Quantum matter without quasiparticles:
strange metals and black holes

Dartmouth College

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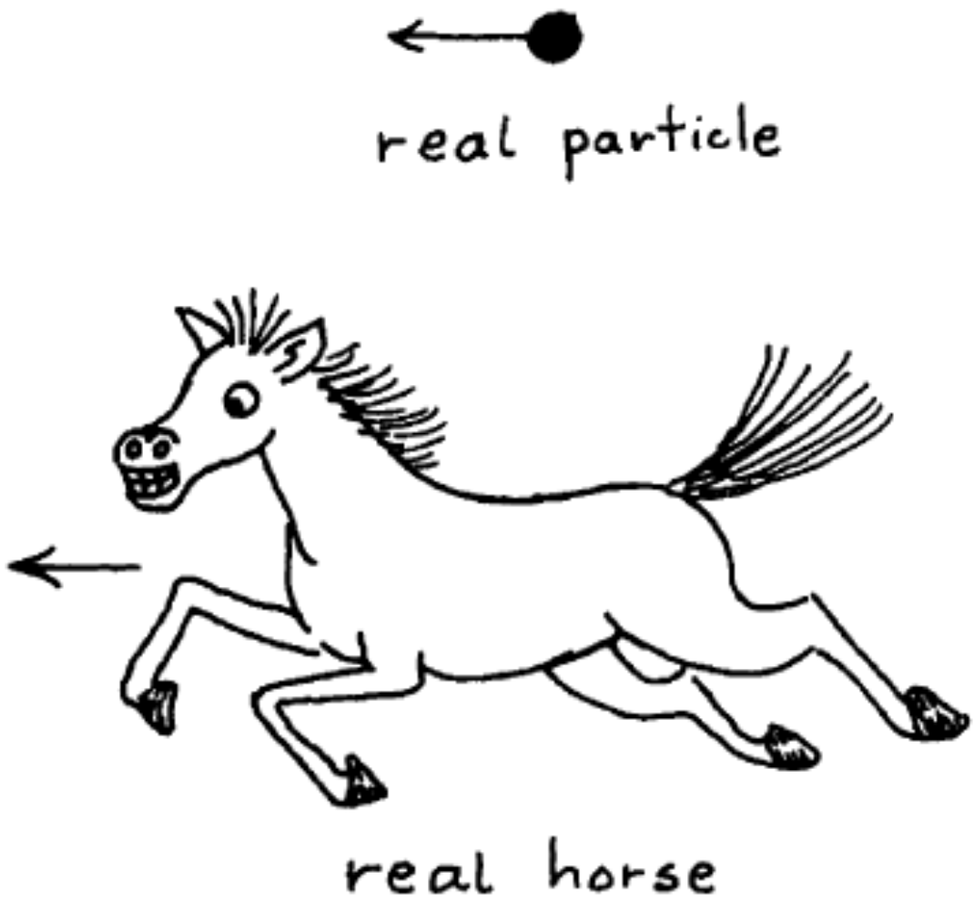
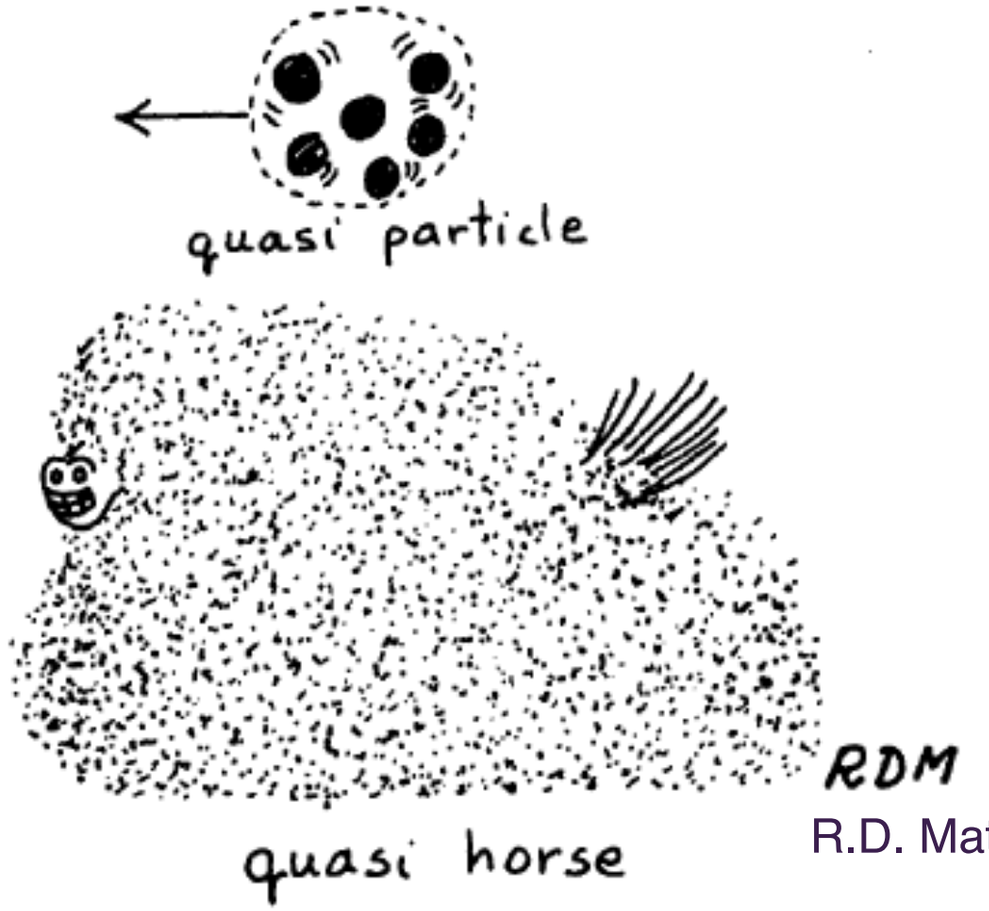
Talk online: sachdev.physics.harvard.edu

Quantum matter <u>with</u> quasiparticles:

A quasiparticle is an "excited lump" in the many-electron state which responds just like an ordinary particle.

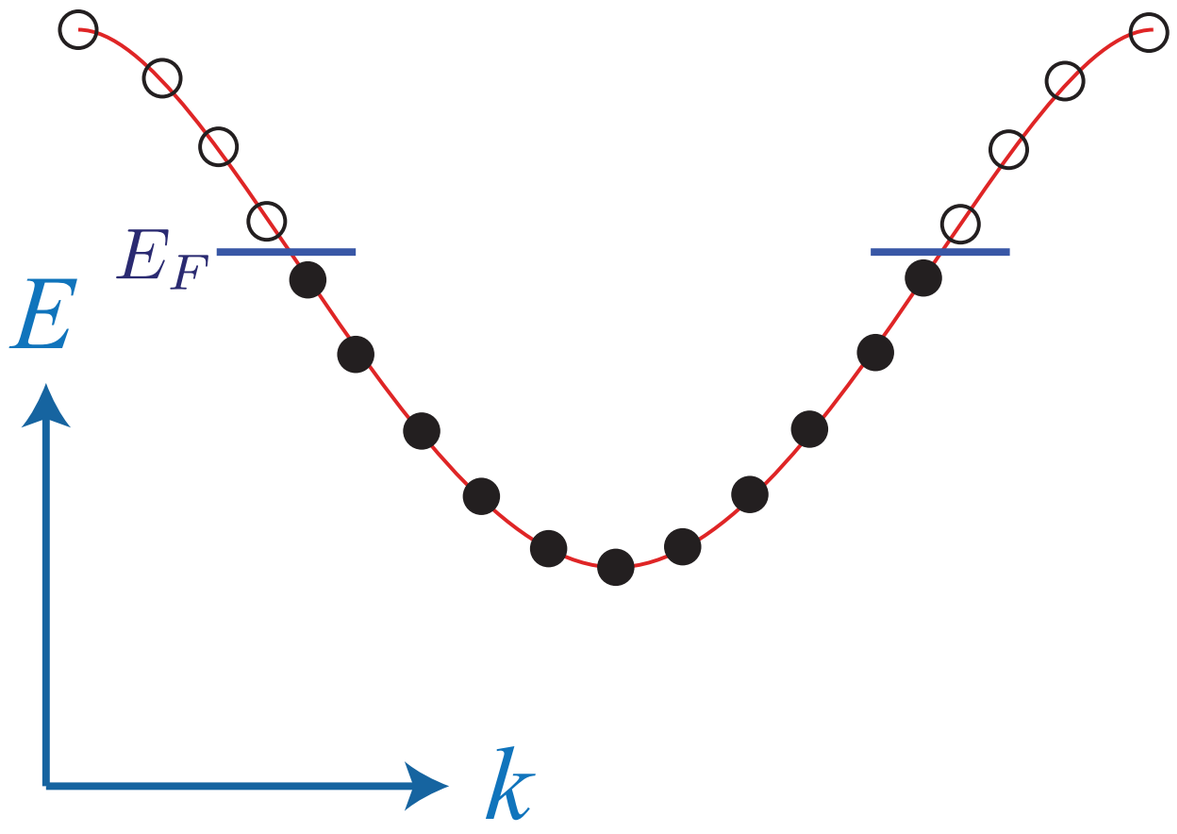




R.D. Mattuck

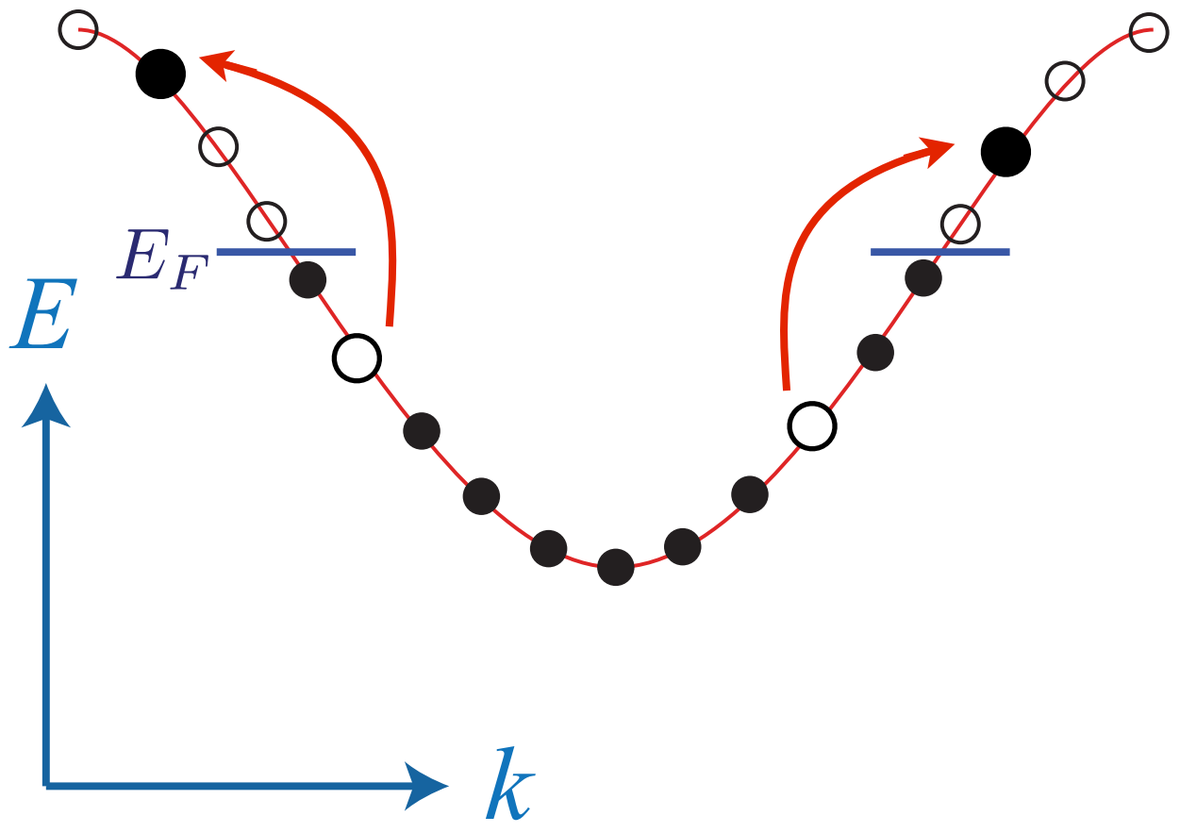
Quantum matter with quasiparticles:

Metals



Quantum matter with quasiparticles:

Metals



Quantum matter with quasiparticles:

