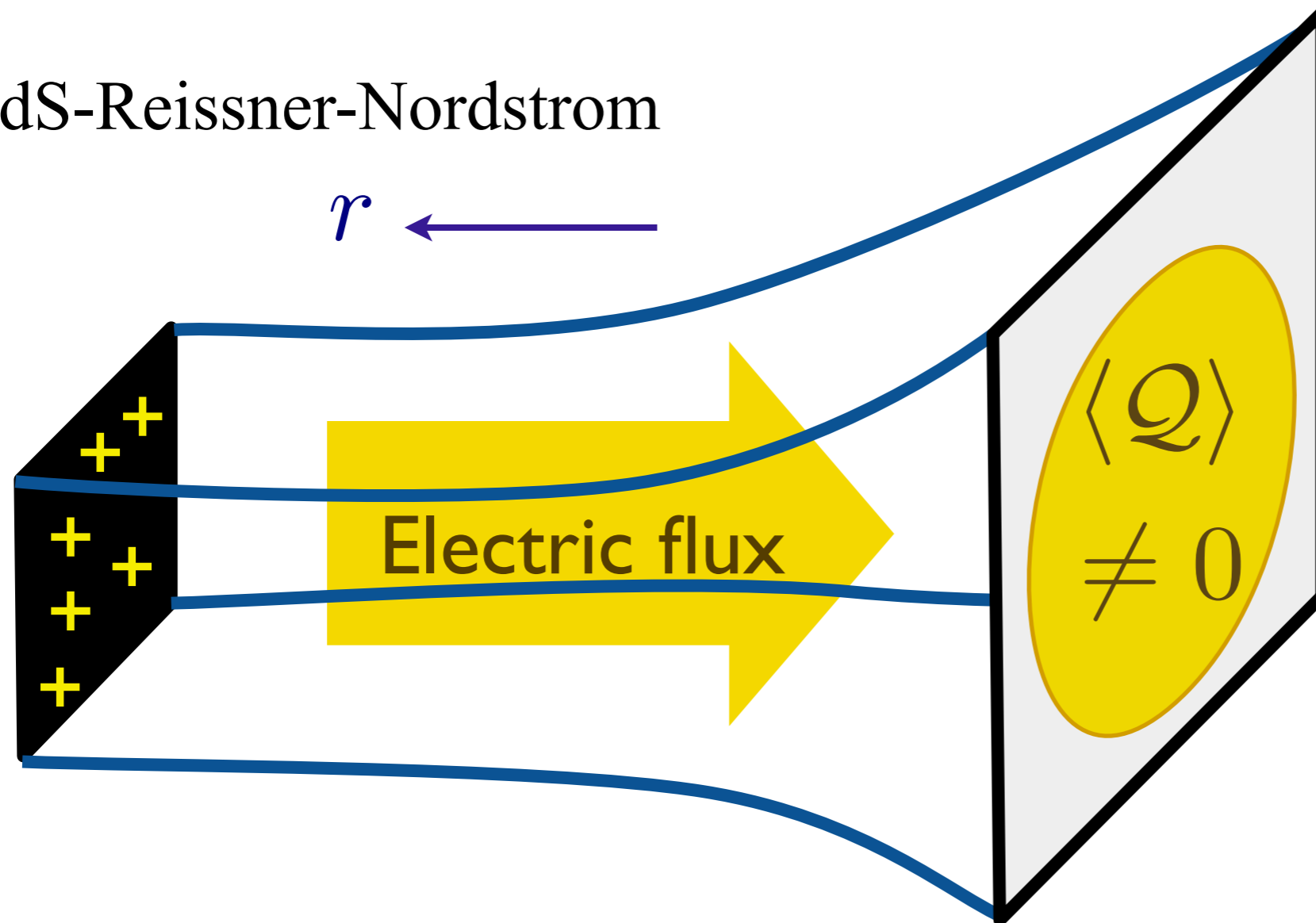


Quantum matter on the boundary with a variable charge density  $\mathcal{Q}$  of a global U(1) symmetry.

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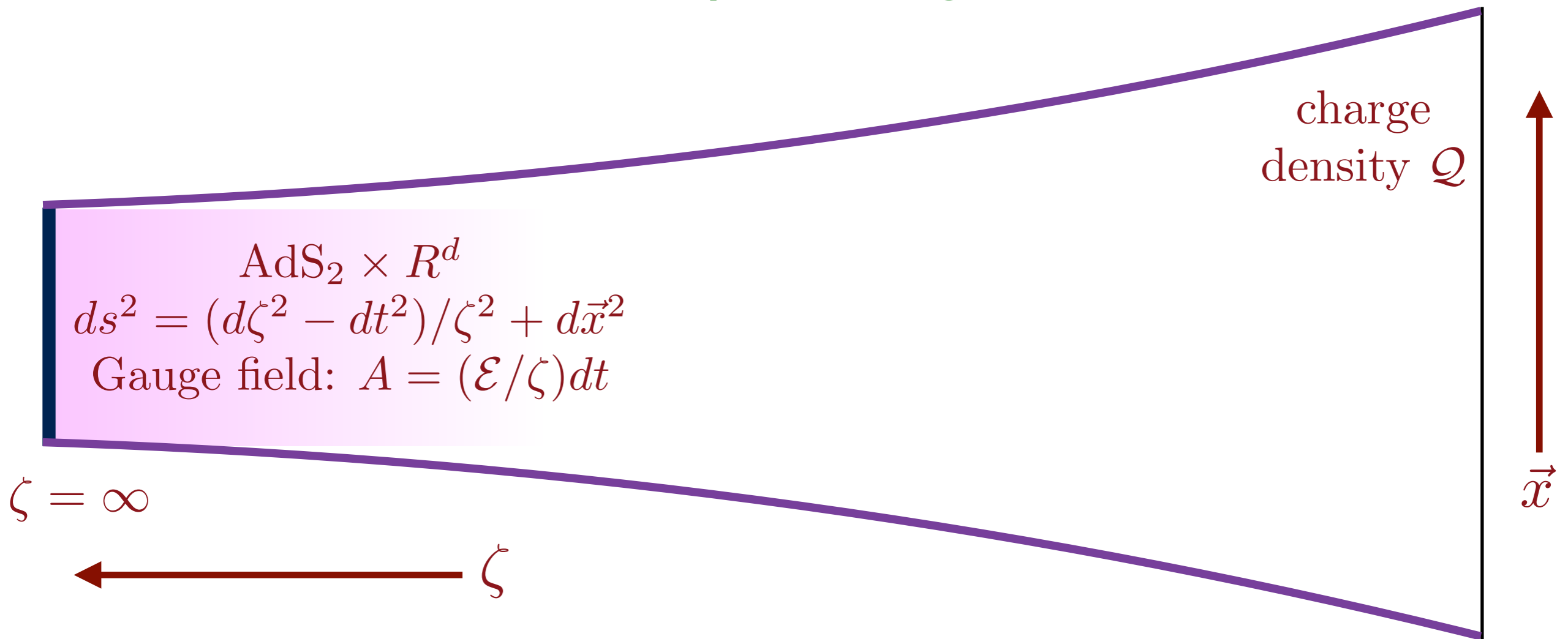
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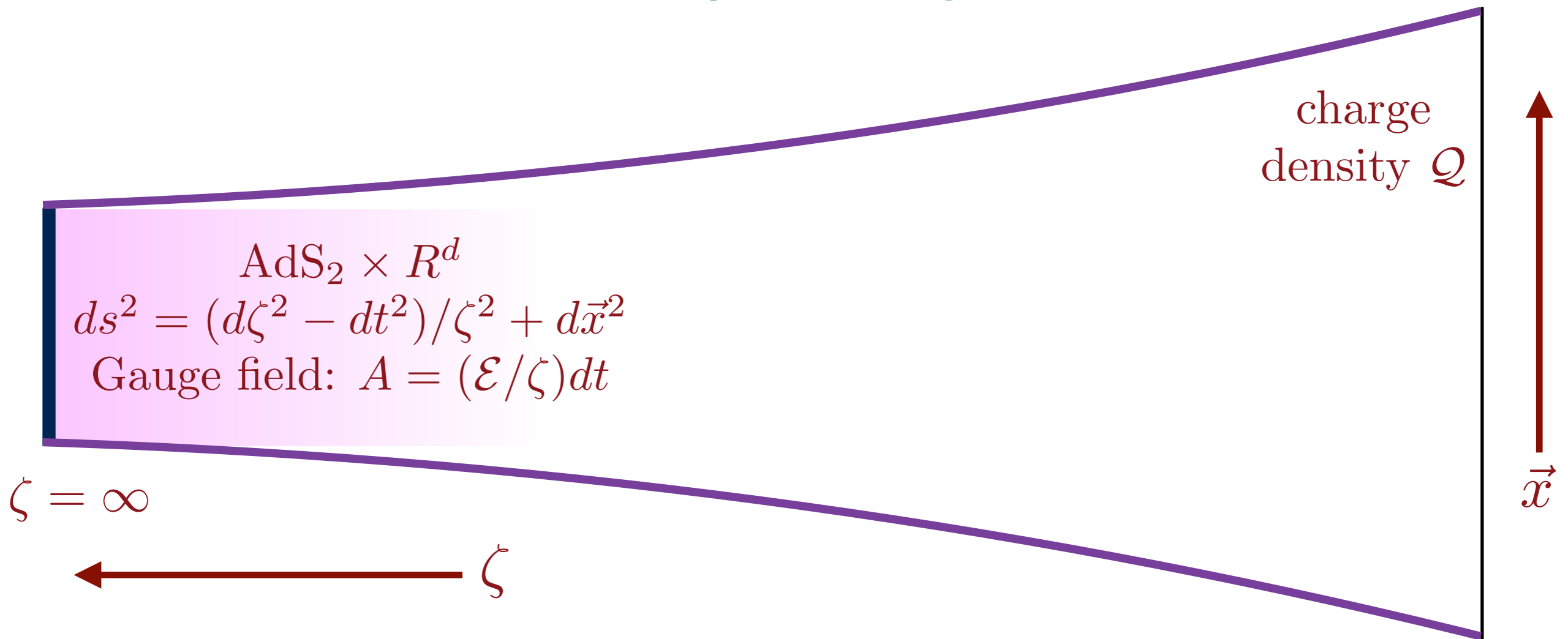
Realizes a strange metal: a state with an unbroken global U(1) symmetry with a continuously variable charge density,  $\mathcal{Q}$ , at  $T = 0$  which does not have any quasiparticle excitations.

