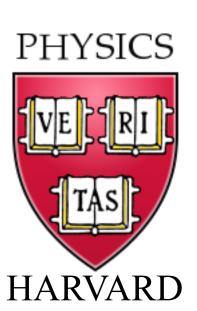
# Universal theory of complex SYK models and extremal charged black holes

Subir Sachdev

Jerusalem Winter School, December 31, 2018

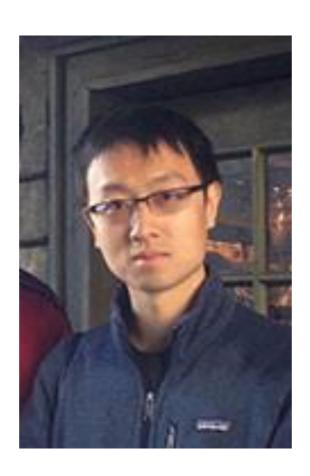








Wenbo Fu



Yingfei Gu



Grigory Tarnopolsky







## I. Quantum matter with quasiparticles: random matrix model

2. Quantum matter without quasiparticles: the complex SYK model

3. Einstein-Maxwell theory of charged black holes in AdS space

4. Fluctuations, and the Schwarzian

#### Quantum matter with quasiparticles: random matrix model

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$$H = \frac{1}{(N)^{1/2}} \sum_{i,j=1}^{N} t_{ij} c_i^{\dagger} c_j - \mu \sum_i c_i^{\dagger} c_i$$

$$c_i c_j + c_j c_i = 0 \quad , \quad c_i c_j^{\dagger} + c_j^{\dagger} c_i = \delta_{ij}$$

$$\frac{1}{N} \sum_i c_i^{\dagger} c_i = \mathcal{Q}$$

 $t_{ij}$  are independent random variables with  $\overline{t_{ij}} = 0$  and  $|\overline{t_{ij}}|^2 = t^2$ 

### Fermions occupying the eigenstates of a $N \times N$ random matrix

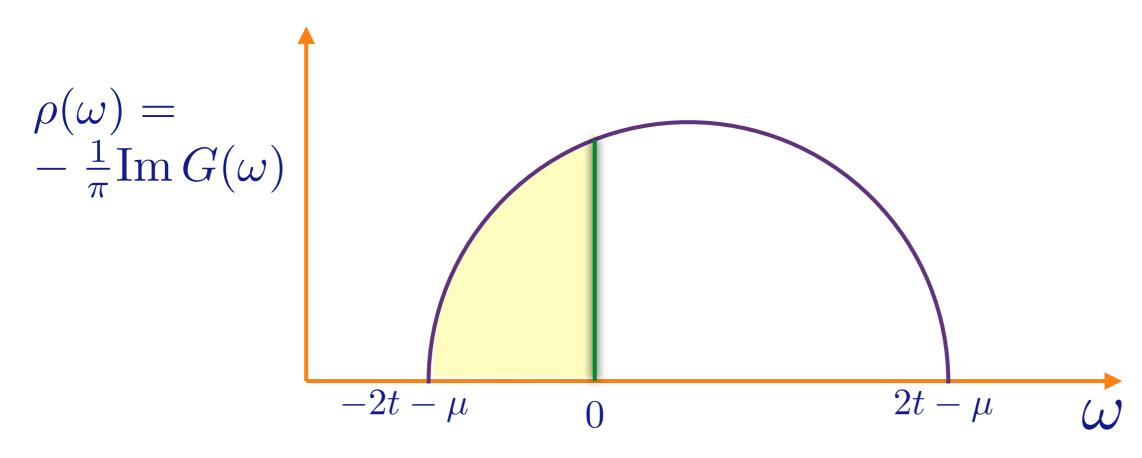
Feynman graph expansion in  $t_{ij...}$ , and graph-by-graph average, yields exact equations in the large N limit:

$$G(\tau) \equiv -T_{\tau} \left\langle c_{i}(\tau) c_{i}^{\dagger}(0) \right\rangle$$

$$G(i\omega) = \frac{1}{i\omega + \mu - \Sigma(i\omega)} , \quad \Sigma(\tau) = t^{2}G(\tau)$$