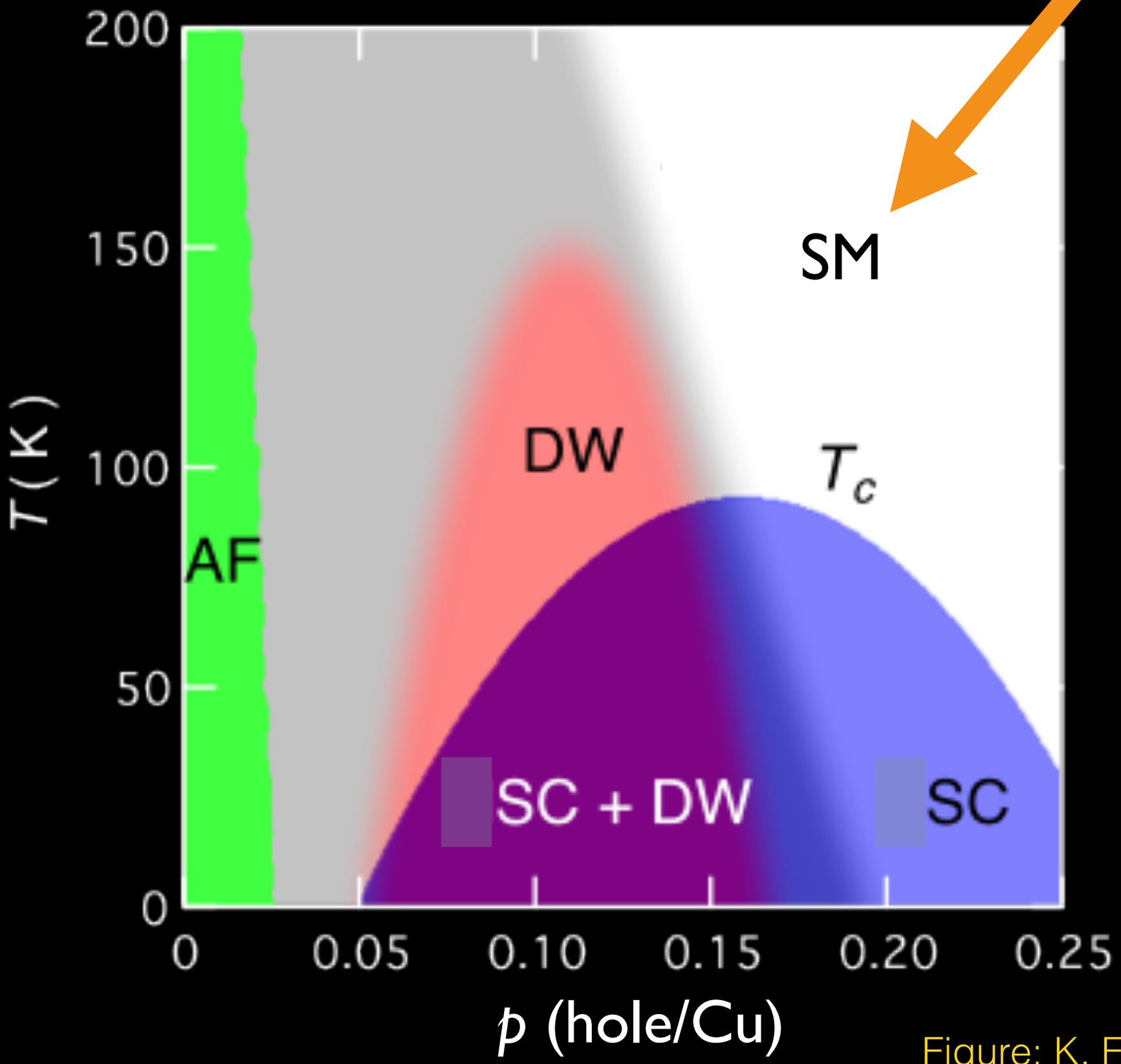
The quasiparticle idea is the key reason for the many successes of quantum condensed matter physics:

- Fermi liquid theory of metals, insulators, semiconductors
- Theory of superconductivity (pairing of quasiparticles)
- Theory of disordered metals and insulators (diffusion and localization of quasiparticles)
- Theory of metals in one dimension (collective modes as quasiparticles)
- Theory of the fractional quantum Hall effect (quasiparticles which are `fractions' of an electron)

Quantum matter <u>without</u> quasiparticles

Strange metal

Entangled electrons lead to "strange" temperature dependence of resistivity and other properties



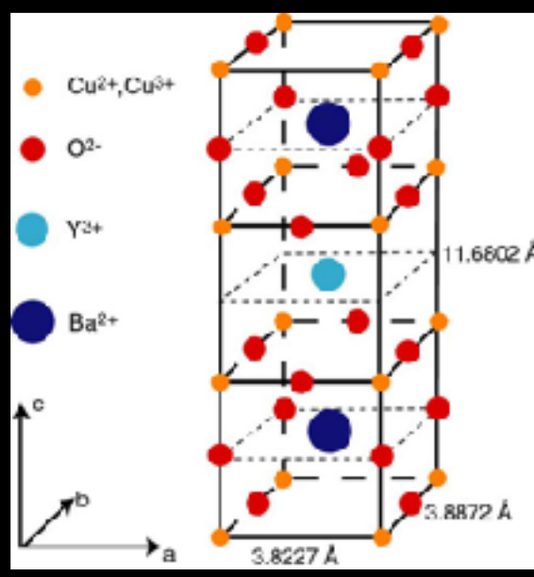


Figure: K. Fujita and J. C. Seamus Davis

"Strange",

"Bad",

or "Incoherent",

metal has a resistivity, ρ, which obeys

$$\rho \sim T,$$

and

$$\rho \gg h/e^2$$

(in two dimensions),
where h/e^2 is the quantum unit of resistance.

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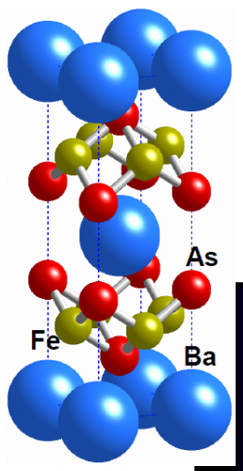
Quantum matter <u>without</u> quasiparticles

Resistivity
$\sim \rho_0 + AT^\alpha$



Strange
(or "bad"
or "incoherent")
Metal
$\rho \sim T$
$\rho \gg h/e^2$

$BaFe_2(As_{1-x}P_x)_2$

AF
+nematic

Superconductivity

T (K)
200
100
0

0
0.2
0.4
0.6
x

α
2.0
1.0
0

Superconductivity in Bad Metals

V. J. Emery and S. A. Kivelson
Phys. Rev. Lett. **74**, 3253 – Published 17 April 1995

S. Kasahara, T. Shibauchi, K. Hashimoto, K. Ikada, S. Tonegawa, R. Okazaki, H. Shishido, H. Ikeda, H. Takeya, K. Hirata, T. Terashima, and Y. Matsuda, *PRB* **81**, 184519 (2010)

Quantum matter with quasiparticles:

- **Quasiparticles are additive excitations:**
  The low-lying excitations of the many-body system can be identified as a set $\{n_\alpha\}$ of quasiparticles with energy ε_α

$$E = \sum_\alpha n_\alpha \varepsilon_\alpha + \sum_{\alpha,\beta} F_{\alpha\beta} n_\alpha n_\beta + \dots$$

In a lattice system of N sites, this parameterizes the energy of $\sim e^{\alpha N}$ states in terms of poly(N) numbers.

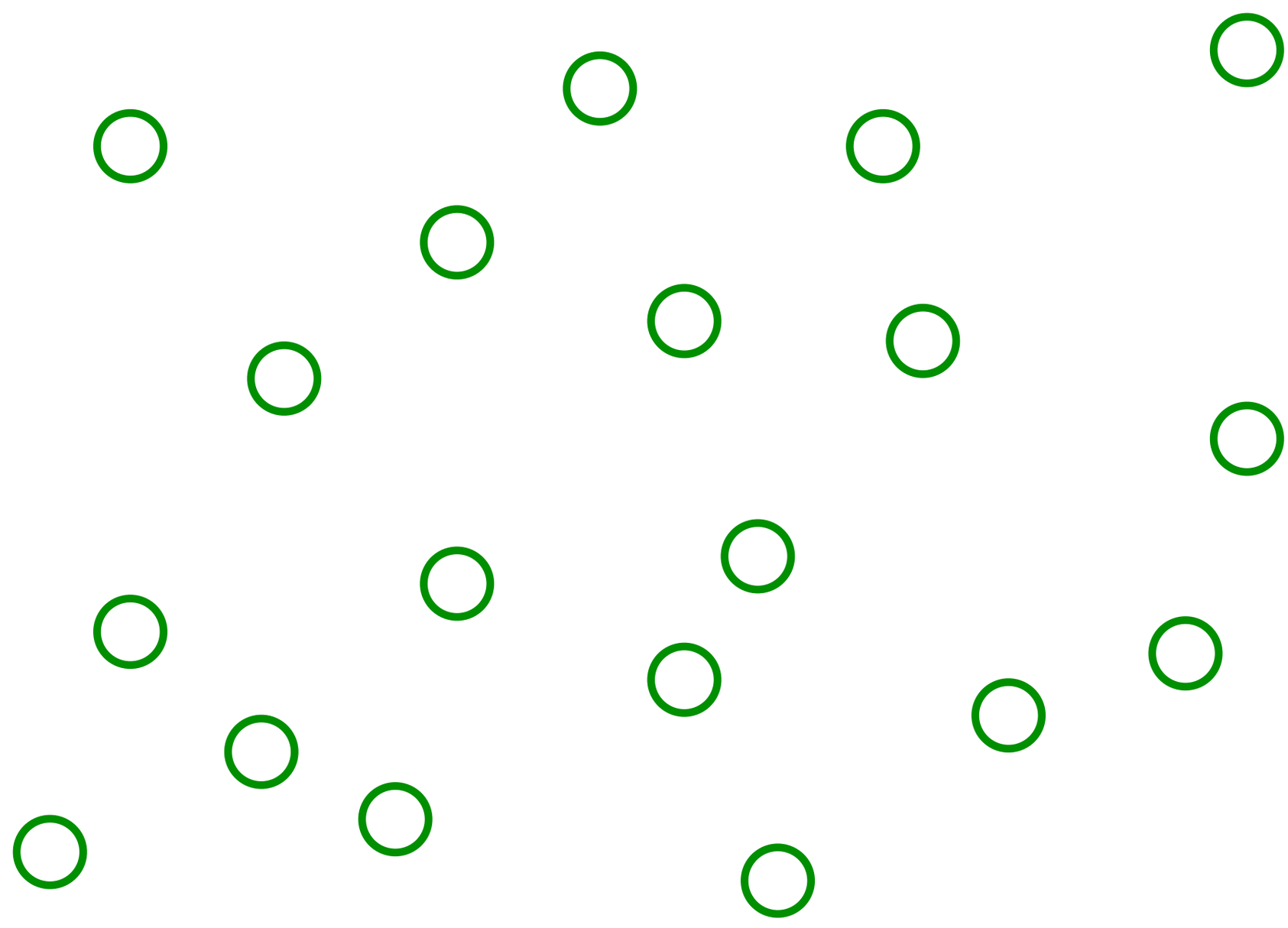
Quantum matter <u>with</u> quasiparticles:

- Quasiparticles eventually collide with each other. Such collisions eventually leads to thermal equilibration in a chaotic quantum state, but the equilibration takes a long time. In a Fermi liquid, this time diverges as

$$\tau_{\rm eq} \sim \frac{\hbar E_F}{(k_B T)^2} \quad , \quad \text{as } T \to 0,$$