# General Relativity of charged black branes



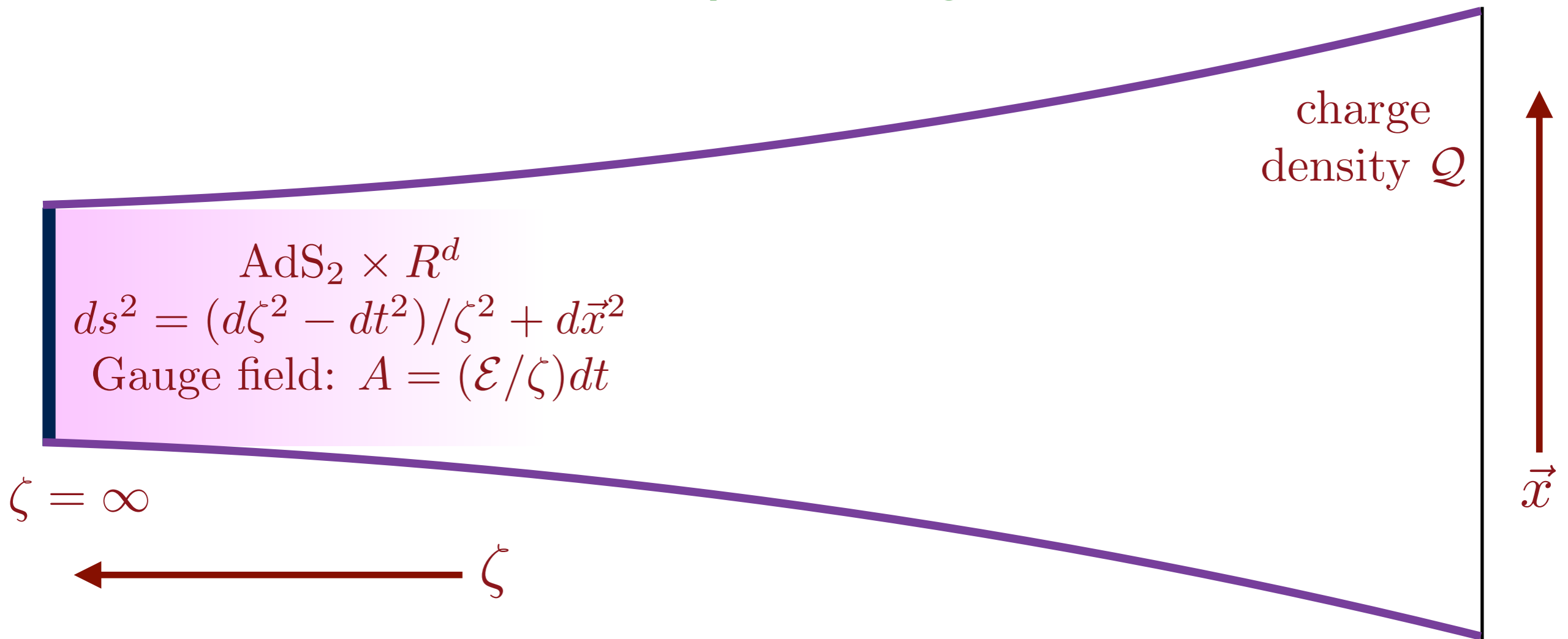
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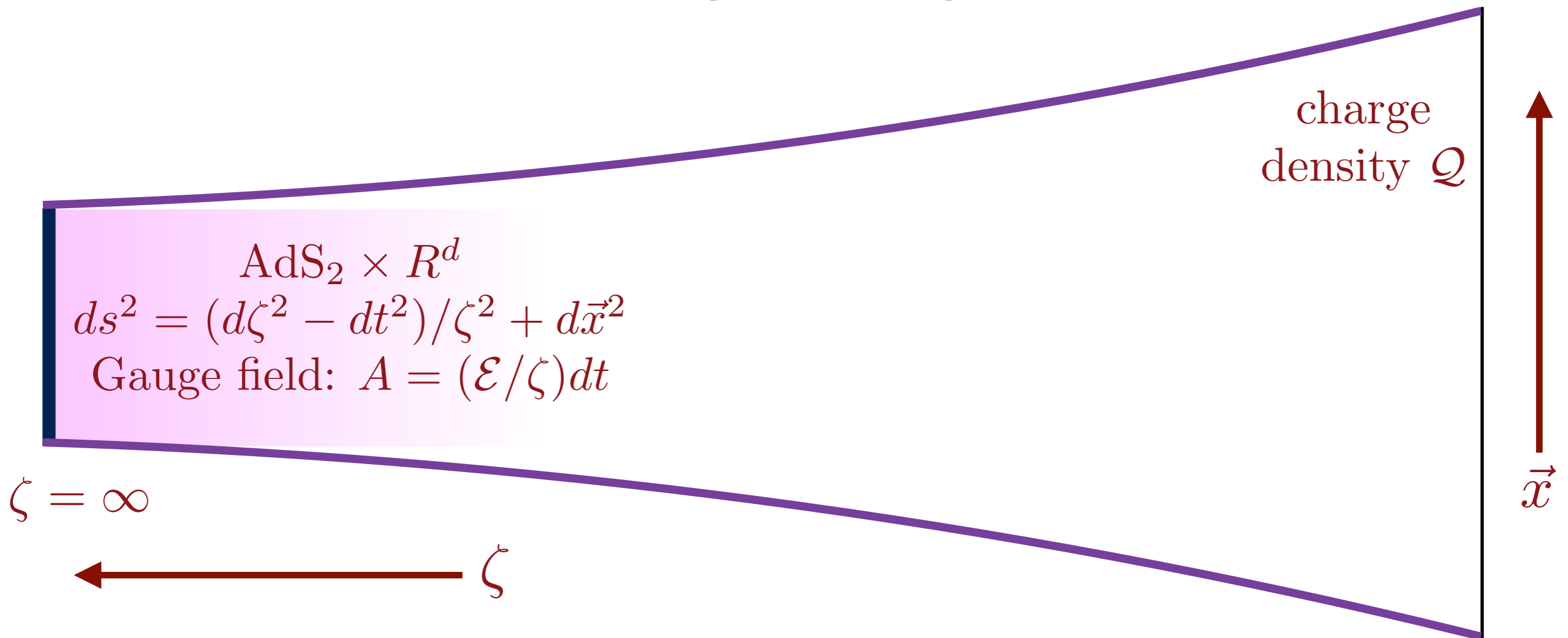
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- Both  $\mathcal{E}$  and  $\mathcal{S}_{\text{BH}}$  are determined by  $Q$ , and both vanish as  $Q \rightarrow 0$ .
- Near the boundary,  $A = \mu dt$ , where  $\mu$  is the chemical potential

# General Relativity of charged black branes

Conformal mapping to  $T > 0$

$$\zeta = \zeta_0$$

charge  
density  $Q$

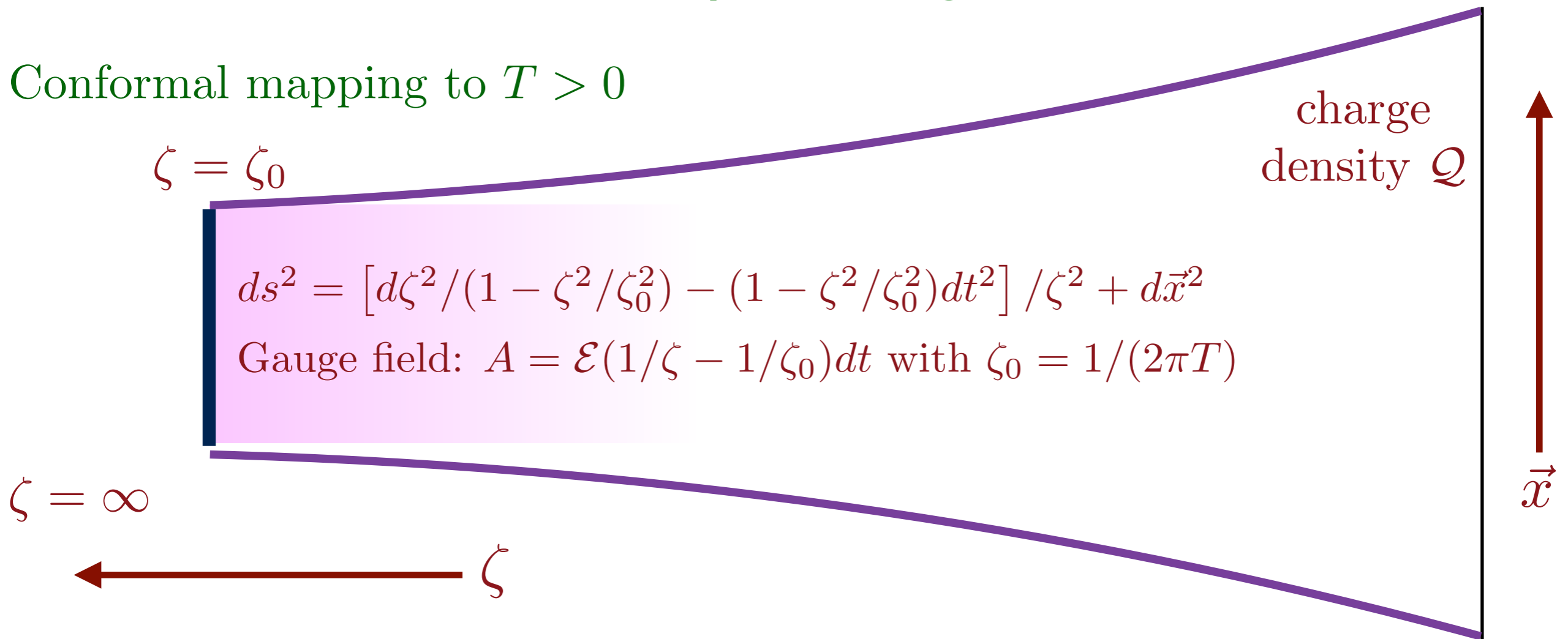
$$ds^2 = [d\zeta^2 / (1 - \zeta^2 / \zeta_0^2) - (1 - \zeta^2 / \zeta_0^2) dt^2] / \zeta^2 + d\vec{x}^2$$

$$\text{Gauge field: } A = \mathcal{E}(1/\zeta - 1/\zeta_0) dt \text{ with } \zeta_0 = 1/(2\pi T)$$