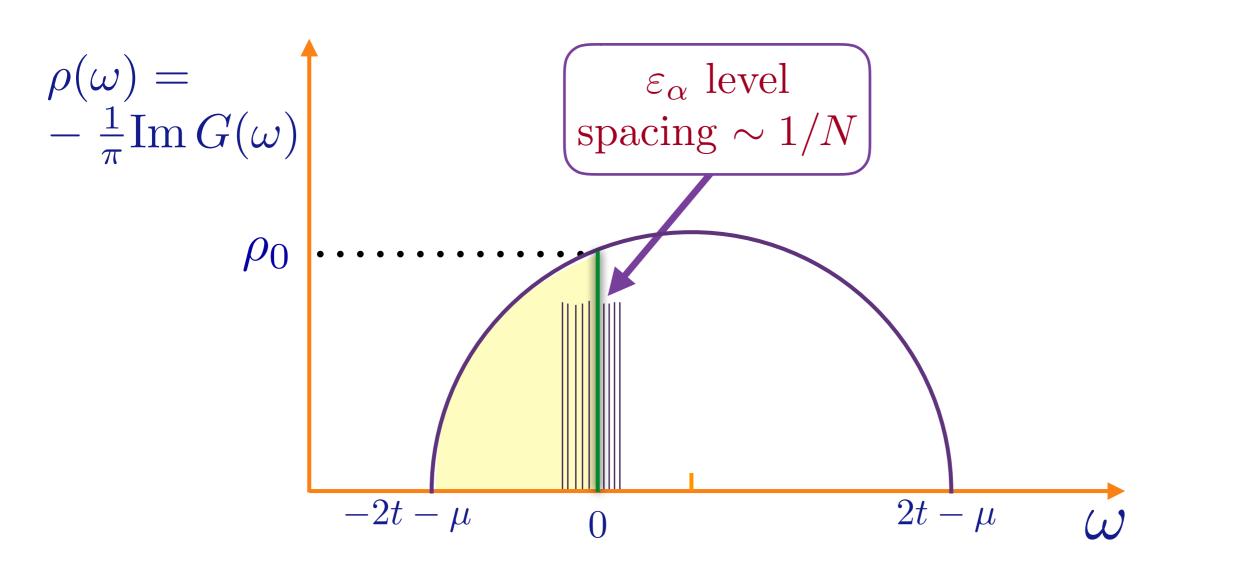
$$G(\tau = 0^{-}) = Q.$$

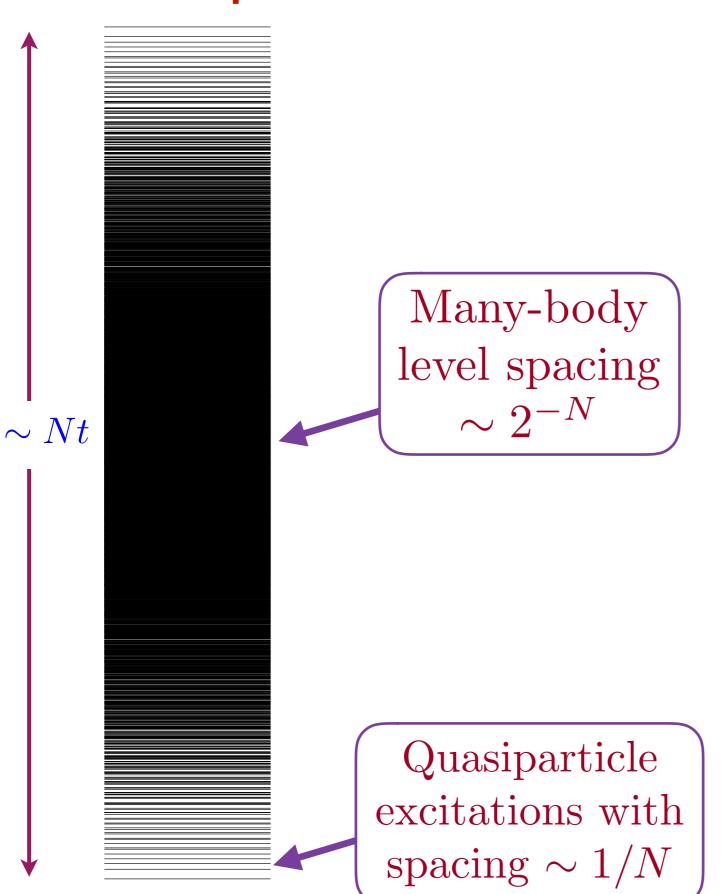
 $G(\omega)$  can be determined by solving a quadratic equation.



Let  $\varepsilon_{\alpha}$  be the eigenvalues of the matrix  $t_{ij}/\sqrt{N}$ . The fermions will occupy the lowest NQ eigenvalues, upto the Fermi energy  $E_F$ . The single-particle density of states is

$$\rho(\omega) = (1/N) \sum_{\alpha} \delta(\omega - \varepsilon_{\alpha}), \text{ and } \rho_0 \equiv \rho(\omega = 0).$$





There are  $2^N$  many body levels with energy

$$E = \sum_{\alpha=1}^{N} n_{\alpha} \varepsilon_{\alpha},$$

where  $n_{\alpha} = 0, 1$ . Shown are all values of E for a single cluster of size N = 12. The  $\varepsilon_{\alpha}$  have a level spacing  $\sim 1/N$ .

The grand potential  $\Omega(T)$  at low T is (from the Sommerfeld expansion)

$$\Omega(T) - E_0 = N\left(-\frac{\pi^2}{6}\rho_0 T^2 + \mathcal{O}(T^4)\right) + \dots$$

where  $\rho_0 \equiv \rho(0)$  is the *single* particle density of states at the Fermi level. We can also define the *many* body density of states, D(E), via

$$Z = e^{-\Omega(T)/T} = \int_{-\infty}^{\infty} dE D(E) e^{-E/T}$$

The inversion from  $\Omega(T)$  to D(E) has to performed with care (it need not commute with the 1/N expansion), and we obtain

$$D(E) \sim \exp\left(\pi\sqrt{\frac{2N\rho_0(E - E_0)}{3}}\right) \quad , \quad E > E_0 \; , \; \frac{1}{N} \ll \rho_0(E - E_0) \ll N$$

and D(E) = 0 for  $E < E_0$ . This is related to the asymptotic growth of the partitions of an integer,  $p(n) \sim \exp(\pi \sqrt{2n/3})$ . Near the lower bound, there are large sample-to-sample fluctuations due to variations in the lowest quasiparticle energies.

Now add weak interactions

$$H = \frac{1}{(N)^{1/2}} \sum_{i,j=1}^{N} t_{ij} c_i^{\dagger} c_j - \mu \sum_{i} c_i^{\dagger} c_i + \frac{1}{(2N)^{3/2}} \sum_{i,j,k,\ell=1}^{N} U_{ij;k\ell} c_i^{\dagger} c_j^{\dagger} c_k c_\ell$$

 $U_{ij;k\ell}$  are independent random variables with  $\overline{U_{ij;k\ell}} = 0$  and  $|U_{ij;k\ell}|^2 = U^2$ . We compute the lifetime of a quasiparticle,  $\tau_{\alpha}$ , in an exact eigenstate  $\psi_{\alpha}(i)$  of the free particle Hamitonian with energy  $\varepsilon_{\alpha}$ . By Fermi's Golden rule, for  $\varepsilon_{\alpha}$  at the Fermi energy

$$\frac{1}{\tau_{\alpha}} = \pi U^{2} \rho_{0}^{2} \int d\varepsilon_{\beta} d\varepsilon_{\gamma} d\varepsilon_{\delta} f(\varepsilon_{\beta}) (1 - f(\varepsilon_{\gamma})) (1 - f(\varepsilon_{\delta})) \delta(\varepsilon_{\alpha} + \varepsilon_{\beta} - \varepsilon_{\gamma} - \varepsilon_{\delta})$$

$$= \frac{\pi^{3} U^{2} \rho_{0}^{2}}{4} T^{2}$$

where  $\rho_0$  is the density of states at the Fermi energy, and  $f(\epsilon) = 1/(e^{\epsilon/T} + 1)$  is the Fermi function.

Fermi liquid state: Two-body interactions lead to a scattering time of quasiparticle excitations from in (random) single-particle eigenstates which diverges as  $\sim T^{-2}$  at the Fermi level.