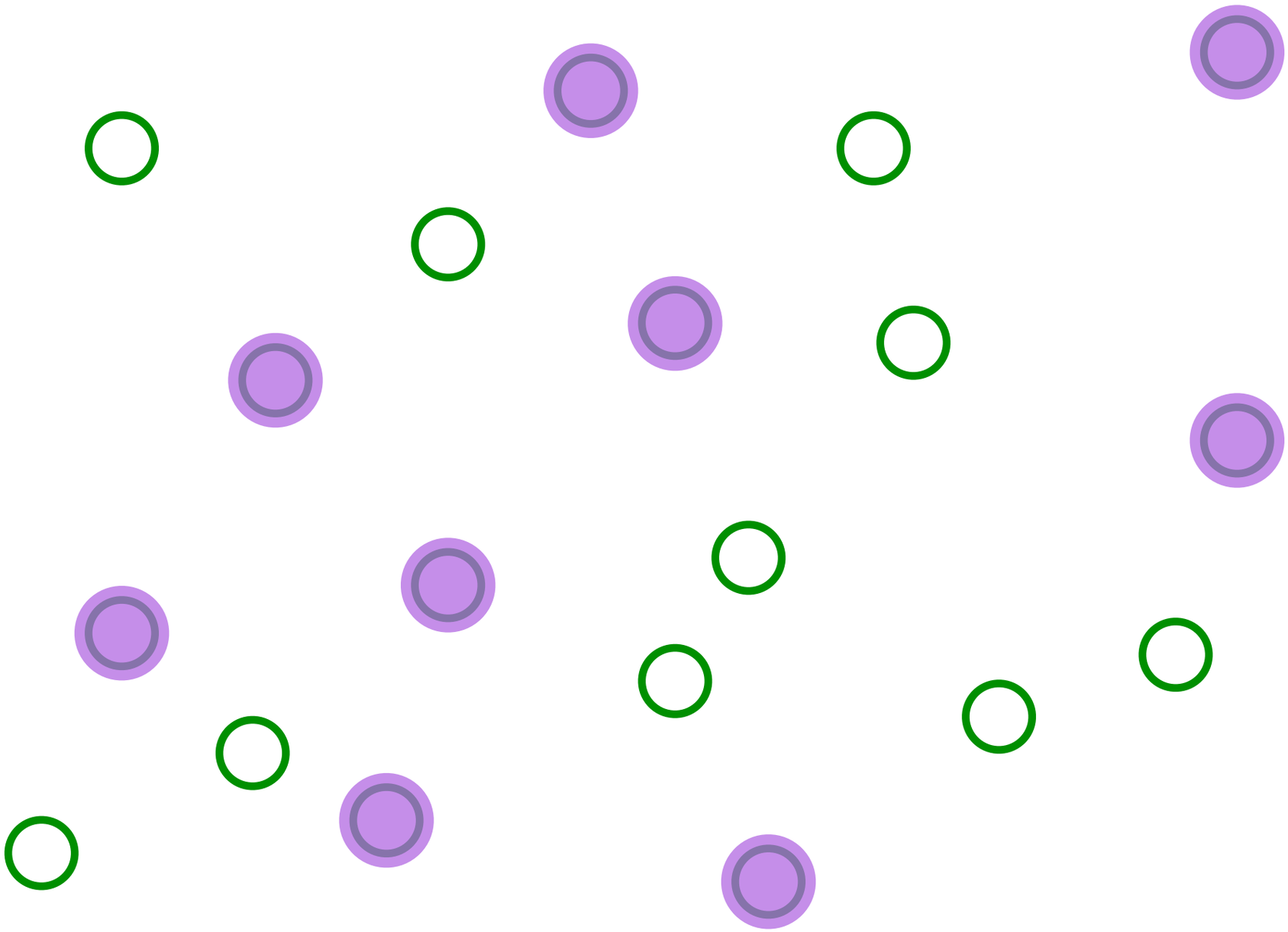
where E_F is the Fermi energy.

A simple model of a metal with quasiparticles



Pick a set of random positions

A simple model of a metal with quasiparticles



Place electrons randomly on some sites

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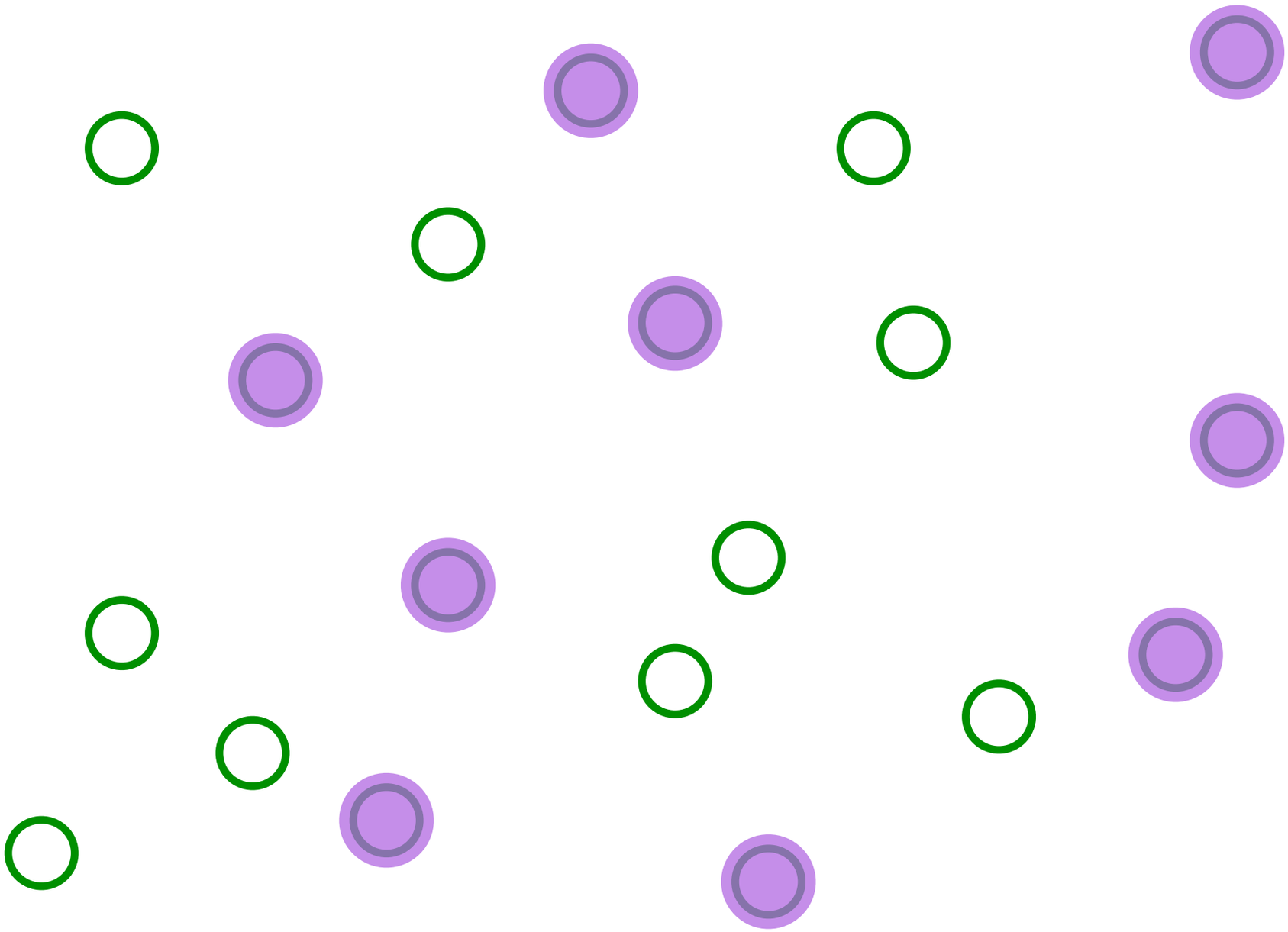

Electrons move one-by-one randomly

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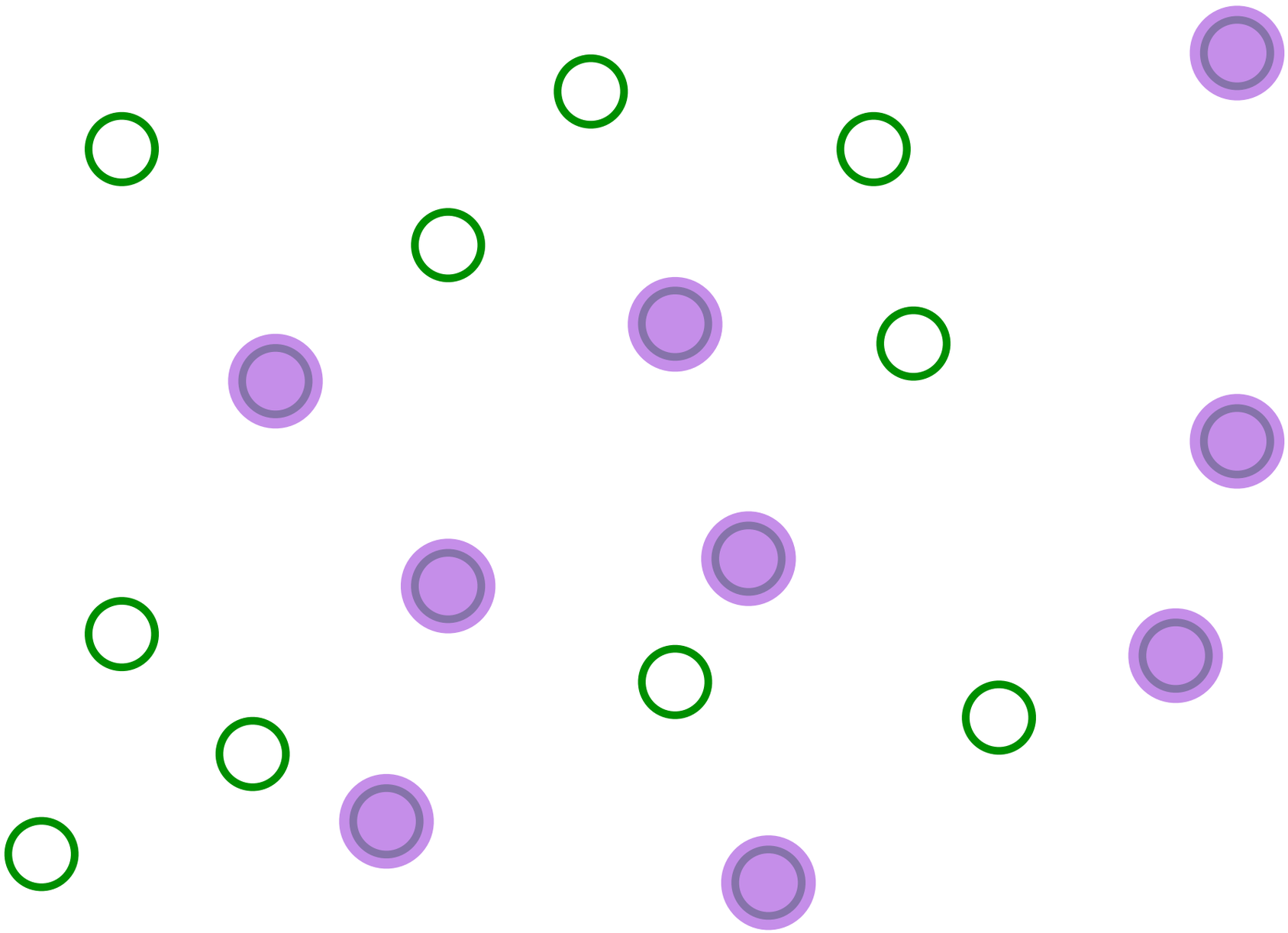

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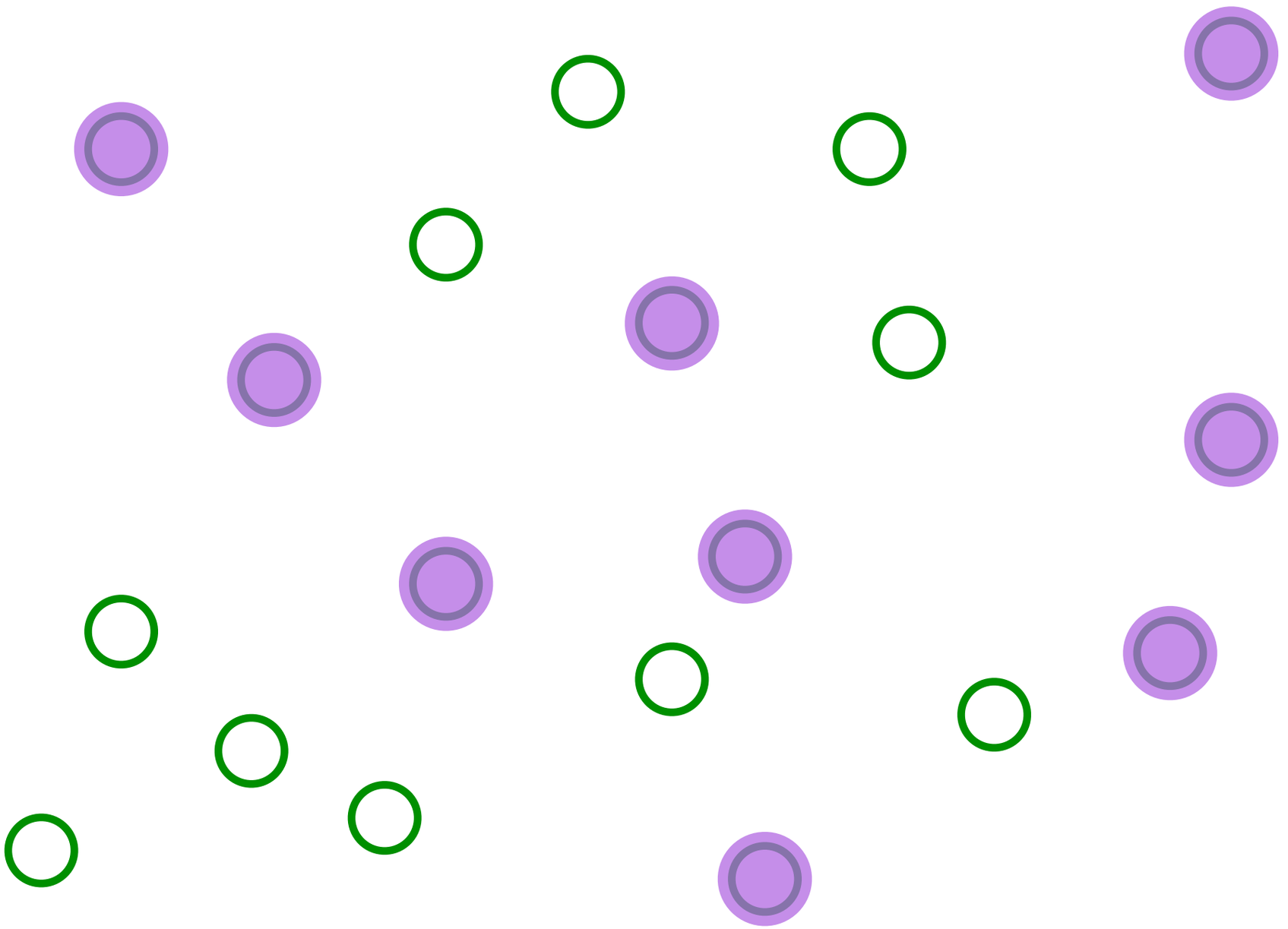