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- Using a thermodynamic Maxwell relation (also obeyed by gravity),

$$\left( \frac{\partial \mathcal{S}_{\text{BH}}}{\partial Q} \right)_T = - \left( \frac{\partial \mu}{\partial T} \right)_Q = 2\pi\mathcal{E}$$

A. Sen

hep-th/0506177

S. Sachdev

1506.05111

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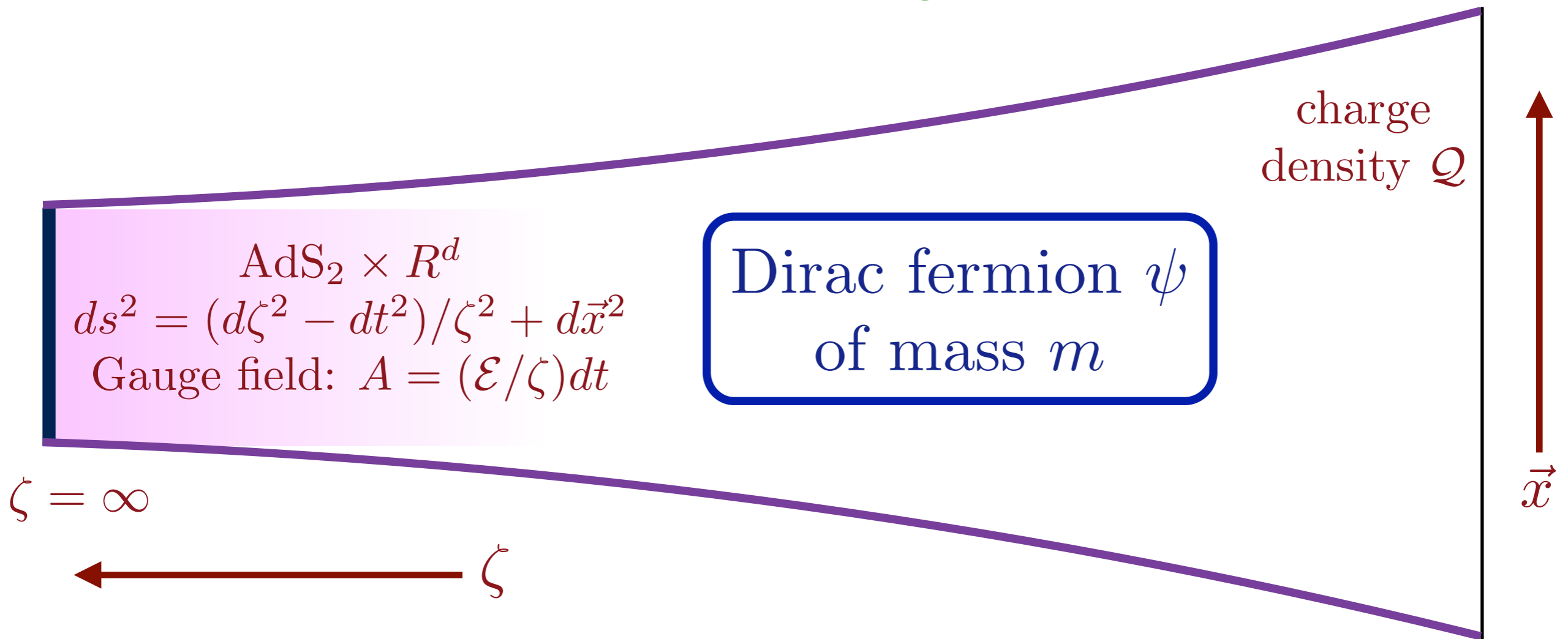
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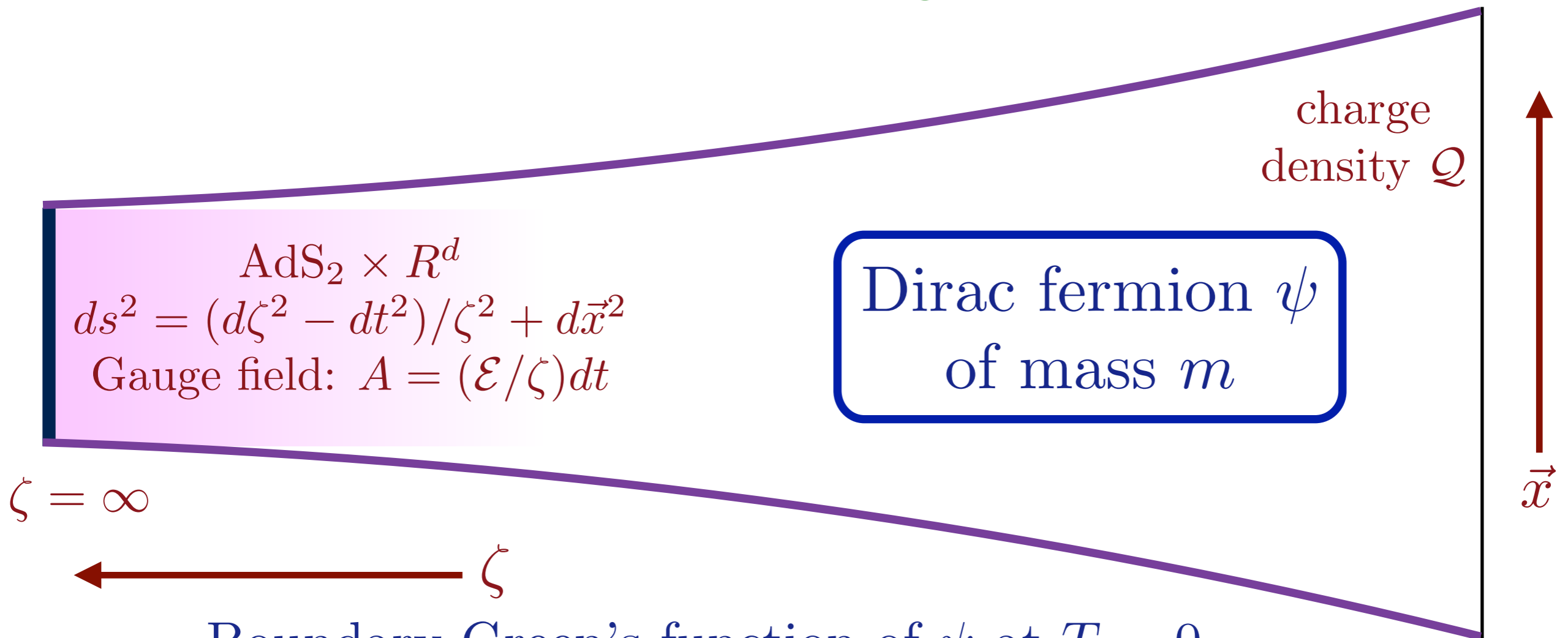
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- Also obeyed by the Wald entropy in higher derivative gravity.