## I. Quantum matter with quasiparticles: random matrix model

2. Quantum matter without quasiparticles: the complex SYK model

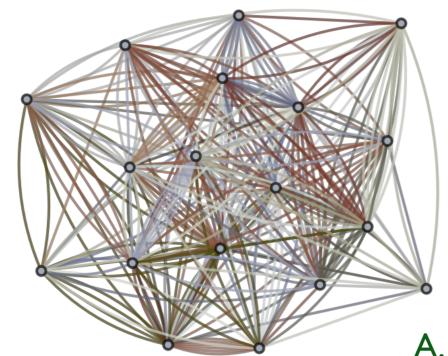
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(See also: the "2-Body Random Ensemble" in nuclear physics; did not obtain the large N limit; T.A. Brody, J. Flores, J.B. French, P.A. Mello, A. Pandey, and S.S.M. Wong, Rev. Mod. Phys. **53**, 385 (1981))

$$H = \frac{1}{(2N)^{3/2}} \sum_{i,j,k,\ell=1}^{N} U_{ij;k\ell} c_i^{\dagger} c_j^{\dagger} c_k c_{\ell} - \mu \sum_i c_i^{\dagger} c_i$$
$$c_i c_j + c_j c_i = 0 \quad , \quad c_i c_j^{\dagger} + c_j^{\dagger} c_i = \delta_{ij}$$
$$\mathcal{Q} = \frac{1}{N} \sum_i c_i^{\dagger} c_i$$

 $U_{ij;k\ell}$  are independent random variables with  $\overline{U_{ij;k\ell}} = 0$  and  $\overline{|U_{ij;k\ell}|^2} = U^2$  $N \to \infty$  yields critical strange metal.

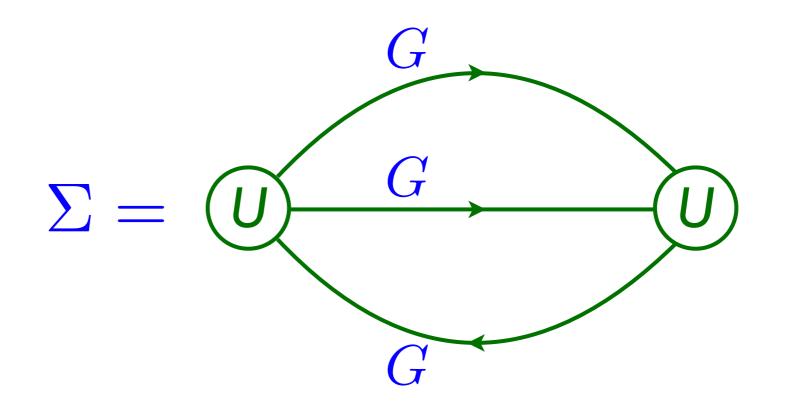


S. Sachdev and J. Ye, PRL **70**, 3339 (1993)

A. Kitaev, unpublished; S. Sachdev, PRX 5, 041025 (2015)

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$$G(i\omega) = \frac{1}{i\omega + \mu - \Sigma(i\omega)} , \quad \Sigma(\tau) = -U^2 G^2(\tau) G(-\tau)$$
$$G(\tau = 0^-) = \mathcal{Q}.$$



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- At general charge Q, there is a spectral symmetry determined by a parameter  $\mathcal{E}$ :

$$G(\tau) \sim \begin{cases} -\tau^{-2\Delta} & \tau > 0 \\ e^{-2\pi\mathcal{E}}(-\tau)^{-2\Delta} & \tau < 0 \end{cases}, \quad T = 0$$

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• There is a universal 'Luttinger relation' between  $-\infty < \mathcal{E} < \infty$  and the total charge  $0 < \mathcal{Q} < 1$ 

$$e^{2\pi\mathcal{E}} = rac{\sin(\pi\Delta + \theta)}{\sin(\pi\Delta - \theta)}$$
 R. In A. C. K. Jenson  $\mathcal{Q} = rac{1}{2} - rac{\theta}{\pi} + \left(\Delta - rac{1}{2}
ight) rac{\sin(2\theta)}{\sin(2\pi\Delta)}$ 

A. Georges, O. Parcollet, and S. Sachdev, PRB **63**, 134406 (2001) R. Davison, Wenbo Fu, A. Georges, Yingfei Gu, K. Jensen, S. Sachdev, PRB **95**, 155131 (2017)

Solution of these equations and corresponding evaluation of the free energy yields the following universal results (i.e. all results are quantitatively unchanged by adding additional higher q fermion terms):

• At T > 0, we obtain a solution with a conformal structure

$$G(\tau) = -A \frac{e^{-2\pi\mathcal{E}T\tau}}{\sqrt{1 + e^{-4\pi\mathcal{E}}}} \left(\frac{T}{\sin(\pi T\tau)}\right)^{1/2} , \quad 0 < \tau < 1/T,$$

where the 'particle-hole asymmetry' is determined by  $\mathcal{E}$ 

A. Georges and O. Parcollet PRB **59**, 5341 (1999) S. Sachdev, PRX **5**, 041025 (2015)

We now examine the behavior of the chemical potential,  $\mu$ , as  $T \to 0$  at fixed  $\mathcal{Q}$ . For this we relate the long-time 'conformal' Greens function, (valid for  $\tau \gg 1/U$ ) to its short-time behavior. In particular at  $|\omega_n| \gg U$  we have

$$G(i\omega_n) = \frac{1}{i\omega_n} - \frac{\mu}{(i\omega_n)^2} + \dots$$

which implies for the spectral density of the Green's function,  $\rho(\Omega)$ 

$$\mu = -\int_{-\infty}^{\infty} \frac{d\Omega}{\pi} \,\Omega\rho(\Omega),$$

which makes it evident that  $\mu$  depends only upon the particle-hole asymmetric part of the spectral density. Next, we can relate the  $\Omega$  integrals to the derivative of the imaginary time correlator

$$\mu = -\partial_{\tau} G(\tau = 0^{+}) - \partial_{\tau} G(\tau = (1/T)^{-}).$$

We pull out an explicitly particle-hole asymmetric part of  $G(\tau)$  by defining

$$G(\tau) \equiv e^{-2\pi \mathcal{E}T\tau} G_c(\tau)$$
 ,  $0 < \sigma < \frac{1}{T}$ .

where  $G_c$  will be given by a particle-hole symmetric conformal form at low T and low  $\omega$ . Then we obtain

$$\mu = 2\pi \mathcal{E}T \left[ G(\tau = 0^+) + G(\tau = (1/T)^-) \right] - \partial_{\tau}g(\tau = 0^+)$$

$$+ \text{ terms dependent on } G_c$$

$$= -2\pi \mathcal{E}T + \text{ terms dependent on } G_c$$

It can be shown that all the terms dependent upon  $G_c$  have a T dependence that is weaker than linear in T provided Q is held fixed. Hence we have

$$\mu = \mu_0 - 2\pi \mathcal{E}T + \text{terms vanishing as } T^p \text{ with } p > 1$$

with  $\mu_0$  a non-universal constant. From this relation we obtain

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$$\left(\frac{\partial \mu}{\partial T}\right)_{\mathcal{Q}} = -2\pi\mathcal{E} \quad , \quad T \to 0,$$