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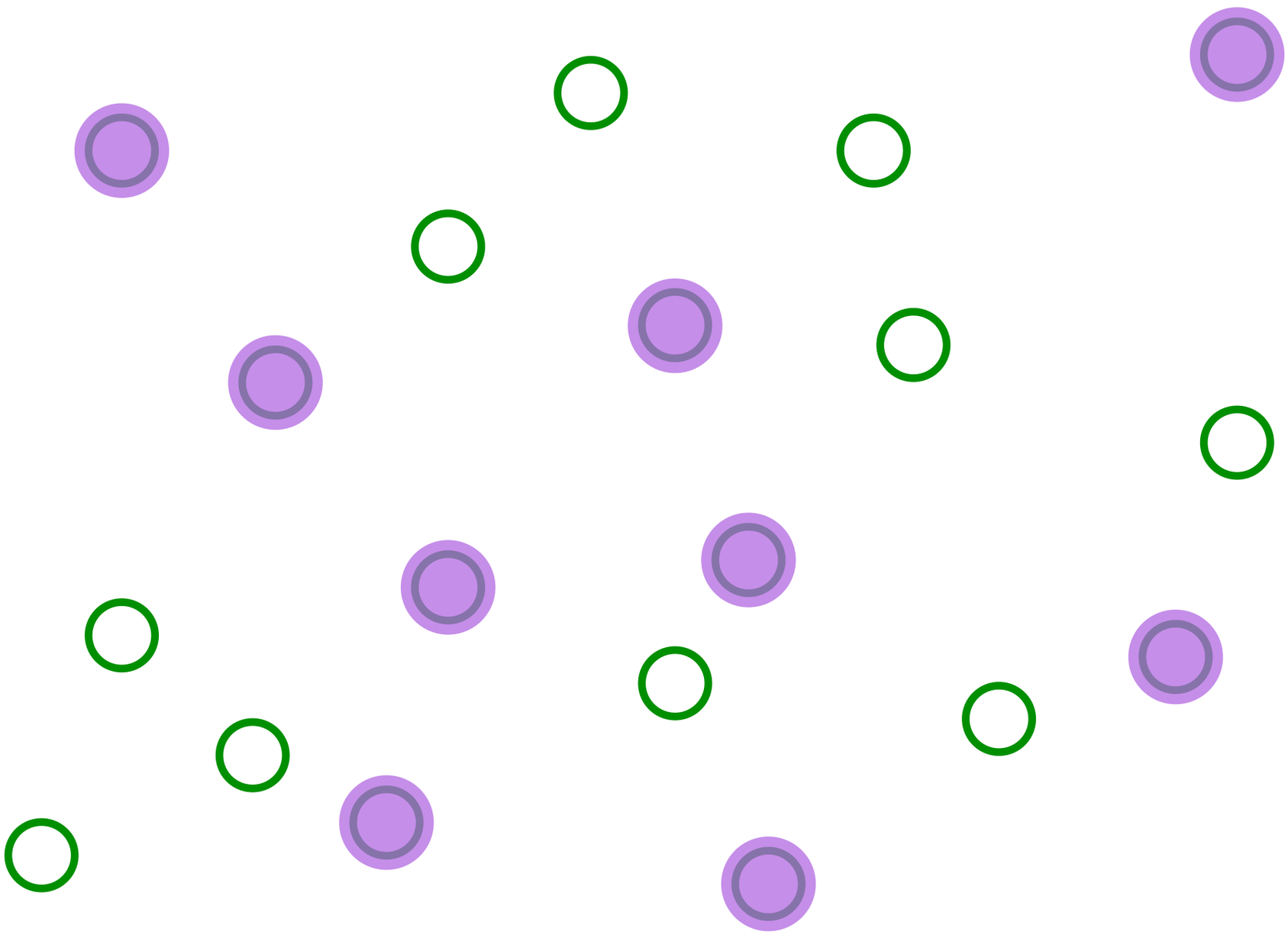


Electrons move one-by-one randomly

A simple model of a metal with quasiparticles

$$H = \frac{1}{(N)^{1/2}} \sum_{i,j=1}^{N} t_{ij} c_i^\dagger c_j + \dots$$

$$c_i c_j + c_j c_i = 0 \quad , \quad c_i c_j^\dagger + c_j^\dagger c_i = \delta_{ij}$$

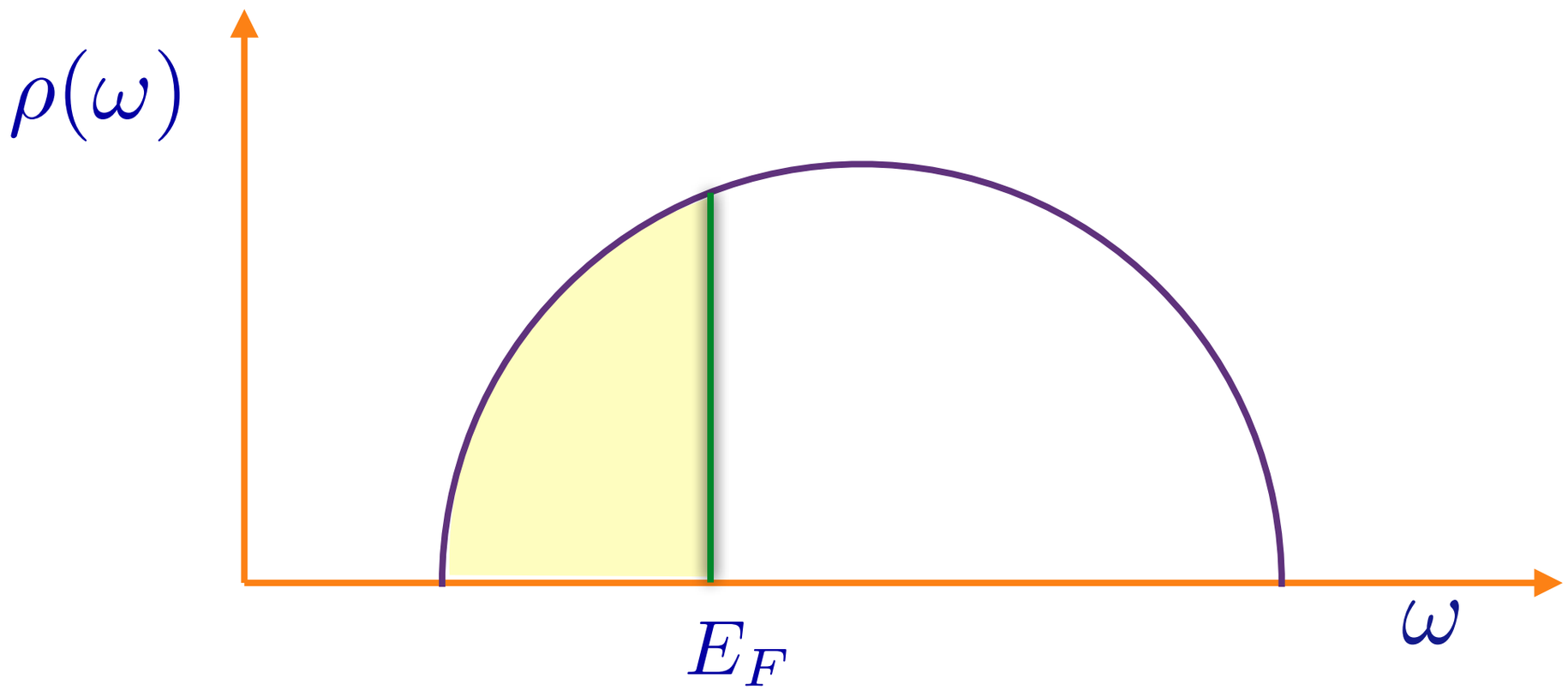
$$\frac{1}{N} \sum_i c_i^\dagger c_i = Q$$

t_{ij} are independent random variables with $\overline{t_{ij}} = 0$ and $\overline{|t_{ij}|^2} = t^2$

Fermions occupying the eigenstates of a
N x *N* random matrix

A simple model of a metal with quasiparticles

Let ε_α be the eigenvalues of the matrix $t_{ij}/\sqrt{N}$. The fermions will occupy the lowest NQ eigenvalues, upto the Fermi energy E_F. The density of states is $\rho(\omega) = (1/N)\sum_\alpha \delta(\omega - \varepsilon_\alpha)$.



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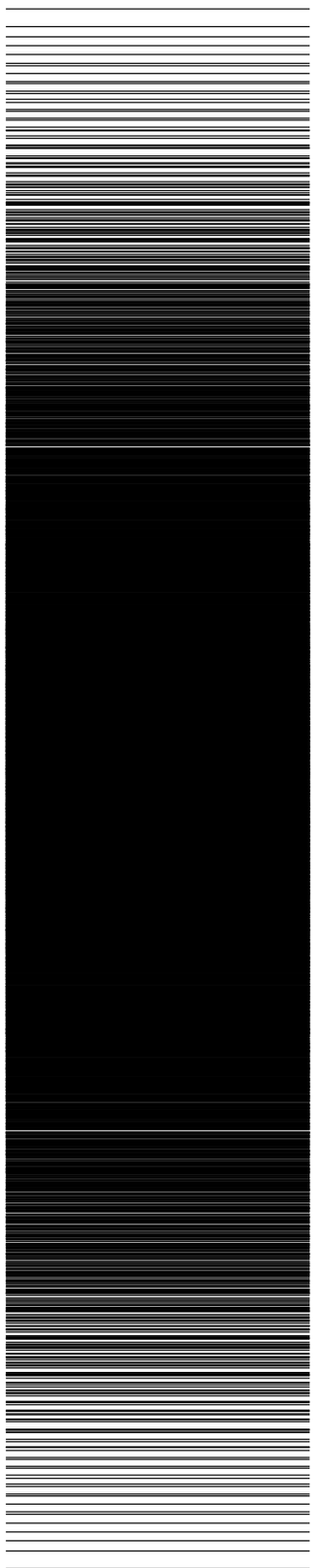

Many-body level spacing $\sim 2^{-N}$

Quasiparticle excitations with spacing $\sim 1/N$

There are 2^N many body levels with energy

$$E = \sum_{\alpha=1}^{N} n_\alpha \varepsilon_\alpha ,$$

where $n_\alpha = 0, 1$. Shown are all values of E for a single cluster of size $N = 12$. The ε_α have a level spacing $\sim 1/N$.