# Quantum fields on charged black branes



# Quantum fields on charged black branes



Boundary Green's function of  $\psi$  at  $T = 0$

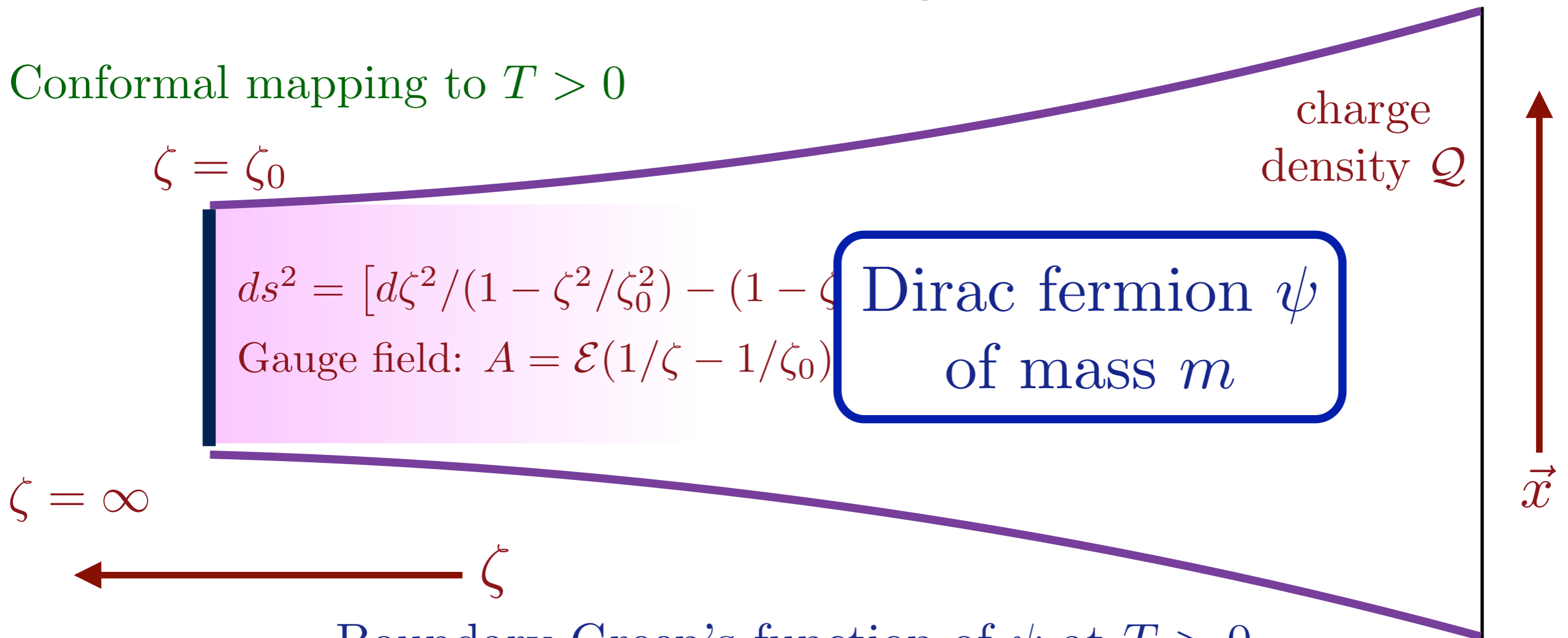
$$\text{Im}G(\omega) \sim \begin{cases} \omega^{-(1-2\Delta)}, & \omega > 0 \\ e^{-2\pi\mathcal{E}} |\omega|^{-(1-2\Delta)}, & \omega < 0. \end{cases}$$

where the fermion scaling dimension  $\Delta$  is a function of  $m$

$\mathcal{E}$  encodes the particle-hole asymmetry

# Quantum fields on charged black branes

Conformal mapping to  $T > 0$



Boundary Green's function of  $\psi$  at  $T > 0$

$$G^R(\omega) = \frac{-iC e^{-i\theta}}{(2\pi T)^{1-2\Delta}} \frac{\Gamma\left(\Delta - \frac{i\omega}{2\pi T} + i\mathcal{E}\right)}{\Gamma\left(1 - \Delta - \frac{i\omega}{2\pi T} + i\mathcal{E}\right)}$$

where  $e^{2\pi\mathcal{E}} = \frac{\sin(\pi\Delta + \theta)}{\sin(\pi\Delta - \theta)}$ .

What is a possible quantum theory on  
the boundary ?

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A critical strange metal state with infinite-range  
interactions obtained in

S. Sachdev and J. Ye, Phys. Rev. Lett. **70**, 3339 (1993)

S. Sachdev, Phys. Rev. Lett. **105**, 151602 (2010)



# Infinite-range strange metals

$$H = \frac{1}{(NM)^{1/2}} \sum_{i,j=1}^N \sum_{\alpha,\beta=1}^M J_{ij} c_{i\alpha}^\dagger c_{i\beta} c_{j\beta}^\dagger c_{j\alpha}$$

$$c_{i\alpha} c_{j\beta} + c_{j\beta} c_{i\alpha} = 0 \quad , \quad c_{i\alpha} c_{j\beta}^\dagger + c_{j\beta}^\dagger c_{i\alpha} = \delta_{ij} \delta_{\alpha\beta}$$

$$\frac{1}{M} \sum_{\alpha} c_{i\alpha}^\dagger c_{i\alpha} = Q$$

$J_{ij}$  are independent random variables with  $\overline{J_{ij}} = 0$  and  $\overline{J_{ij}^2} = J^2$   
 $N \rightarrow \infty$  at  $M = 2$  yields spin-glass ground state.

$N \rightarrow \infty$  and then  $M \rightarrow \infty$  yields critical strange metal

S. Sachdev and J. Ye, Phys. Rev. Lett. **70**, 3339 (1993)

**OR**

$$H = \frac{1}{(2N)^{3/2}} \sum_{i,j,k,l=1}^N J_{ij;kl} c_i^\dagger c_j^\dagger c_k c_l - \mu \sum_i c_i^\dagger c_i$$

$$c_i c_j + c_j c_i = 0 \quad , \quad c_i c_j^\dagger + c_j^\dagger c_i = \delta_{ij}$$

$$Q = \frac{1}{N} \sum_i c_i^\dagger c_i$$

$J_{ij;kl}$  are independent random variables with  $\overline{J_{ij;kl}} = 0$  and  $\overline{|J_{ij;kl}|^2} = J^2$   
 $N \rightarrow \infty$  yields same critical strange metal; simpler to study numerically

A. Kitaev, unpublished; S. Sachdev, arXiv:1506.05111

## Infinite-range strange metals

Feynman graph expansion in  $J_{ij..}$ , and graph-by-graph average, yields exact equations in the large  $N$  limit:

$$G(i\omega) = \frac{1}{i\omega + \mu - \Sigma(i\omega)} \quad , \quad \Sigma(\tau) = -J^2 G^2(\tau) G(-\tau)$$
$$G(\tau = 0^-) = \mathcal{Q}.$$

Low frequency analysis shows that the solutions must be gapless and obey

$$\Sigma(z) = \mu - \frac{1}{A} \sqrt{z} + \dots \quad , \quad G(z) = \frac{A}{\sqrt{z}}$$

for some complex  $A$ . Let us also define  $\tilde{\Sigma}(z) = \Sigma(z) - \mu$ .

# Infinite-range strange metals

At frequencies  $\ll J$ , the equations for  $G$  and  $\Sigma$  can be written as

$$\int d\tau_2 G(\tau_1, \tau_2) \tilde{\Sigma}(\tau_2, \tau_3) = -\delta(\tau_1 - \tau_3)$$
$$\tilde{\Sigma}(\tau_1, \tau_2) = -J^2 [G(\tau_1, \tau_2)]^2 G(\tau_2, \tau_1)$$

These equations are invariant under the reparametrization and gauge transformations

$$\tau = f(\sigma)$$

$$G(\tau_1, \tau_2) = [f'(\sigma_1) f'(\sigma_2)]^{-1/4} \frac{g(\sigma_1)}{g(\sigma_2)} G(\sigma_1, \sigma_2)$$

$$\tilde{\Sigma}(\tau_1, \tau_2) = [f'(\sigma_1) f'(\sigma_2)]^{-3/4} \frac{g(\sigma_1)}{g(\sigma_2)} \tilde{\Sigma}(\sigma_1, \sigma_2)$$

where  $f(\sigma)$  and  $g(\sigma)$  are arbitrary functions.

These equations and invariances have similarities to those of the large  $N$  limit of quantum spins at the spatial boundary of a CFT<sub>2</sub> (multi-channel Kondo problems)

A. Georges and O. Parcollet  
PRB 59, 5341 (1999)  
A. Kitaev, unpublished  
S. Sachdev, arXiv:1506.05111

O. Parcollet, A. Georges, G. Kotliar, and A. Sengupta  
PRB 58, 3794 (1998)

# Infinite-range strange metals

From these expressions we obtain the Green's function

$$G^R(\omega) = \frac{-iC e^{-i\theta}}{(2\pi T)^{1-2\Delta}} \frac{\Gamma\left(\Delta - \frac{i\omega}{2\pi T} + i\mathcal{E}\right)}{\Gamma\left(1 - \Delta - \frac{i\omega}{2\pi T} + i\mathcal{E}\right)}$$

where  $\Delta = 1/4$  and  $e^{2\pi\mathcal{E}} = \frac{\sin(\pi\Delta + \theta)}{\sin(\pi\Delta - \theta)}$

and  $Q = \frac{1}{4}(3 - \tanh(2\pi\mathcal{E})) - \frac{1}{\pi} \tan^{-1}(e^{2\pi\mathcal{E}})$ .

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A. Georges, O. Parcollet, and S. Sachdev Phys. Rev. B **63**, 134406 (2001)

# Infinite-range strange metals

The entropy per site,  $\mathcal{S}$ , has a non-zero limit as  $T \rightarrow 0$ , and is similar to universal boundary entropy of the Kondo problem.

N. Andrei and C. Destri, PRL **52**, 364 (1984).

A. M. Tsvelick, J. Phys. C **18**, 159 (1985).

I. Affleck and A. W. W. Ludwig, PRL **67**, 161 (1991).

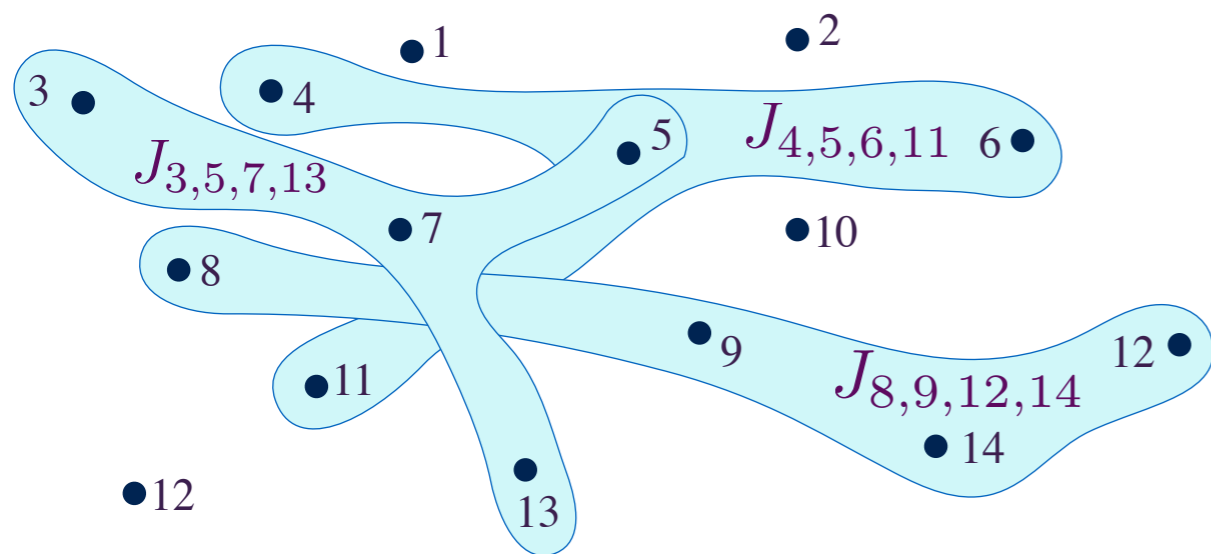
This entropy obeys

$$\left(\frac{\partial \mathcal{S}}{\partial \mathcal{Q}}\right)_T = - \left(\frac{\partial \mu}{\partial T}\right)_{\mathcal{Q}} = 2\pi\mathcal{E}$$