Using a Maxwell relation we then have

$$\frac{1}{N} \left( \frac{\partial S}{\partial \mathcal{Q}} \right)_T = 2\pi \mathcal{E} \neq 0 \text{ as } T \to 0.$$

Solution of these equations and corresponding evaluation of the free energy yields the following universal results (i.e. all results are quantitatively unchanged by adding additional higher q fermion terms):

• There is a non-vanishing entropy in the zero temperature limit

$$S(T \to 0) = Ns_0 + \dots$$

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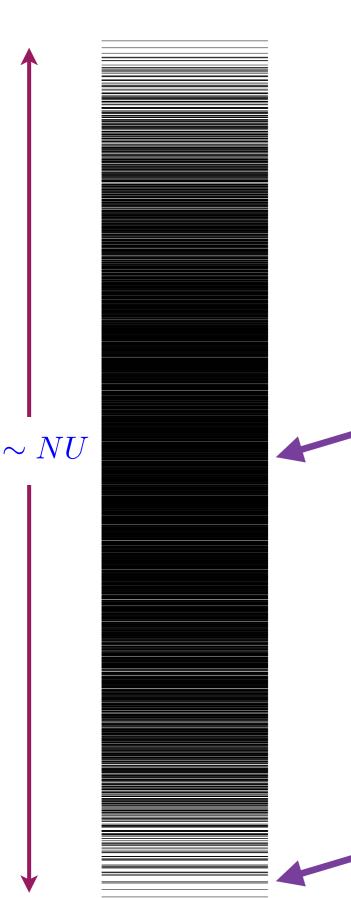
$$S(T \to 0) = Ns_0 + \dots$$

• The saddle point equations imply the relation

$$\frac{ds_0}{d\mathcal{Q}} = 2\pi\mathcal{E}$$

Integrating this relation from  $s_0 = 0$ , Q = 0, allows us to compute  $s_0$  as a function of Q.

A. Georges, O. Parcollet, and S. Sachdev, PRB 63, 134406 (2001)



Many-body level spacing  $\sim$  $2^{-N} = e^{-N \ln 2}$ 

Non-quasiparticle

excitations with

spacing  $\sim e^{-Ns_0}$ 

There are  $2^N$  many body levels with energy E. Shown are all values of E for a single cluster of size N=12. The  $T\to 0$  state has an entropy  $S_{GPS}=Ns_0$ , where  $s_0<\ln 2$  is determined by integrating

$$\frac{ds_0}{d\mathcal{Q}} = 2\pi\mathcal{E} \,.$$

At 
$$Q = 1/2$$
,

$$s_0 = \frac{G}{\pi} + \frac{\ln(2)}{4} = 0.464848\dots$$

where G is Catalan's constant.

GPS: A. Georges, O. Parcollet, and S. Sachdev, PRB **63**, 134406 (2001)

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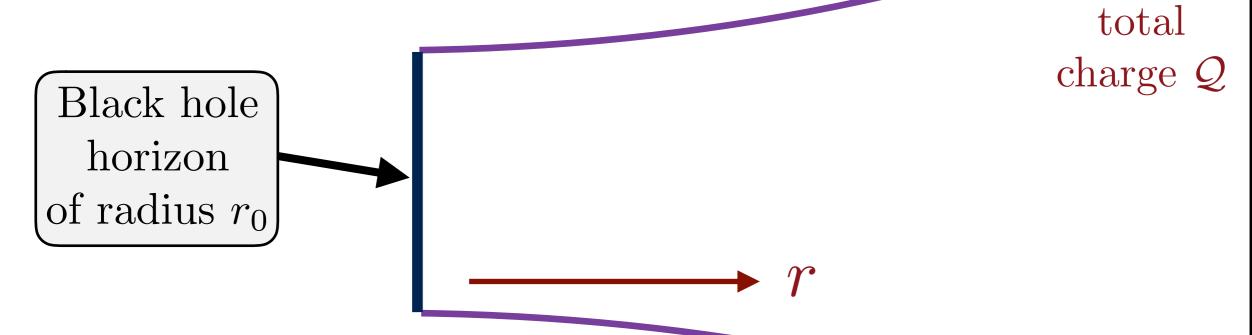
2. Quantum matter without quasiparticles: the complex SYK model

S. Sachdev, Phys. Rev. Lett. **I 05**, 151602 (2010) S. Sachdev, PRX **5**, 041025 (2015)

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$$S_{EM} = \frac{1}{2\kappa^2} \int d^{d+2}x \sqrt{-g} \left( \mathcal{R} + \frac{6}{L^2} - \frac{L^2}{g_F^2} F^2 \right)$$



Solutions of  $S_{EM}$  have metric and gauge field (F = dA)

$$ds^{2} = -V(r)dt^{2} + r^{2}d\Omega_{d}^{2} + \frac{dr^{2}}{V(r)} , \quad A = \mu \left(1 - \frac{r_{0}^{d-1}}{r^{d-1}}\right)dt$$
$$V(r) = 1 + \frac{r^{2}}{L^{2}} + \frac{\Theta^{2}}{r^{2d-2}} - \frac{M}{r^{d-1}}.$$

where  $d\Omega_d^2$  is the metric of the d-sphere. All parameters of the solution are determined in terms of the chemical potential  $\mu$ , and the Hawking temperature of horizon, T.

A. Chamblin, R. Emparan, C.V. Johnson, and R.C. Myers, PRD **60**, 064018 (1999)

In the  $T \to 0$  limit, at fixed  $\mu$ , we obtain a charged black hole solution with radius  $r_0(T \to 0, \mu) = R_h$ . All properties of this black hole can be expressed in terms of  $R_h$ 

• The total charge in the black hole is

$$Q = \frac{R_h^{d-1} \sqrt{2d \left[ (d+1)R_h^2 + (d-1)L^2 \right]}}{\kappa^2 g_F}$$

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• The total charge in the black hole is

$$Q = \frac{R_h^{d-1} \sqrt{2d \left[ (d+1)R_h^2 + (d-1)L^2 \right]}}{\kappa^2 g_F}$$

• The Bekenstein-Hawking entropy remains finite as  $T \to 0$  ( $s_d$  is the area of the d-dimensional surface of a unit sphere)

$$S(T \to 0) = s_0 + \dots$$
 ,  $s_0 = \frac{2\pi s_d}{\kappa^2} R_h^d$ 

In the  $T \to 0$  limit, at fixed  $\mu$ , we obtain a charged black hole solution with radius  $r_0(T \to 0, \mu) = R_h$ . All properties of this black hole can be expressed in terms of  $R_h$ 