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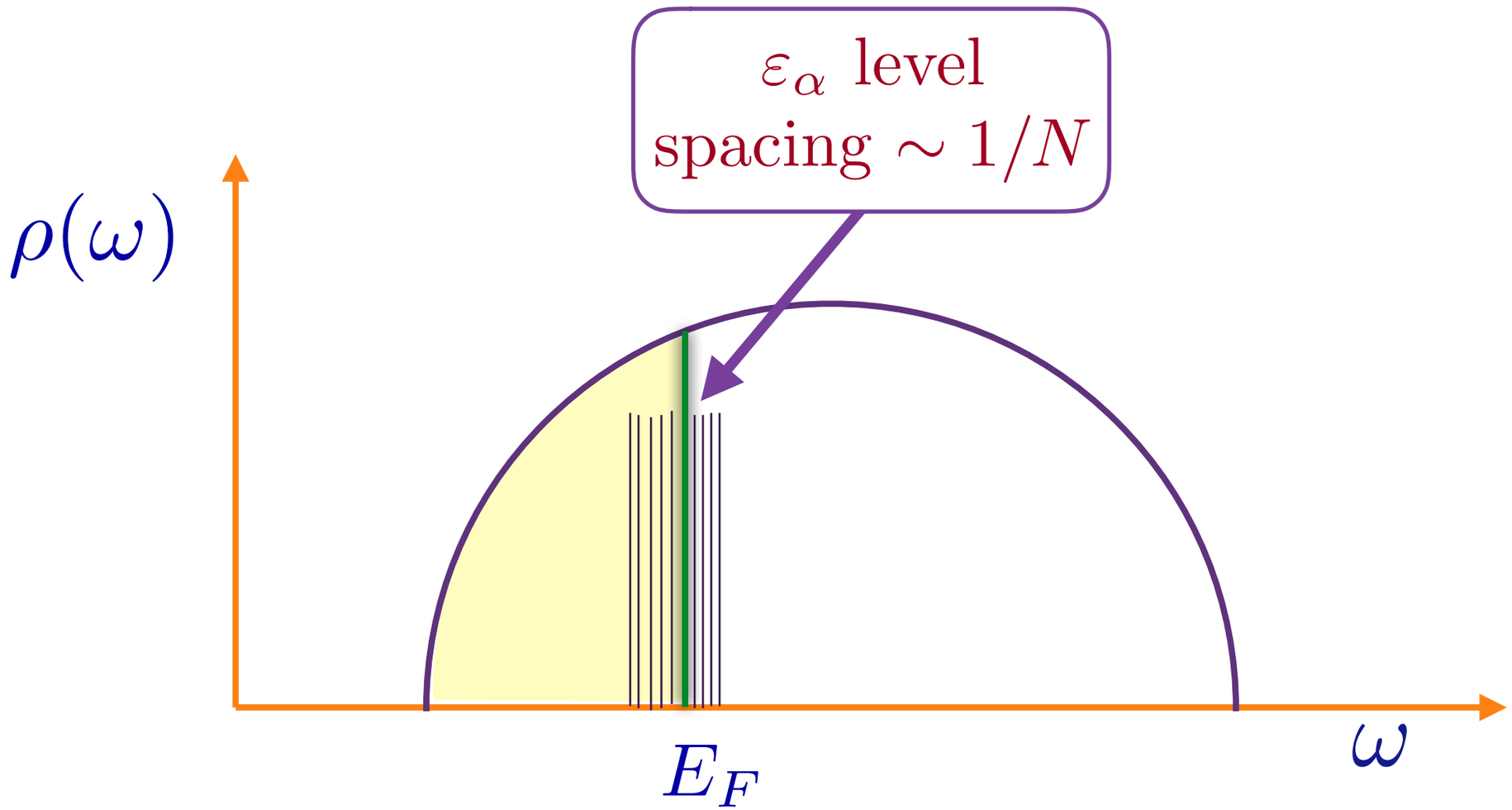

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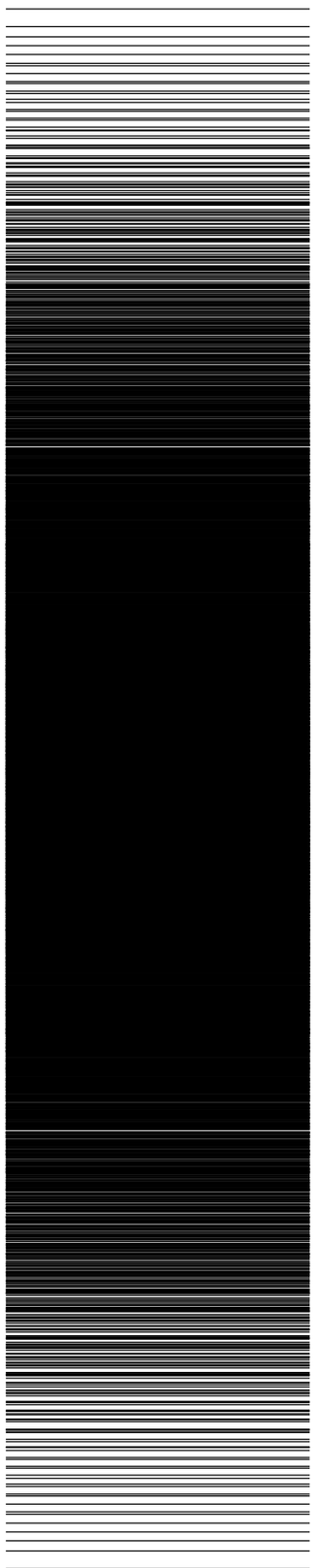

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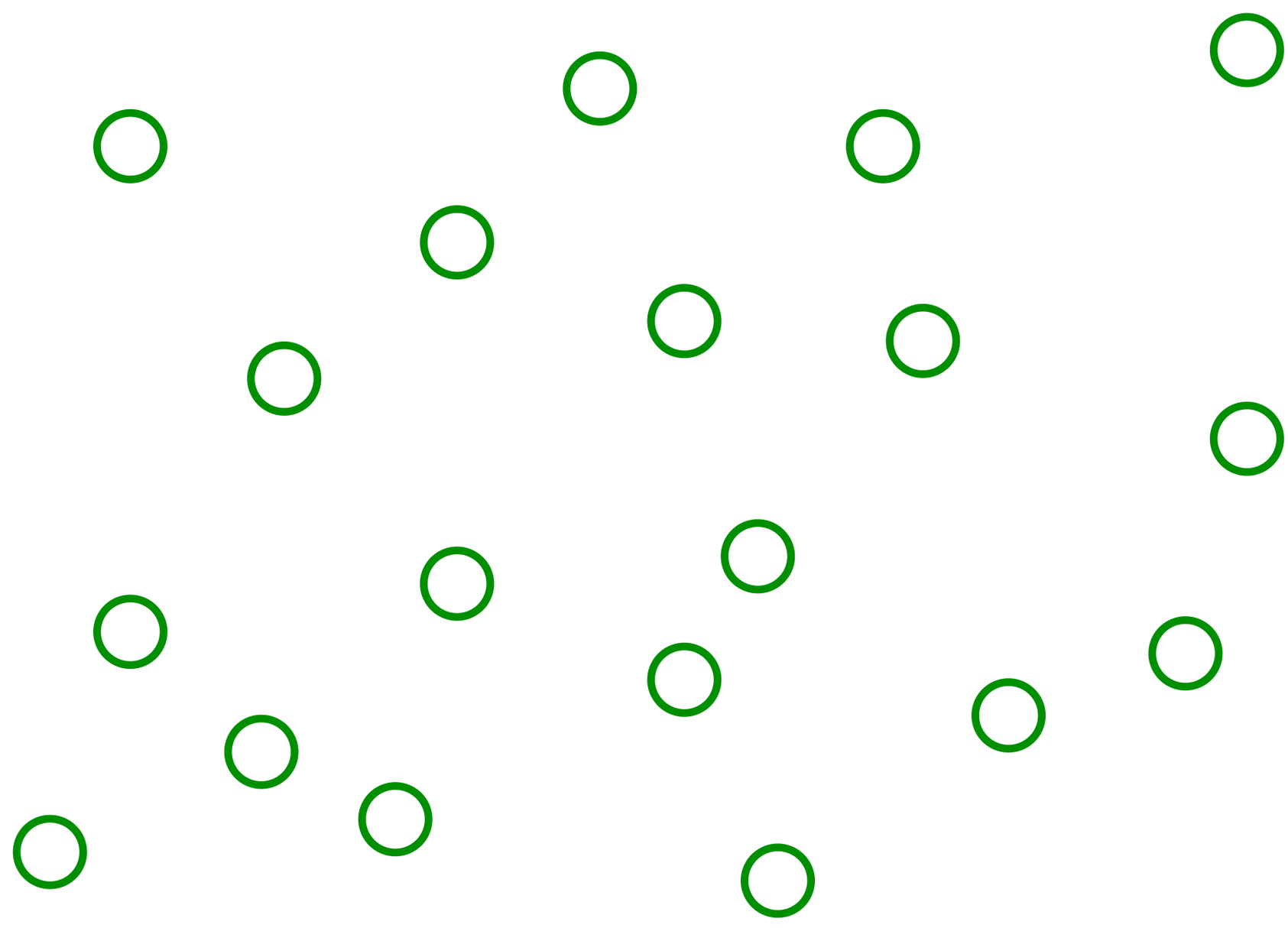
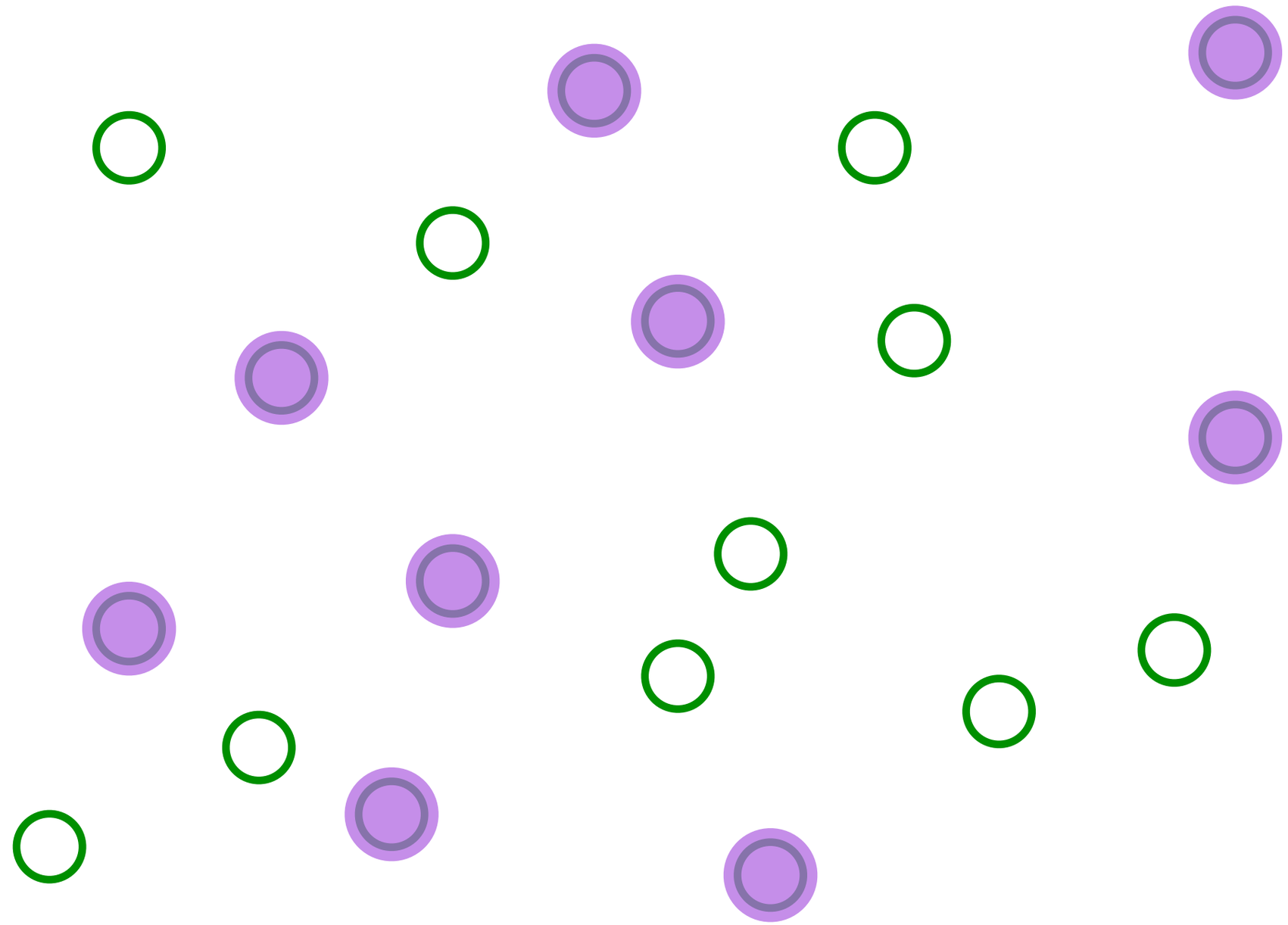
The Sachdev-Ye-Kitaev (SYK) model