O. Parcollet, A. Georges, G. Kotliar, and A. Sengupta Phys. Rev. B **58**, 3794 (1998)

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$J_{ij;kl}$  independent  
random numbers

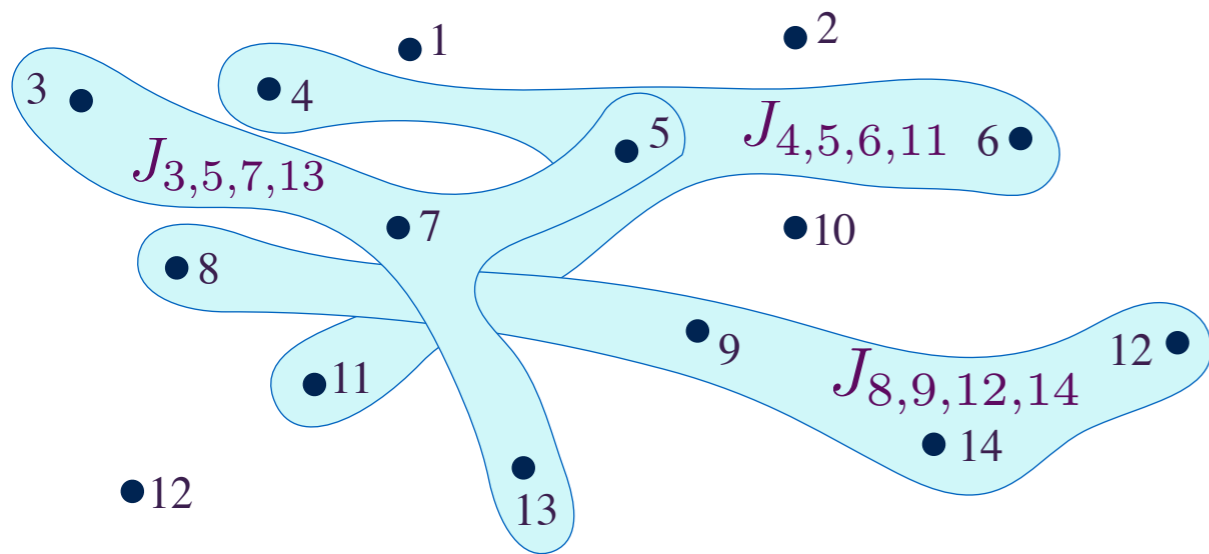
## An infinite-range model of a strange metal

S. Sachdev and J. Ye, Phys. Rev. Lett. **70**, 3339 (1993)

A. Kitaev, unpublished

S. Sachdev, arXiv:1506.05111

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Local fermion density of states

$$\rho(\omega) \sim \begin{cases} \omega^{-1/2}, & \omega > 0 \\ e^{-2\pi\mathcal{E}} |\omega|^{-1/2}, & \omega < 0. \end{cases}$$

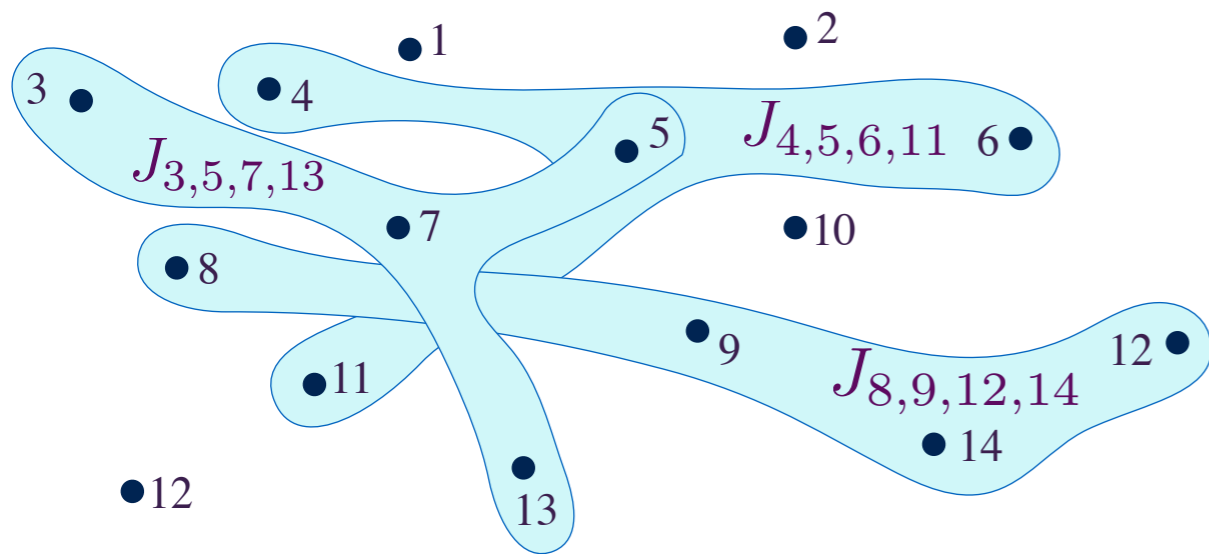
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Known 'equation of state'  
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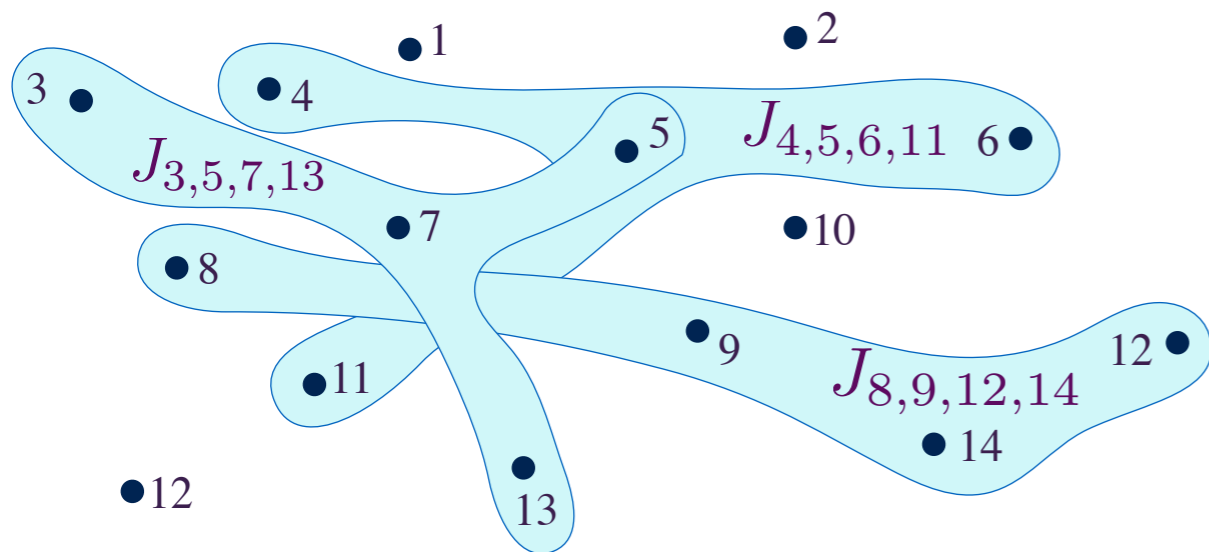
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Microscopic zero temperature  
entropy density,  $\mathcal{S}$ , obeys

$$\frac{\partial \mathcal{S}}{\partial Q} = 2\pi\mathcal{E}$$

$$c_i c_j + c_j c_i = 0$$

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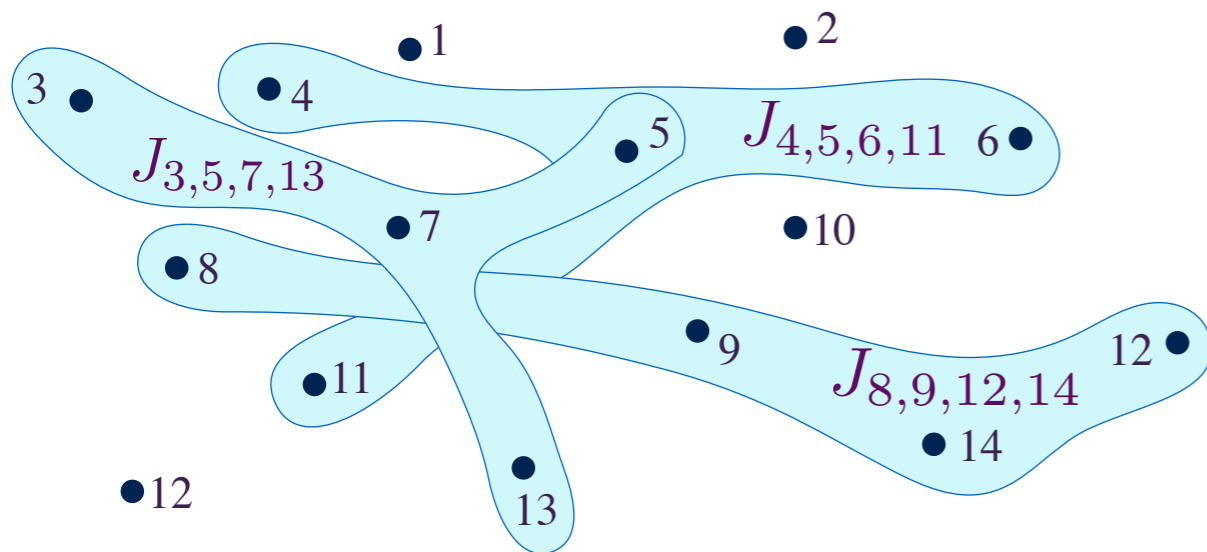
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Einstein-Maxwell theory  
+ cosmological constant