• In the near-horizon region, we change co-ordinates from r to  $\zeta$  so that

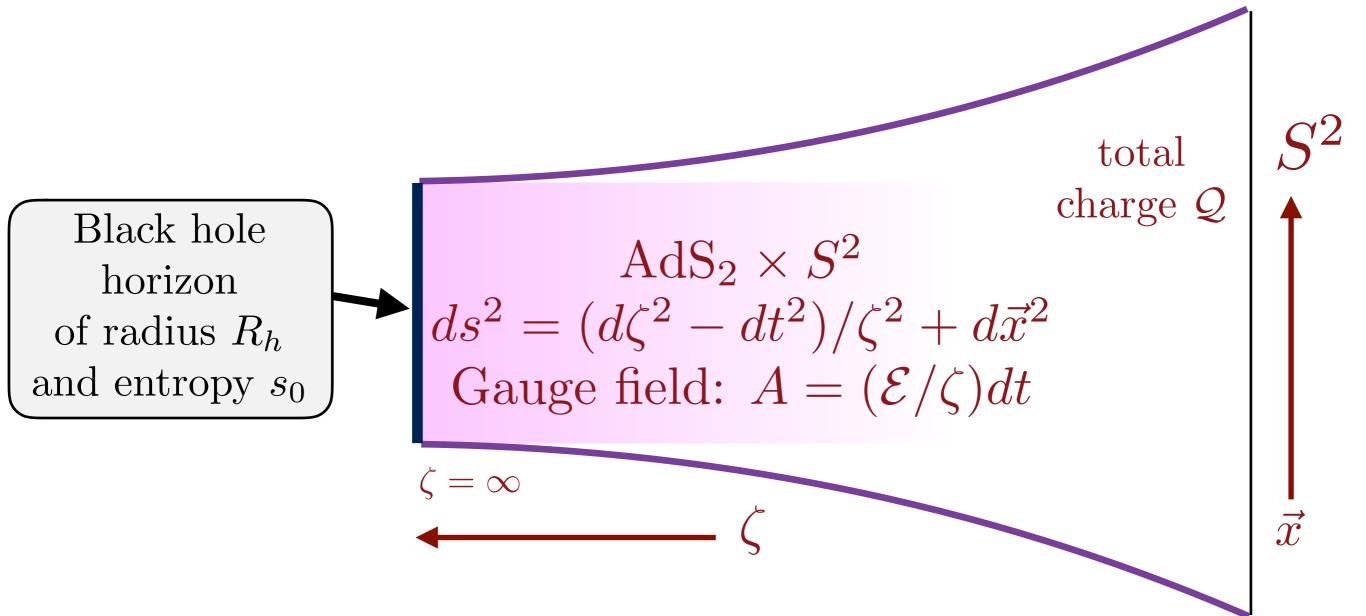
$$r - R_h = \frac{R_2^2}{\zeta}$$
 ,  $R_2 = \frac{LR_h}{\sqrt{d(d+1)R_h^2 + (d-1)^2L^2}}$ .

Then the near-horizon metric becomes  $AdS_2 \times S_d$ , with

$$ds^{2} = R_{2}^{2} \left[ \frac{-dt^{2} + d\zeta^{2}}{\zeta^{2}} \right] + R_{h}^{2} d\Omega_{d}^{2} , \quad A = \frac{\mathcal{E}}{\zeta} dt.$$

where the dimensionless electric field  $\mathcal{E}$  is

$$\mathcal{E} = \frac{g_F R_h \sqrt{2d \left[ (d+1) R_h^2 + (d-1) L^2 \right]}}{2 \left[ d(d+1) R_h^2 + (d-1)^2 L^2 \right]}.$$



• The entropy  $s_0$ , the charge  $\mathcal{Q}$ , and the dimensionless electric field  $\mathcal{E}$  obey

$$\frac{ds_0}{d\mathcal{Q}} = 2\pi\mathcal{E}$$

#### Dimensional reduction to 2D gravity

We neglect all dependence of the metric on angular co-ordinates, introduce a scalar field  $\phi(r,\tau)$ , and write the 4D metric as

$$ds^2 = \frac{1}{\sqrt{\phi}} h_{ab} dx^a dx^b + \phi d\Omega_2^2$$

where  $a, b = x, \tau$  and  $h_{ab}$  is a 2D metric. Then the 4D action  $S_{EM}$  reduces to an action for 2D gravity

$$S_{JT} = \frac{2\pi}{\kappa^2} \int d^2x \sqrt{-h} \left( \phi \mathcal{R}_{2D} + V(\phi) - \frac{Z(\phi)L^2}{g_F^2} F^2 \right)$$

with

$$V(\phi) = \frac{2}{\sqrt{\phi}} + \frac{6\sqrt{\phi}}{L^2} \quad , \quad Z(\phi) = \phi^{3/2}$$

This describes Jackiw-Teitelbaum 2D gravity, along with a 2D electromagnetic field.

#### Probe fermion in the AdS2 near horizon

• A probe fermion has a near-horizon Green's function with a conformal structure

$$G(\tau) = -A \frac{e^{-2\pi\mathcal{E}T\tau}}{\sqrt{1 + e^{-4\pi\mathcal{E}}}} \left(\frac{T}{\sin(\pi T\tau)}\right)^{1/2} , \quad 0 < \tau < 1/T,$$

where the 'particle-hole asymmetry' is determined by  $\mathcal{E}$ . This is identical to the complex SYK model.

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#### The SYK model

$$G(i\omega) = \frac{1}{i\omega + \mu - \Sigma(i\omega)} , \quad \Sigma(\tau) = -U^2 G^2(\tau) G(-\tau)$$
$$\Sigma(z) = \mu - \frac{1}{A} \sqrt{z} + \dots , \quad G(z) = \frac{A}{\sqrt{z}}$$

#### The SYK model

$$G(i\omega) = \frac{1}{(\omega + \chi - \Sigma(i\omega))}, \quad \Sigma(\tau) = -U^2 G^2(\tau) G(-\tau)$$

$$\Sigma(z) = \chi - \frac{1}{A} \sqrt{z} + \dots, \quad G(z) = \frac{A}{\sqrt{z}}$$

At frequencies  $\ll U$ , the  $i\omega + \mu$  can be dropped, and without it equations are invariant under the reparametrization and gauge transformations. The singular part of the self-energy and the Green's function obey

$$\int_0^\beta d\tau_2 \, \Sigma_{\text{sing}}(\tau_1, \tau_2) G(\tau_2, \tau_3) = -\delta(\tau_1 - \tau_3)$$
$$\Sigma_{\text{sing}}(\tau_1, \tau_2) = -U^2 G^2(\tau_1, \tau_2) G(\tau_2, \tau_1)$$