Pick a set of random positions

The SYK model



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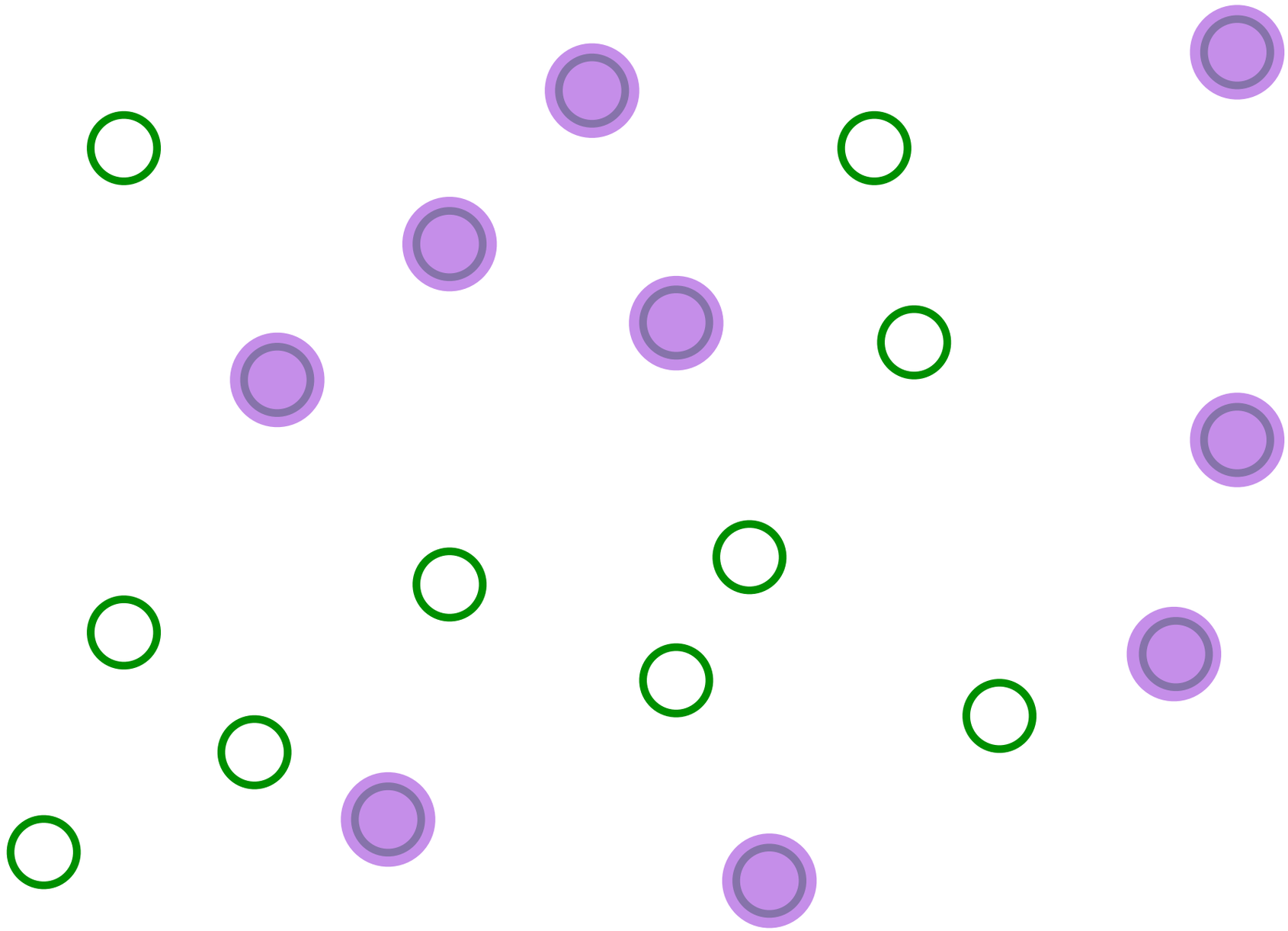
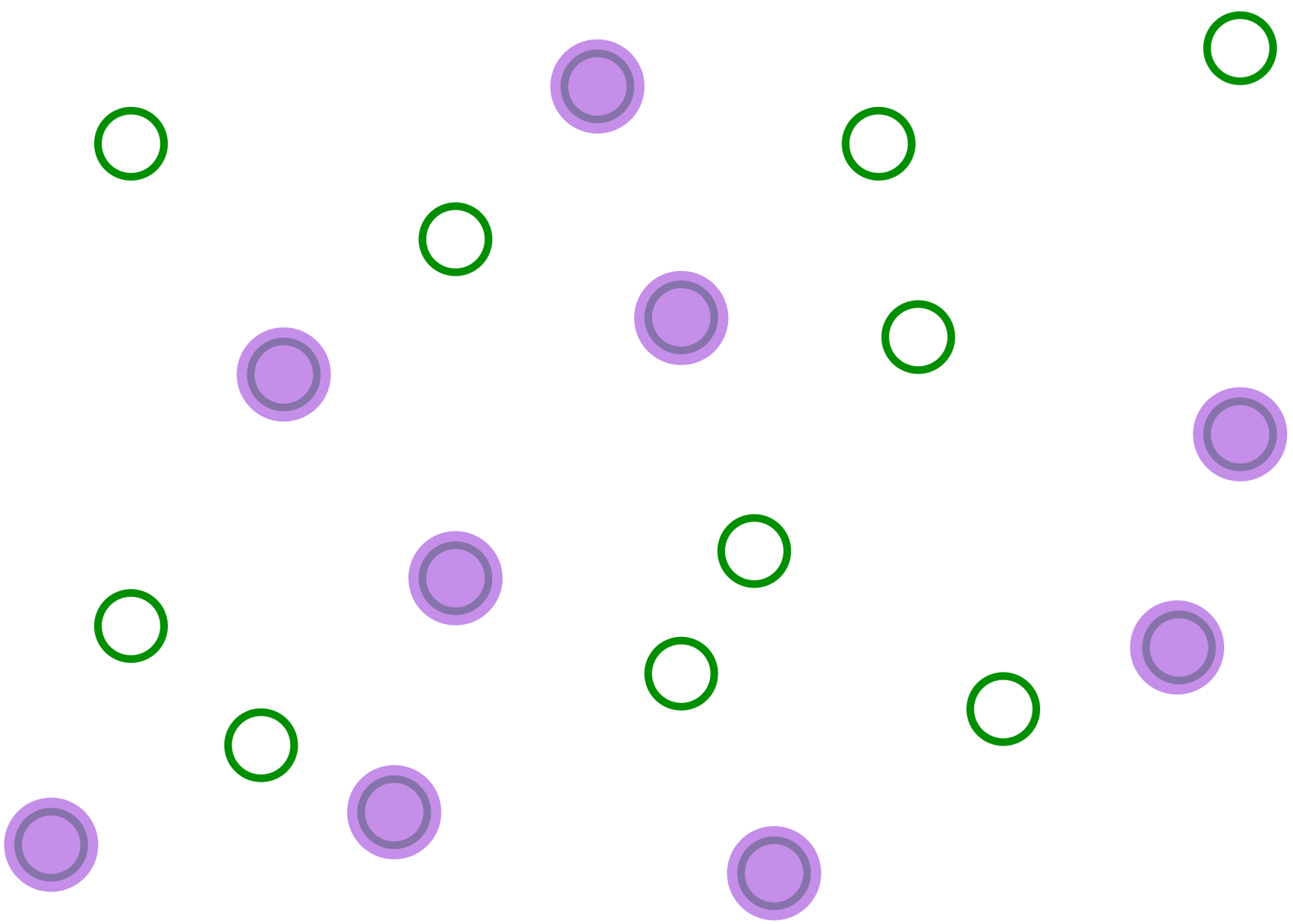


Entangle electrons pairwise randomly

The SYK model



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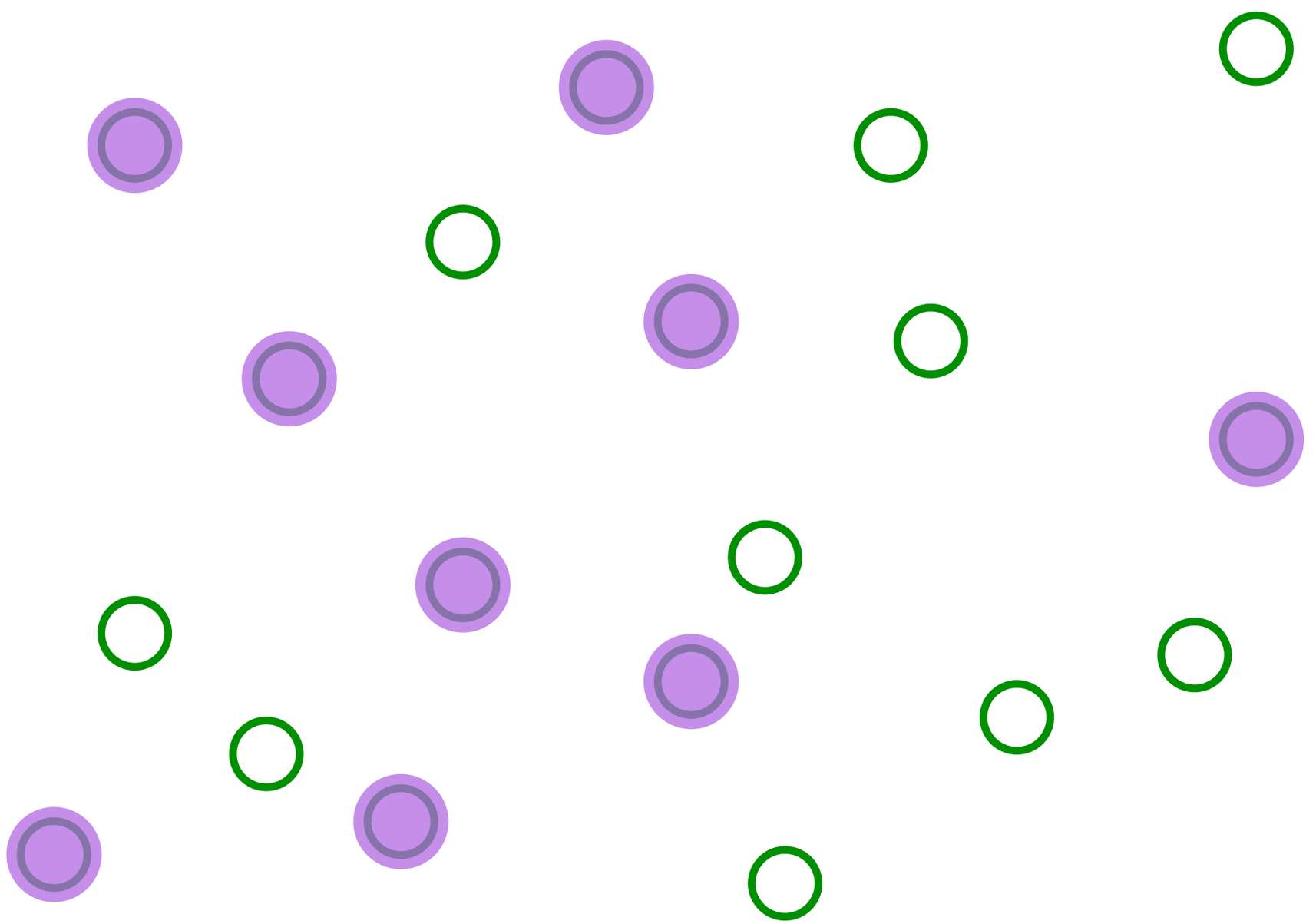