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$$Q = \frac{1}{N} \sum_i \langle c_i^\dagger c_i \rangle.$$

Local fermion density of states

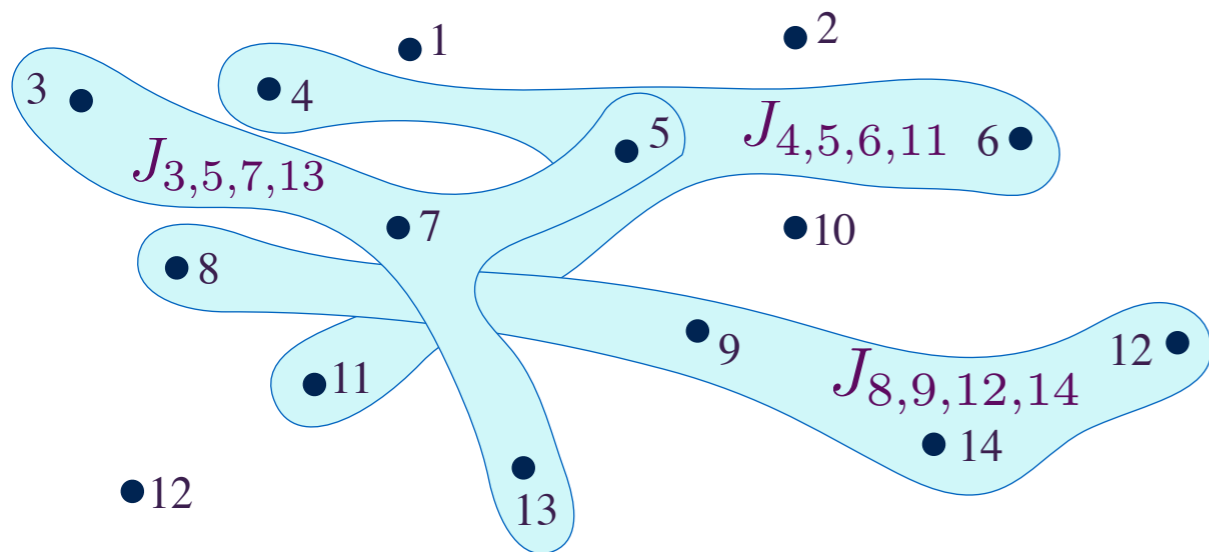
$$\rho(\omega) \sim \begin{cases} \omega^{-1/2}, & \omega > 0 \\ e^{-2\pi\mathcal{E}} |\omega|^{-1/2}, & \omega < 0. \end{cases}$$

Known 'equation of state'  
determines  $\mathcal{E}$  as a function of  $Q$

Microscopic zero temperature  
entropy density,  $\mathcal{S}$ , obeys

$$\frac{\partial \mathcal{S}}{\partial Q} = 2\pi\mathcal{E}$$

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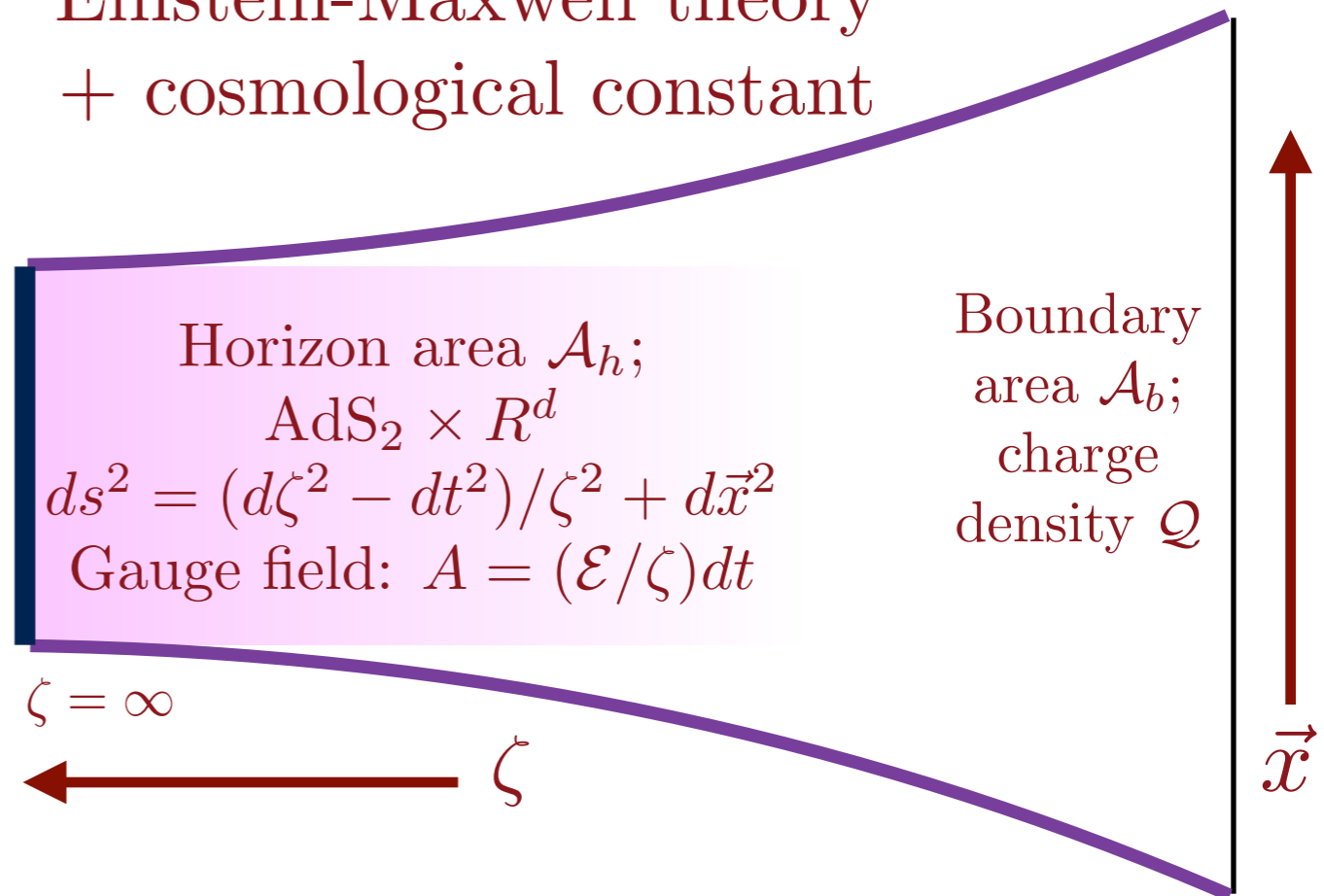
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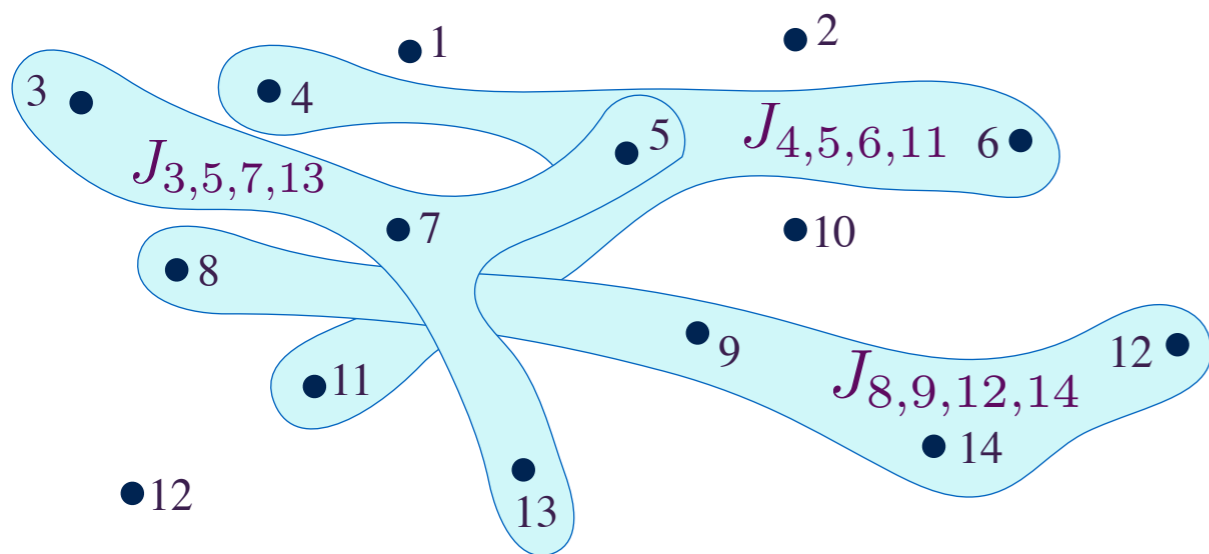
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Einstein-Maxwell theory  
+ cosmological constant