## The SYK model

$$\int_0^\beta d\tau_2 \, \Sigma_{(\tau_1, \tau_2)} G(\tau_2, \tau_3) = -\delta(\tau_1 - \tau_3)$$
$$\Sigma(\tau_1, \tau_2) = -U^2 G^2(\tau_1, \tau_2) G(\tau_2, \tau_1)$$

These equations are invariant under

$$\tau = f(\sigma)$$

$$G(\tau_1, \tau_2) = \left[ f'(\sigma_1) f'(\sigma_2) \right]^{-1/4} \frac{g(\sigma_1)}{g(\sigma_2)} \widetilde{G}(\sigma_1, \sigma_2)$$

$$\Sigma(\tau_1, \tau_2) = \left[ f'(\sigma_1) f'(\sigma_2) \right]^{-3/4} \frac{g(\sigma_1)}{g(\sigma_2)} \widetilde{\Sigma}(\sigma_1, \sigma_2)$$

where  $f(\sigma)$  and  $g(\sigma)$  are arbitrary functions. By using  $f(\sigma) = \tan(\pi T \sigma)/(\pi T)$  we can now obtain the T > 0 solution from the T = 0 solution.

## The SYK model

Let us write the large N saddle point solutions of S as

$$G_s(\tau_1 - \tau_2) \sim (\tau_1 - \tau_2)^{-1/2}$$
  
 $\Sigma_s(\tau_1 - \tau_2) \sim (\tau_1 - \tau_2)^{-3/2}.$ 

The saddle point will be invariant under a reperamaterization  $f(\tau)$  when choosing  $G(\tau_1, \tau_2) = G_s(\tau_1 - \tau_2)$  leads to a transformed  $\widetilde{G}(\sigma_1, \sigma_2) = G_s(\sigma_1 - \sigma_2)$  (and similarly for  $\Sigma$ ). It turns out this is true only for the SL(2, R) transformations under which

$$f(\tau) = \frac{a\tau + b}{c\tau + d}$$
 ,  $ad - bc = 1$ .

So the (approximate) reparametrization symmetry is spontaneously broken down to SL(2, R) by the saddle point.

• The saddle-point

$$G(\tau_1 - \tau_2) = -A \frac{e^{-2\pi \mathcal{E}T(\tau_1 - \tau_2)}}{\sqrt{1 + e^{-4\pi \mathcal{E}}}} \left( \frac{T}{\sin(\pi T(\tau_1 - \tau_2))} \right)^{2\Delta}$$

is invariant only under PSL(2, R) transformations which map the thermal circle onto itself, and an associated gauge transformation

$$\frac{\tan(\pi T f(\tau))}{\pi T} = \frac{a \frac{\tan(\pi T \tau)}{\pi T} + b}{c \frac{\tan(\pi T \tau)}{\pi T} + d} , \quad ad - bc = 1,$$

$$-i\phi(\tau) = -i\phi_0 + 2\pi \mathcal{E}T(\tau - f(\tau))$$

A. Kitaev, 2015

### Infinite-range (SYK) model without quasiparticles

After introducing replicas  $a = 1 \dots n$ , and integrating out the disorder, the partition function can be written as

$$Z = \int \mathcal{D}c_{ia}(\tau) \exp\left[-\sum_{ia} \int_{0}^{\beta} d\tau \, c_{ia}^{\dagger} \left(\frac{\partial}{\partial \tau} - \mu\right) c_{ia}\right]$$
$$-\frac{U^{2}}{4N^{3}} \sum_{ab} \int_{0}^{\beta} d\tau d\tau' \left|\sum_{i} c_{ia}^{\dagger}(\tau) c_{ib}(\tau')\right|^{4}.$$

For simplicity, we neglect the replica indices, and introduce the identity

$$1 = \int \mathcal{D}\Sigma(\tau_1, \tau_2) \exp\left[-N \int_0^\beta d\tau_1 d\tau_2 \Sigma(\tau_1, \tau_2) \left(G(\tau_2, \tau_1) + \frac{1}{N} \sum_i c_i(\tau_2) c_i^{\dagger}(\tau_1)\right)\right].$$

### Infinite-range (SYK) model without quasiparticles

Then the partition function can be written as a path integral with an action S analogous to a Luttinger-Ward functional

$$Z = \int \mathcal{D}G(\tau_{1}, \tau_{2}) \mathcal{D}\Sigma(\tau_{1}, \tau_{2}) \exp(-NS)$$

$$S = \ln \det \left[\delta(\tau_{1} - \tau_{2})(\partial_{\tau_{1}} + \mu) - \Sigma(\tau_{1}, \tau_{2})\right]$$

$$+ \int d\tau_{1} d\tau_{2} \Sigma(\tau_{1}, \tau_{2}) \left[G(\tau_{2}, \tau_{1}) + (U^{2}/2)G^{2}(\tau_{2}, \tau_{1})G^{2}(\tau_{1}, \tau_{2})\right]$$

At frequencies  $\ll U$ , the time derivative in the determinant is less important, and without it the path integral is invariant under the reparametrization and gauge transformations

A. Georges and O. Parcollet

 $\tau = f(\sigma)$  A. Kitaev, 2015 S. Sachdev, PRX **5,** 041025 (2015)  $G(\tau_1,\tau_2) = \left[f'(\sigma_1)f'(\sigma_2)\right]^{-1/4} \frac{g(\sigma_1)}{g(\sigma_2)} \, G(\sigma_1,\sigma_2)$ 

$$\Sigma(\tau_1, \tau_2) = [f'(\sigma_1)f'(\sigma_2)]^{-3/4} \frac{g(\sigma_1)}{g(\sigma_2)} \Sigma(\sigma_1, \sigma_2)$$

where  $f(\sigma)$  and  $g(\sigma)$  are arbitrary functions.