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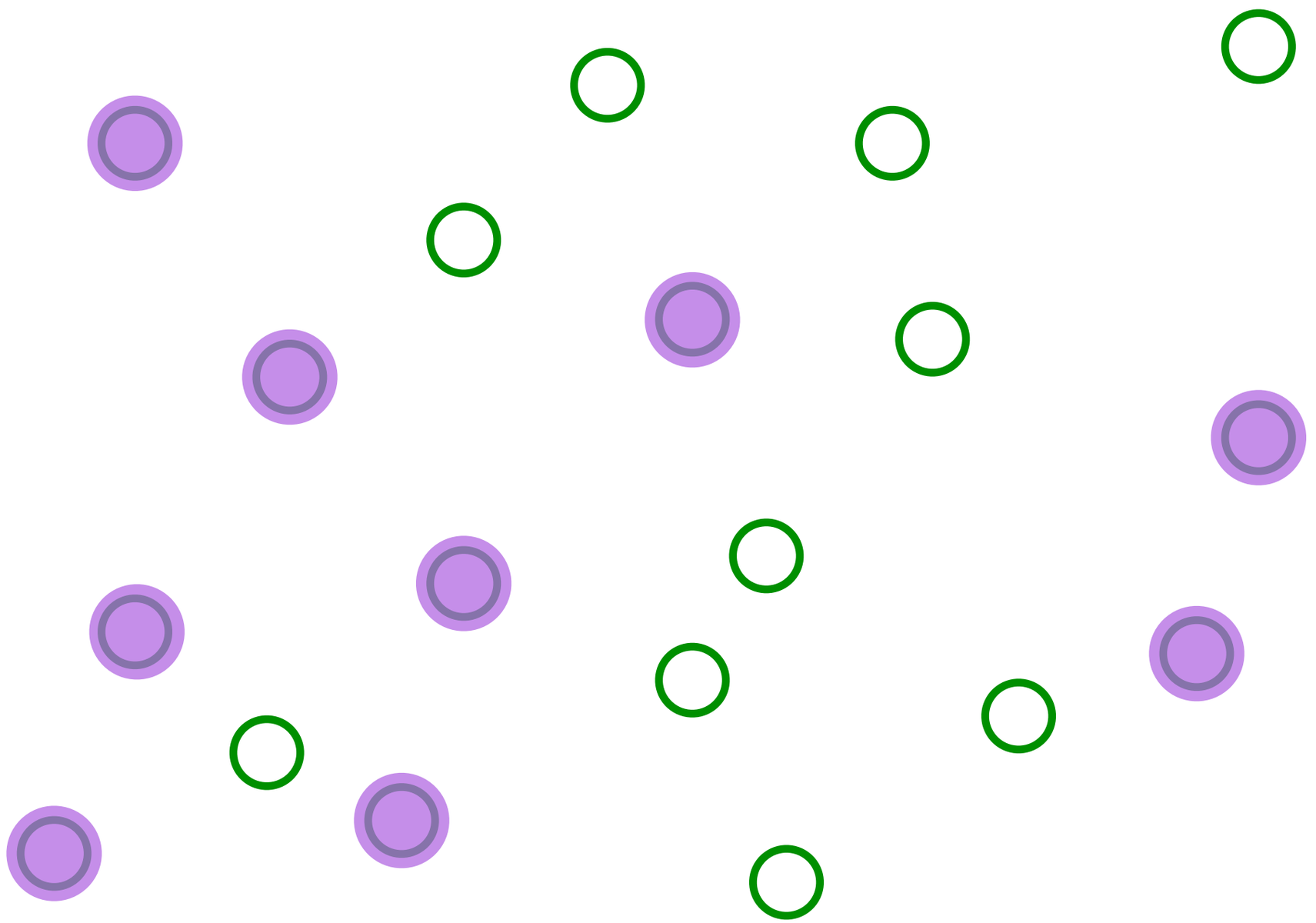

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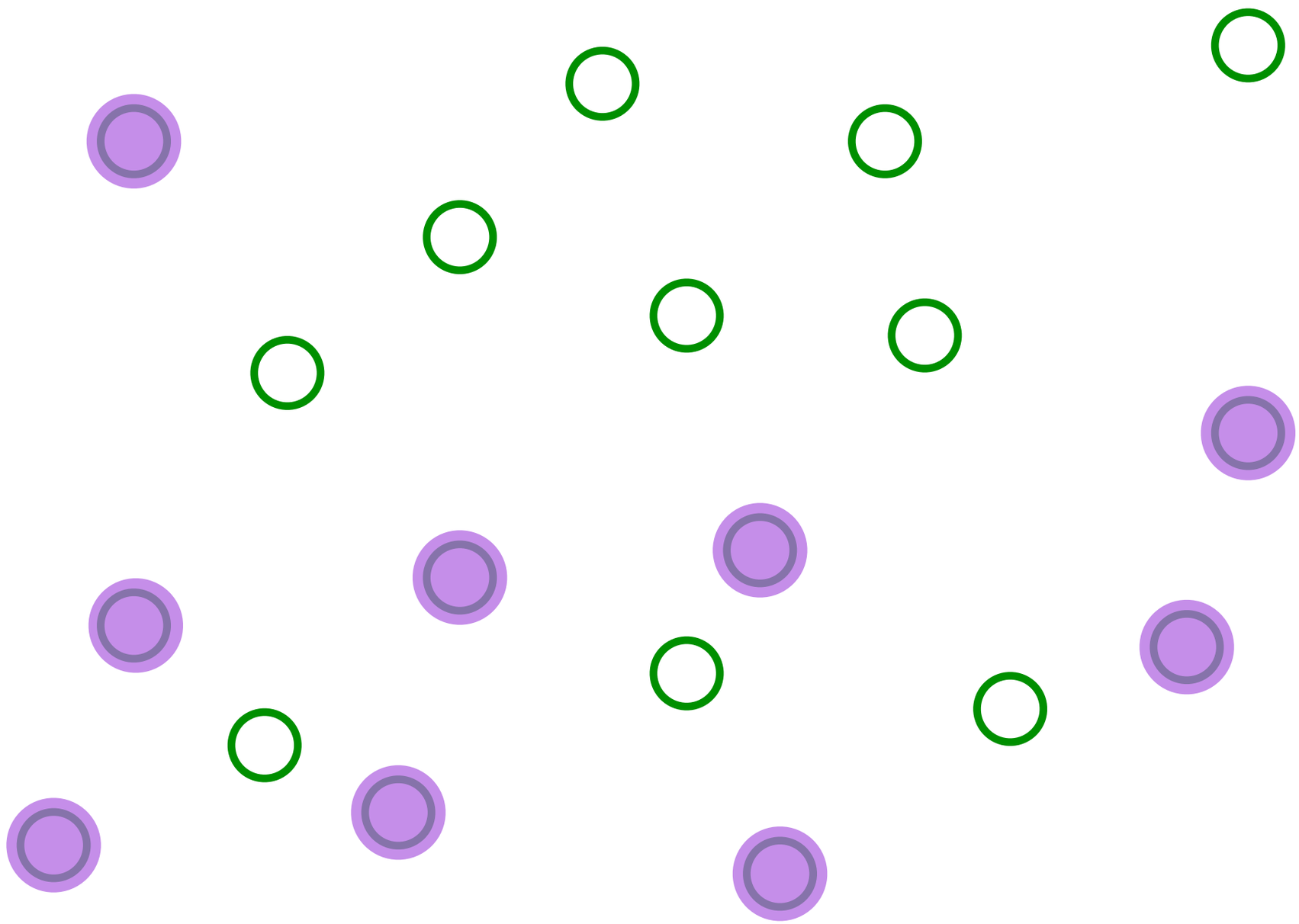

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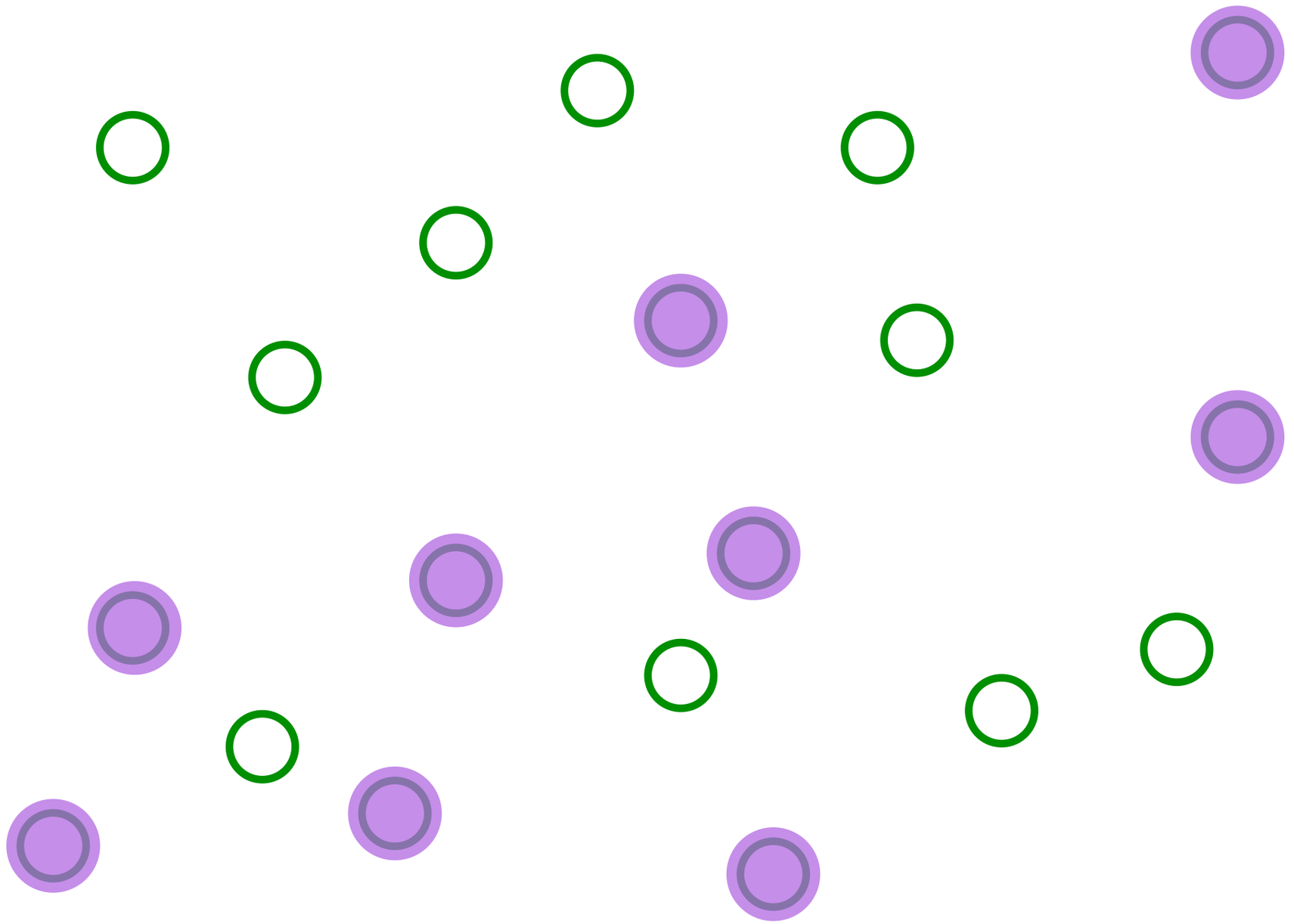

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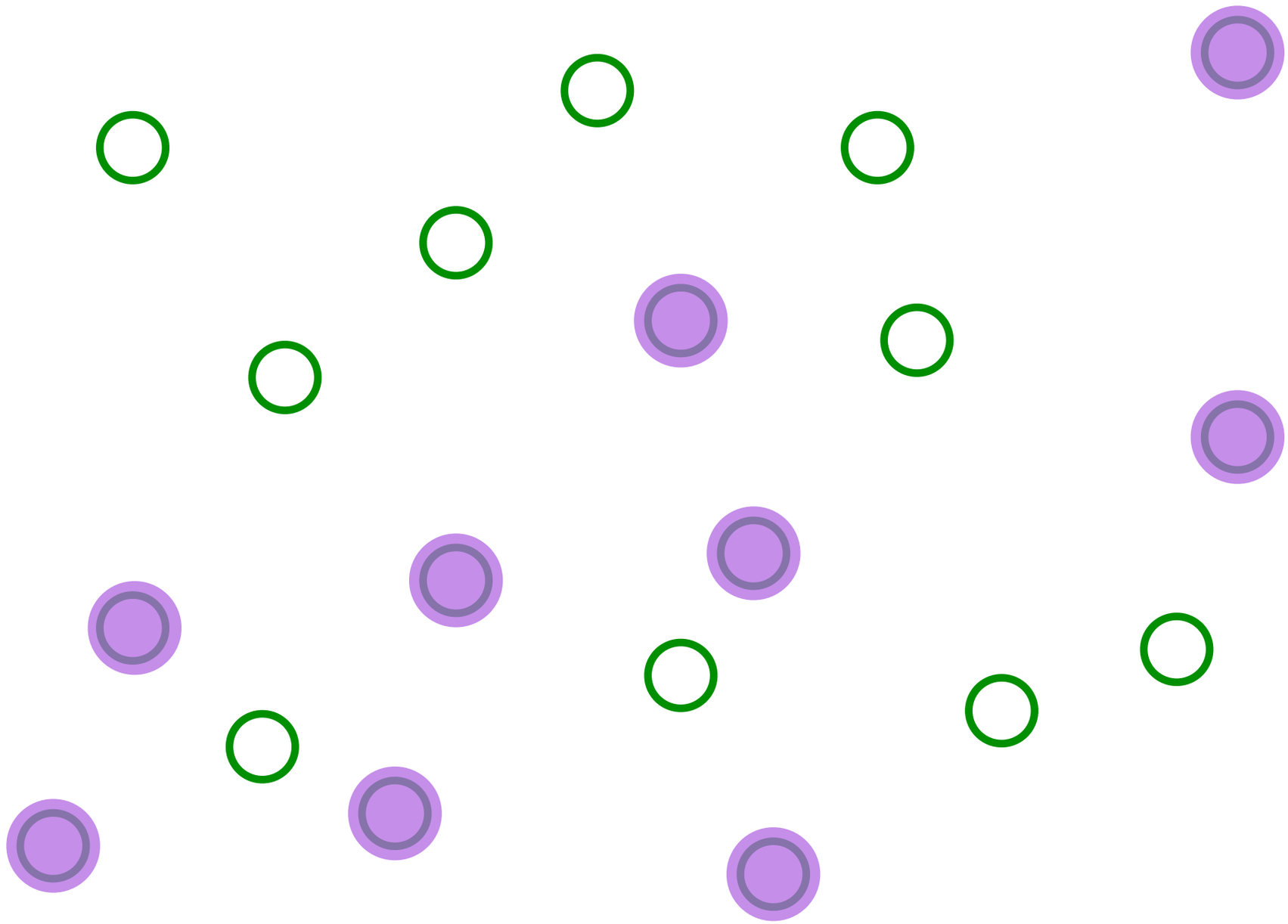


This describes both a strange metal and a black hole!