A. Chamblin, R. Emparan, C.V. Johnson, and R.C. Myers  
Phys. Rev. D **60**, 064018 (1999)

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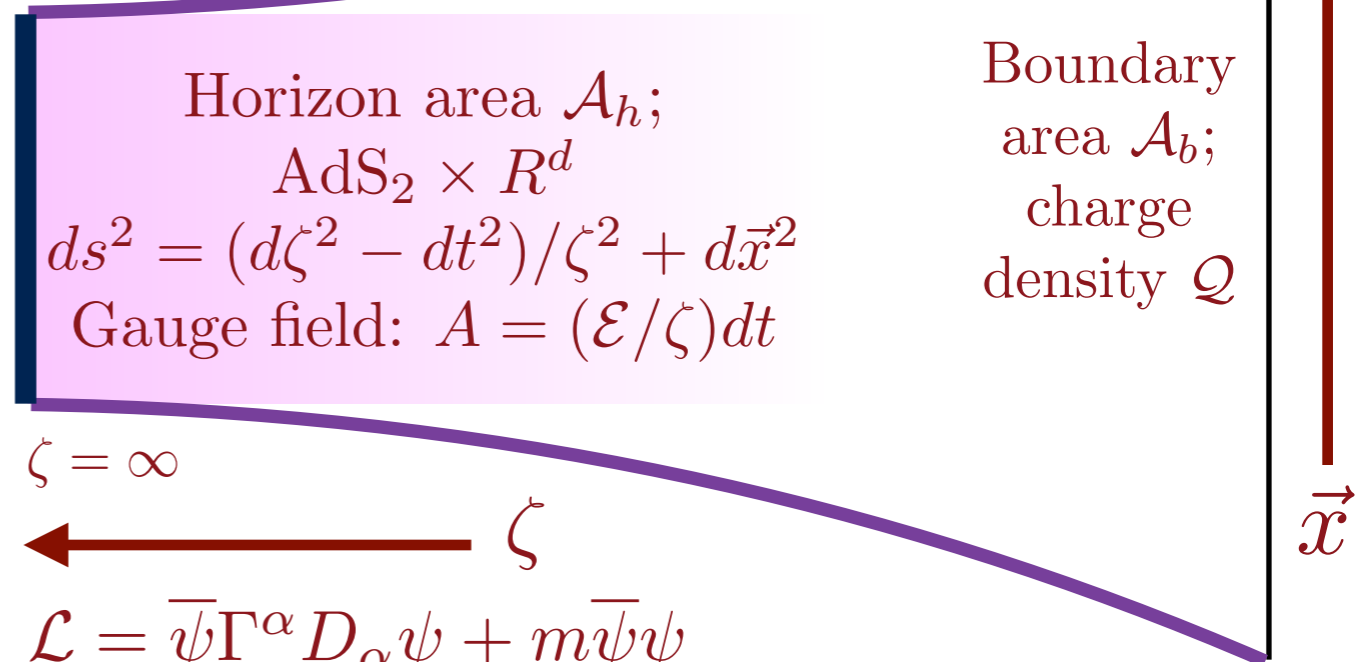
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$$\zeta = \infty$$

$$\leftarrow \zeta$$

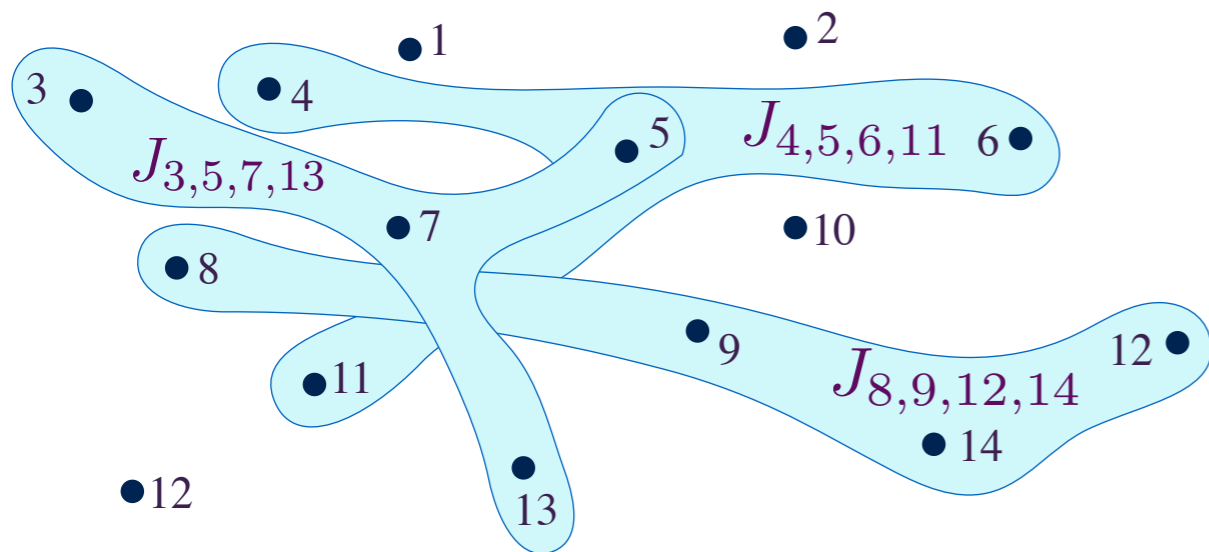
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T. Faulkner, Hong Liu, J. McGreevy, and D. Vegh  
Phys. Rev. D 83, 125002 (2011)

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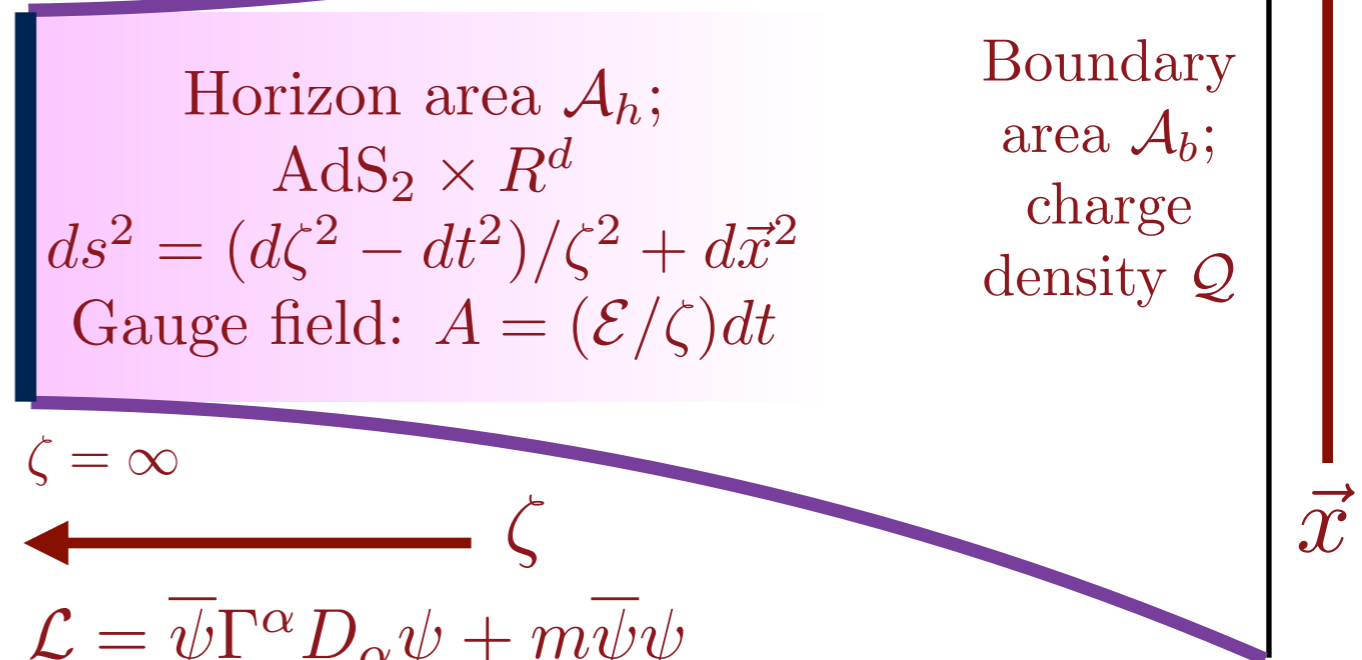
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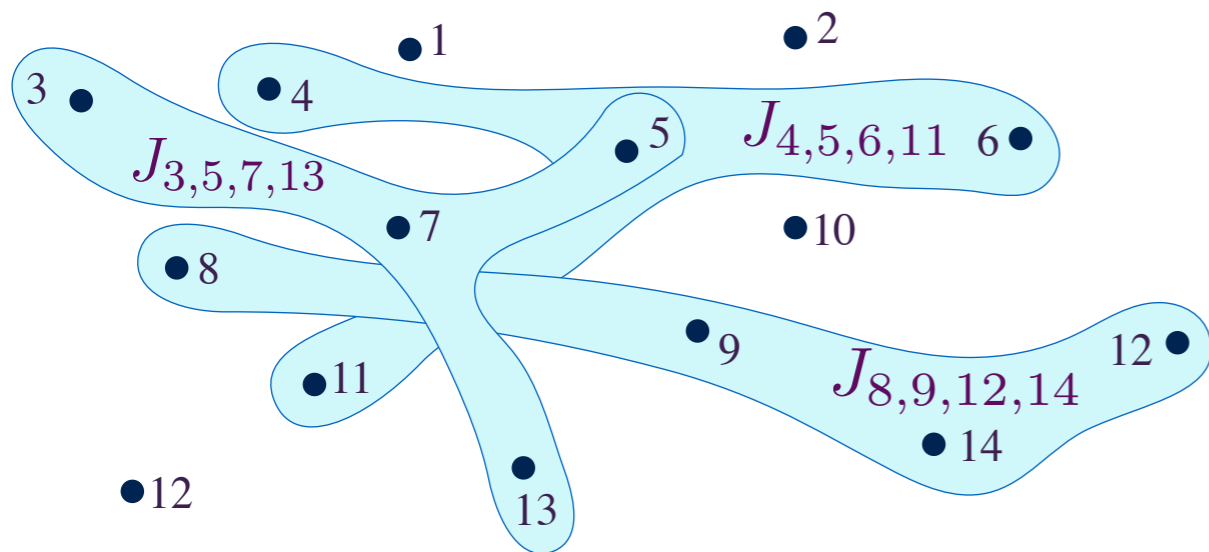
'Equation of state' relating  $\mathcal{E}$  and  $Q$  depends upon the geometry of spacetime far from the  $\text{AdS}_2$