### The SYK model

#### Reparametrization and phase zero modes

We can write the path integral for the SYK model as

$$\mathcal{Z} = \int \mathcal{D}G(\tau_1, \tau_1) \mathcal{D}\Sigma(\tau_1, \tau_2) e^{-NS[G, \Sigma]}$$

for a known action  $S[G, \Sigma]$ . We find the saddle point,  $G_s$ ,  $\Sigma_s$ , and only focus on the "Nambu-Goldstone" modes associated with breaking reparameterization and U(1) gauge symmetries by writing

$$G(\tau_1, \tau_2) = [f'(\tau_1)f'(\tau_2)]^{1/4}G_s(f(\tau_1) - f(\tau_2))e^{i\phi(\tau_1) - i\phi(\tau_2)}$$

(and similarly for  $\Sigma$ ). Then the path integral is approximated by

$$\mathcal{Z} = \int \mathcal{D}f(\tau)\mathcal{D}\phi(\tau)e^{-NS_{\text{eff}}[f,\phi]}.$$

J. Maldacena and D. Stanford, arXiv:1604.07818;

R. Davison, Wenbo Fu, A. Georges, Yingfei Gu, K. Jensen, S. Sachdev, arXiv. 1612.00849; S. Sachdev, PRX 5, 041025 (2015); J. Maldacena, D. Stanford, and Zhenbin Yang, arXiv:1606.01857;

K. Jensen, arXiv: 1605.06098; J. Engelsoy, T.G. Mertens, and H. Verlinde, arXiv: 1606.03438

Symmetry arguments, and explicit computations, show that the effective action is

$$S_{\text{eff}}[f,\phi] = \frac{NK}{2} \int_0^{1/T} d\tau (\partial_\tau \phi + i(2\pi \mathcal{E}T)\partial_\tau f)^2 - \frac{N\gamma}{4\pi^2} \int_0^{1/T} d\tau \left\{ \tan(\pi T f(\tau)), \tau \right\},$$

where  $f(\tau)$  is a monotonic map from [0, 1/T] to [0, 1/T], the couplings  $K, \gamma$ , and  $\mathcal{E}$  can be related to thermodynamic derivatives and we have used the Schwarzian:

$$\{g,\tau\} \equiv \frac{g'''}{g'} - \frac{3}{2} \left(\frac{g''}{g'}\right)^2.$$

Specifically, an argument constraining the effective at T=0 is

$$S_{\text{eff}}\left[f(\tau) = \frac{a\tau + b}{c\tau + d}, \phi(\tau) = 0\right] = 0,$$

and this is origin of the Schwarzian.

J. Maldacena and D. Stanford, arXiv:1604.07818;

R. Davison, Wenbo Fu, A. Georges, Yingfei Gu, K. Jensen, S. Sachdev, PRB **95**, 155131 (2017); A. Gaikwad, L.K. Joshi, G. Mandal, and S.R. Wadia, arXiv:1802.07746

An exact path integral over the effective action leads to the following physical consequences

• The ground state energy with fermion number NQ + p (p integer) varies as

$$E_p = E_0 + \frac{p^2}{2NK}$$

This identifies K with the compressibility  $K = d\mathcal{Q}/d\mu$  at T = 0.

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• The low temperature corrections to the entropy at fixed  $\mathcal{Q}$  are

$$S(T \to 0, \mathcal{Q}) = N \left[ s_0 + \gamma T + \ldots \right] + 2 \ln(U/T) \ldots$$

This defines  $\gamma$  as the co-efficient of the linear-in-T specific heat (at fixed  $\mathcal{Q}$ )

D. Stanford and E. Witten, arXiv:1703.04612; A. Kitaev and S. Josephine Suh, arXiv:1711.08467

An exact path integral over the effective action leads to the following physical consequences

• The many-body density of states, D(E), is related to the grand potential,  $\Omega(T)$  by

$$Z = e^{-\Omega(T)/T} = \int_{-\infty}^{\infty} dE D(E) e^{-E/T}$$

We obtain

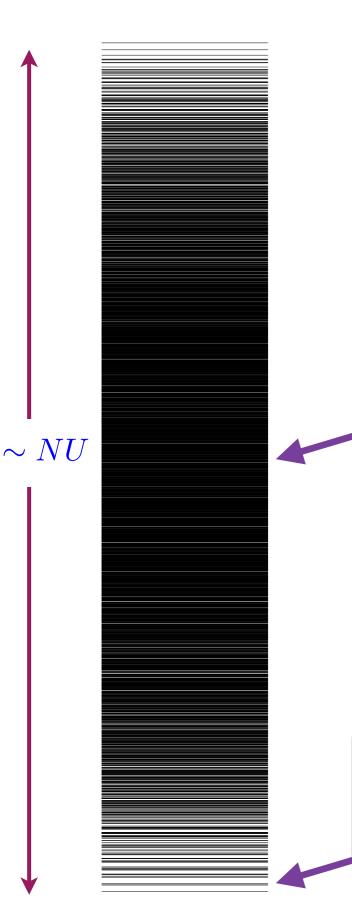
$$D(E) = \sum_{p=-\infty}^{\infty} e^{2\pi p\mathcal{E}} d(E - E_p)$$

where NQ + p is the integer fermion number,

$$d(E) \sim \exp(Ns_0) \sinh\left(\sqrt{2N\gamma E}\right)$$
 ,  $E > 0$  ,  $e^{-cN} \ll \gamma E \ll N$ 

There are exponentially more low energy states than for the quasiparticle case, and D(E) self-averages down to energies exponentially small in N.

## The complex SYK model



Many-body level spacing  $\sim$  $2^{-N} = e^{-N \ln 2}$ 

Non-quasiparticle

excitations with

spacing  $\sim e^{-Ns_0}$ 

There are  $2^N$  many body levels with energy E. Shown are all values of E for a single cluster of size N=12. The  $T\to 0$  state has an entropy  $S_{GPS}=Ns_0$ , where  $s_0<\ln 2$  is determined by integrating

$$\frac{ds_0}{d\mathcal{Q}} = 2\pi\mathcal{E} \,.$$

At 
$$Q = 1/2$$
,

$$s_0 = \frac{G}{\pi} + \frac{\ln(2)}{4} = 0.464848\dots$$

where G is Catalan's constant.

GPS: A. Georges, O. Parcollet, and S. Sachdev, PRB **63**, 134406 (2001)

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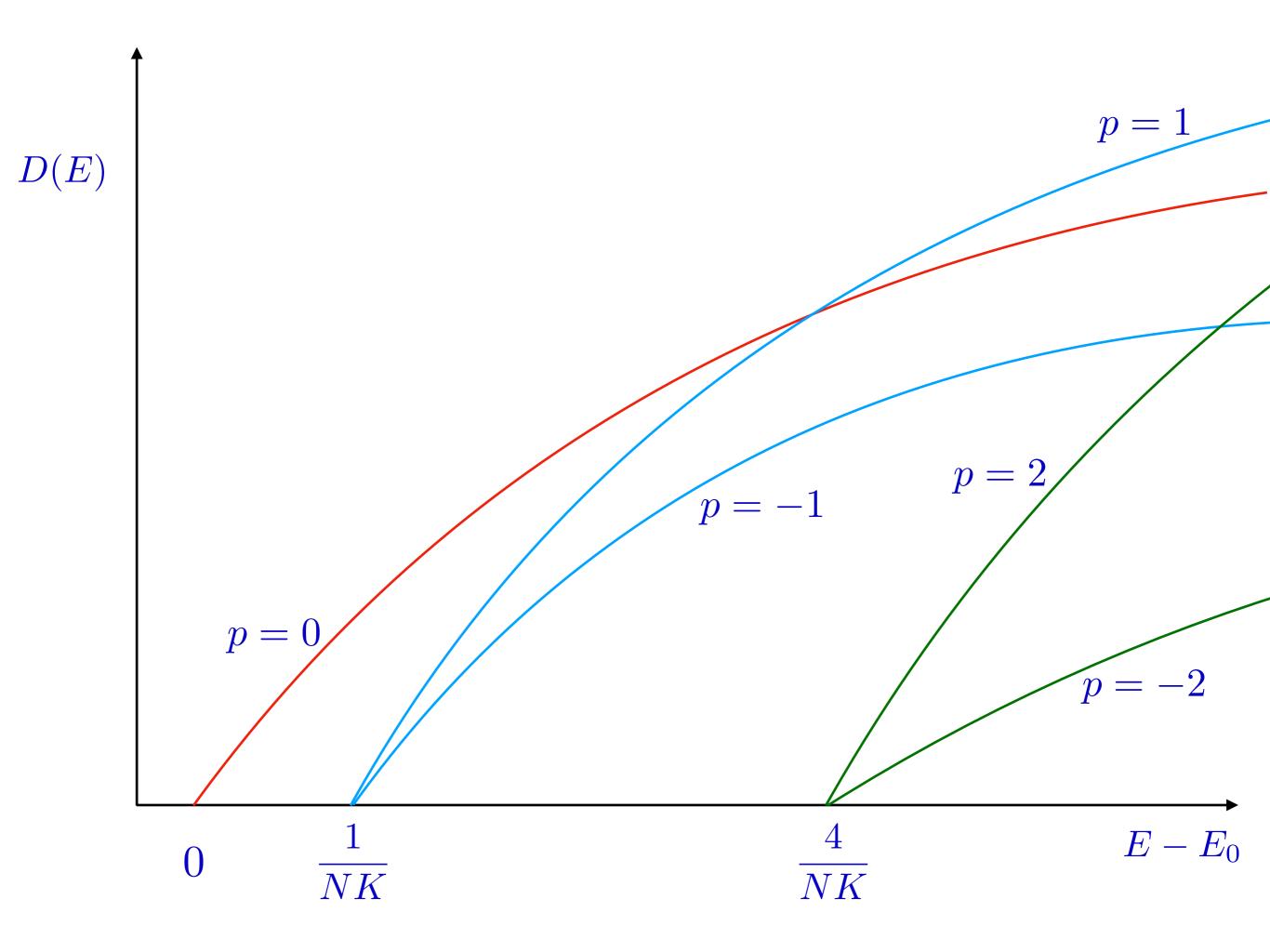
An exact path integral over the effective action leads to the following physical consequences

• At charge NQ + p, the prefactor of the  $\sinh(\sqrt{2N\gamma(E - E_p)})$  term is

$$\exp\left[Ns_0(\mathcal{Q}) + 2\pi p\mathcal{E}\right] \approx \exp\left[Ns_0(\mathcal{Q} + p/N)\right]$$

using

$$\frac{ds_0}{d\mathcal{Q}} = 2\pi\mathcal{E}$$



## The Schwarzian theory and black holes

- Reparameterization invariance is a defining property of Einstein's theory of gravity
- In imaginary time,  $AdS_2$  is the homogeneous hyperbolic space: two-dimensional surface of constant negative curvature. Its metric is invariant under SL(2,R)

$$ds^2 = (d\tau^2 + d\zeta^2)/\zeta^2$$
 is invariant under

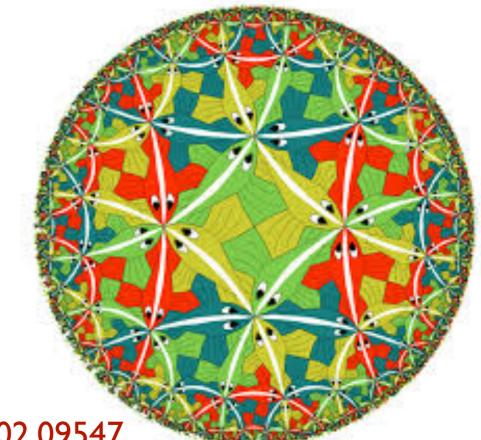
$$\tau' + i\zeta' = \frac{a(\tau + i\zeta) + b}{c(\tau + i\zeta) + d} \text{ with } ad - bc = 1.$$



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Semiclassical fluctuations about the saddle-point of Einstein-Maxwell theory of a charged black holes in  $d \geq 2$  spatial dimensions lead to the same Schwarzian+phase theory of fluctuations.



P. Nayak, A. Shukla, R.M. Soni, S.P. Trivedi, and V. Vishal, arXiv: 1802.09547

U. Moitra, S. P. Trivedi, and V. Vishal, arXiv: 1808.08239

P. Chaturvedi, Yingfei Gu, Wei Song, Boyang Yu, arXiv:1808.08062

A. Gaikwad, L.K. Joshi, G. Mandal, and S.R. Wadia, arXiv:1802.07746

## The Schwarzian theory and black holes

• The Einstein-Maxwell theory leads to the following parameters for the Schwarzian+phase theory

$$K = \left. \frac{d\mathcal{Q}}{d\mu} \right|_{T=0} = \frac{2(d-1)L^2 s_d R_h^{d-3} \left[ d(d+1)R_h^2 + (d-1)^2 L^2 \right]}{(d+1)g_F^2 \kappa^2}$$

$$S(T \to 0, \mathcal{Q}) = s_0 + \gamma T + \dots$$

$$\gamma = \frac{4\pi^2 ds_d L^2 R_h^{d+1}}{\kappa^2 (d(d+1)R_h^2 + (d-1)^2 L^2)}.$$



# Quantum matter without quasiparticles

- Planckian dynamics is realized in the 'solvable' SYK models
- Black holes thermalize in a time  $\sim \hbar/(k_B T_H)$ , where  $T_H$  is the Hawking temperature.
- A Schwarzian theory of a time reparameterization mode, with SL(2,R) symmetry, describes the quantum dynamics of
  - the SYK models
  - black holes with near-extremal AdS<sub>2</sub> horizons