Eliminate  $r_0$  between

$$Q = \frac{r_0^{d-1} \sqrt{2d [(d-1)R^2 + (d+1)r_0^2]}}{\kappa^2 g_F}$$

$$\mathcal{E} = \frac{g_F r_0 \sqrt{2d [(d-1)R^2 + (d+1)r_0^2]}}{2 [(d-1)^2 R^2 + d(d+1)r_0^2]}$$

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Local fermion density of states

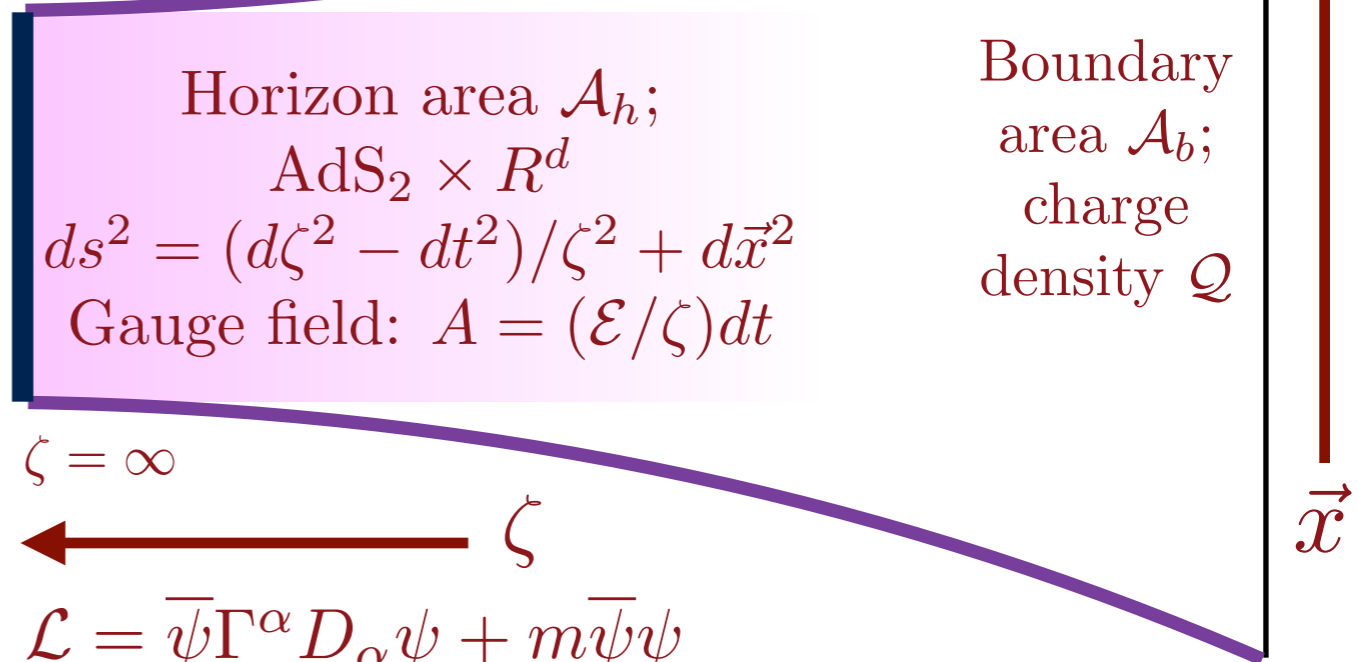
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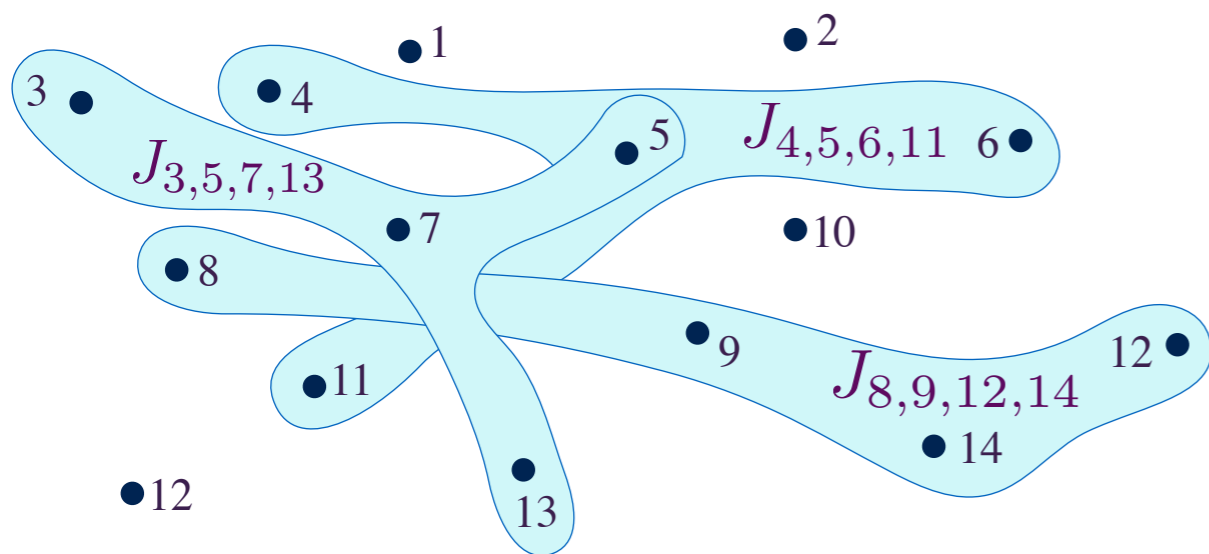
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‘Equation of state’ relating  $\mathcal{E}$  and  $Q$  depends upon the geometry of spacetime far from the  $\text{AdS}_2$

Black hole thermodynamics (classical general relativity) yields

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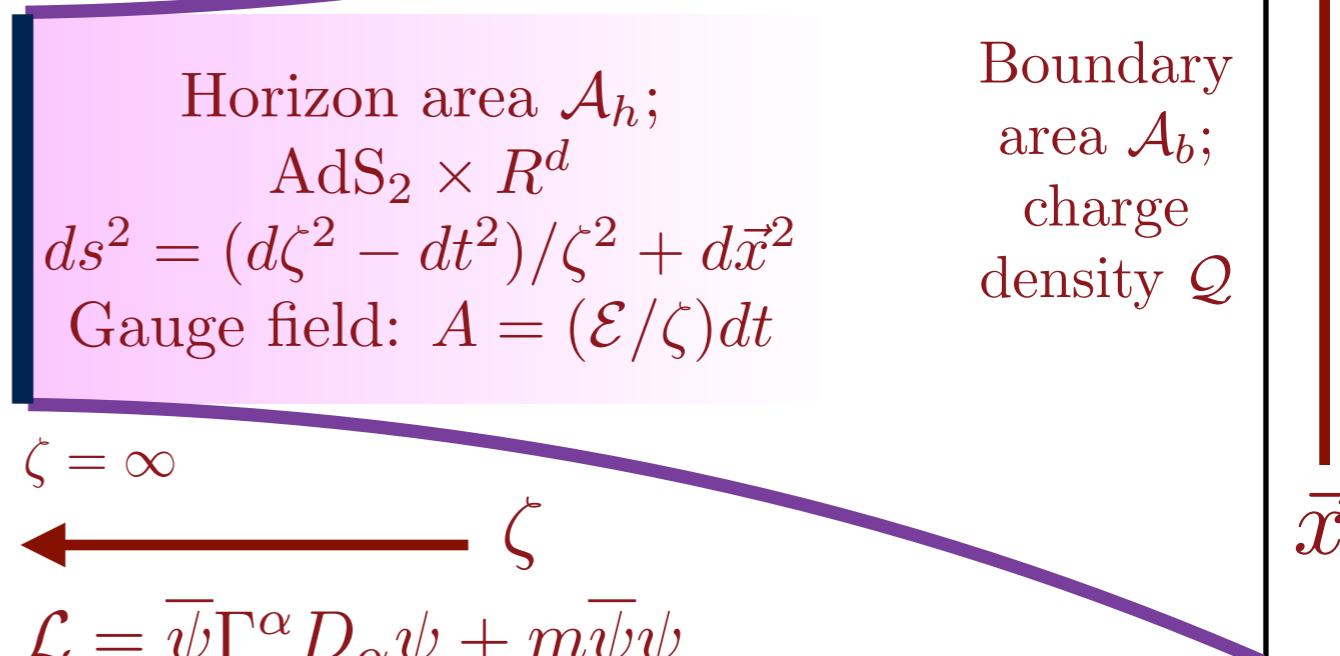
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Evidence for AdS<sub>2</sub> gravity dual of  $H$

Einstein-Maxwell theory + cosmological constant



Horizon area  $\mathcal{A}_h$ ;

AdS<sub>2</sub> × R<sup>d</sup>

$$ds^2 = (d\zeta^2 - dt^2)/\zeta^2 + d\vec{x}^2$$

$$\text{Gauge field: } A = (\mathcal{E}/\zeta)dt$$

Boundary area  $\mathcal{A}_b$ ;

charge density  $\mathcal{Q}$

$$\zeta = \infty$$

$$\zeta$$

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