The SYK model

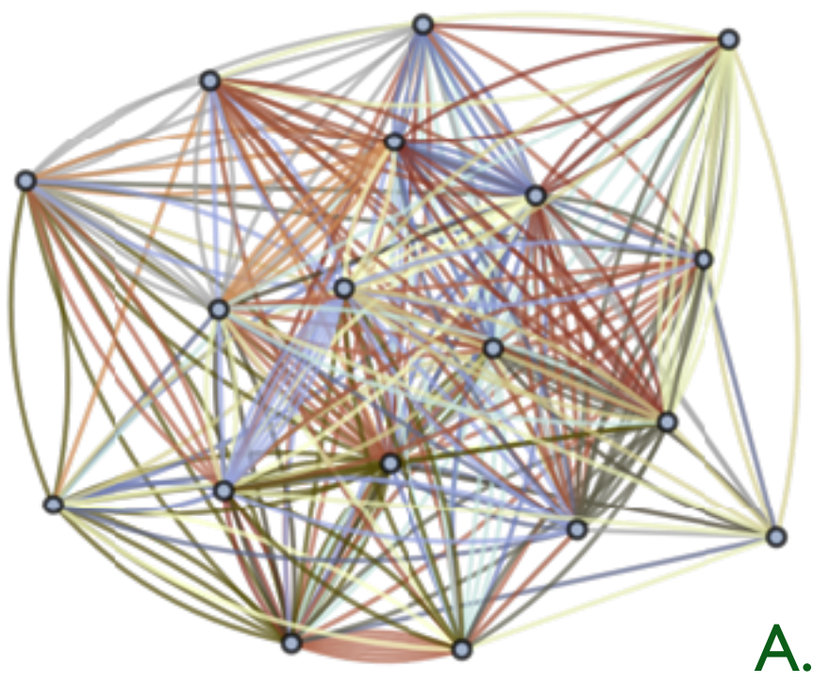
(See also: the "2-Body Random Ensemble" in nuclear physics; did not obtain the large N limit;
T.A. Brody, J. Flores, J.B. French, P.A. Mello, A. Pandey, and S.S.M. Wong, Rev. Mod. Phys. **53**, 385 (1981))

$$H = \frac{1}{(2N)^{3/2}} \sum_{i,j,k,\ell=1}^{N} U_{ij;k\ell}\, c_i^\dagger c_j^\dagger c_k c_\ell - \mu \sum_i c_i^\dagger c_i$$

$$c_i c_j + c_j c_i = 0 \quad , \quad c_i c_j^\dagger + c_j^\dagger c_i = \delta_{ij}$$

$$Q = \frac{1}{N} \sum_i c_i^\dagger c_i$$

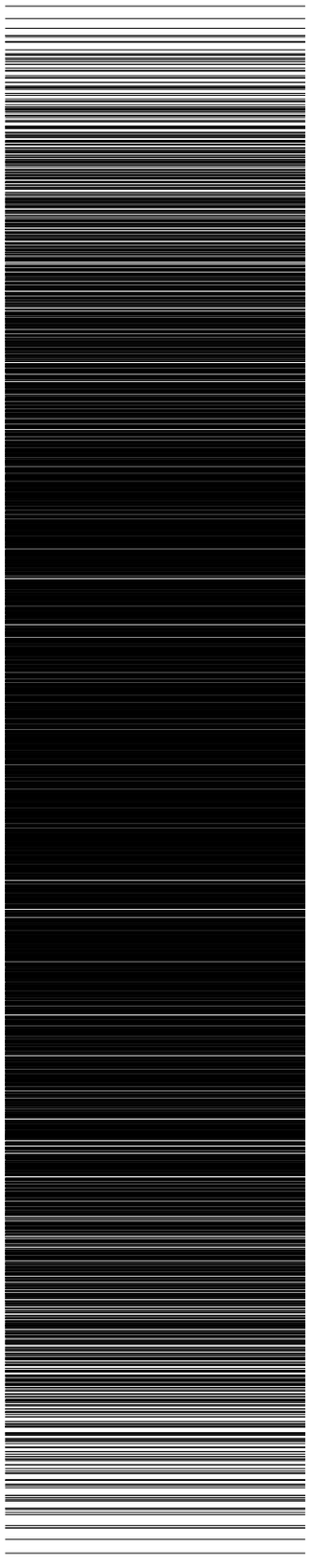
$U_{ij;k\ell}$ are independent random variables with $\overline{U_{ij;k\ell}} = 0$ and $\overline{|U_{ij;k\ell}|^2} = U^2$
$N \to \infty$ yields critical strange metal.



S. Sachdev and J. Ye, PRL **70**, 3339 (1993)

A. Kitaev, unpublished; S. Sachdev, PRX **5**, 041025 (2015)

The SYK model



Many-body level spacing $\sim 2^{-N} = e^{-N \ln 2}$

Non-quasiparticle excitations with spacing $\sim e^{-N s_0}$

There are 2^N many body levels with energy E, which do not admit a quasiparticle decomposition. Shown are all values of E for a single cluster of size $N = 12$. The $T \to 0$ state has an entropy $S_{GPS} = N s_0$ with

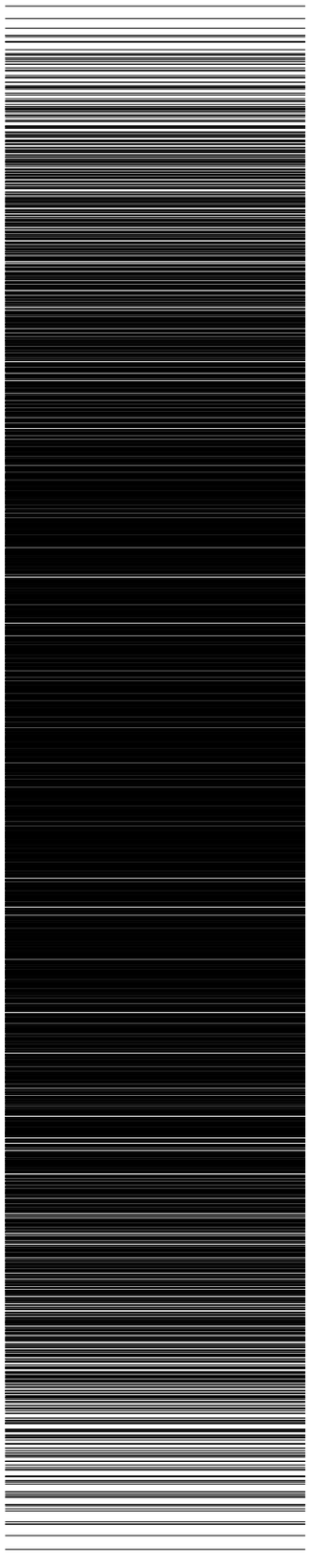
$$s_0 = \frac{G}{\pi} + \frac{\ln(2)}{4} = 0.464848\ldots < \ln 2$$

where G is Catalan's constant, for the half-filled case $Q = 1/2$.

GPS: A. Georges, O. Parcollet, and S. Sachdev, PRB **63**, 134406 (2001)

W. Fu and S. Sachdev, PRB **94**, 035135 (2016)

The SYK model



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No quasiparticles !

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The SYK model

- Low energy, many-body density of states
  $\rho(E) \sim e^{N s_0} \sinh(\sqrt{2(E - E_0)N\gamma})$

A. Georges, O. Parcollet, and S. Sachdev, PRB **63**, 134406 (2001)
D. Stanford and E. Witten, 1703.04612
A. M. Garica-Garcia, J.J.M. Verbaarschot, 1701.06593
D. Bagrets, A. Altland, and A. Kamenev, 1607.00694

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A. Kitaev, unpublished
J. Maldacena and D. Stanford, 1604.07818

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  $G \sim (T/\sin(\pi k_B T \tau/\hbar))^{1/2}$

A. Georges and O. Parcollet PRB **59**, 5341 (1999)

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- The last property indicates $\tau_{\rm eq} \sim \hbar/(k_B T)$, and this has been found in a recent numerical study.

A. Eberlein, V. Kasper, S. Sachdev, and J. Steinberg, arXiv:1706.07803

Quantum matter without quasiparticles:

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S. Sachdev,
Quantum Phase Transitions,
Cambridge (1999)

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$$\tau_{\rm eq} = \# \frac{\hbar}{k_B T}$$

- Systems without quasiparticles are the fastest possible in reaching local equilibrium, and all many-body quantum systems obey, as $T \to 0$

$$\tau_{\rm eq} > C \frac{\hbar}{k_B T} \, .$$

S. Sachdev,
Quantum Phase Transitions,
Cambridge (1999)

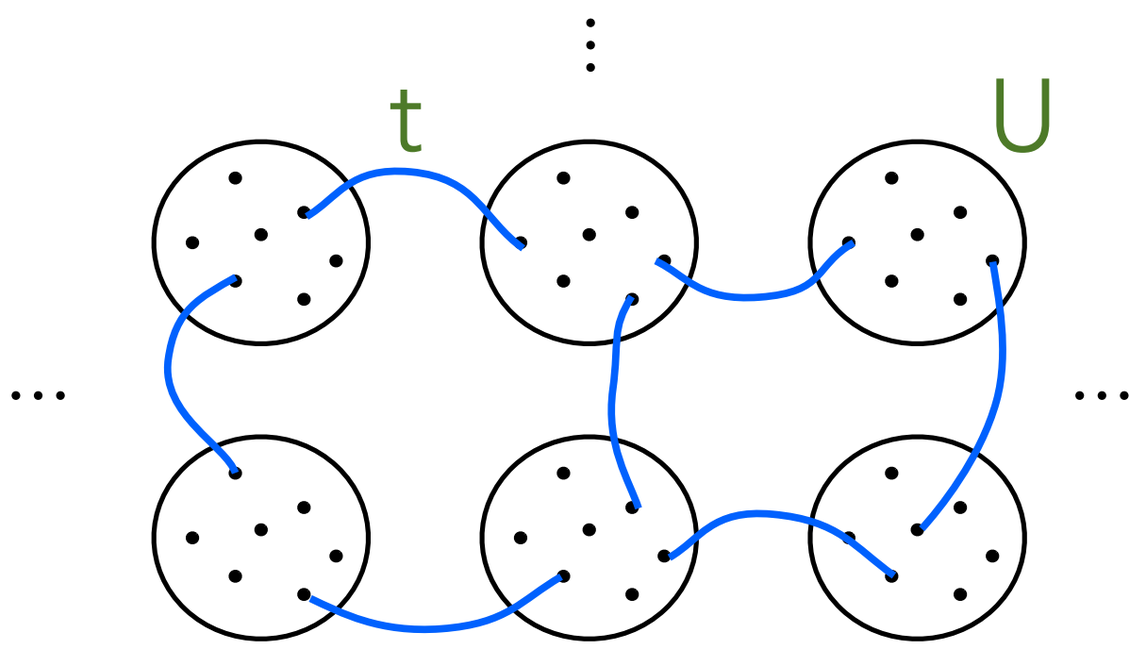
  - In Fermi liquids $\tau_{\rm eq} \sim 1/T^2$, and so the bound is obeyed as $T \to 0$.
  - This bound rules out quantum systems with *e.g.* $\tau_{\rm eq} \sim \hbar/(J k_B T)^{1/2}$.
  - There is no bound in classical mechanics ($\hbar \to 0$). By cranking up frequencies, we can attain equilibrium as quickly as we desire.

See also A. Georges and O. Parcollet PRB **59**, 5341 (1999)

arXiv:1705.00117

Title: A strongly correlated metal built from Sachdev-Ye-Kitaev models

Authors: Xue-Yang Song, Chao-Ming Jian, Leon Balents



$$H = \sum_x \sum_{i<j,k<l} U_{ijkl,x} c^\dagger_{ix} c^\dagger_{jx} c_{kx} c_{lx} + \sum_{\langle xx' \rangle} \sum_{i,j} t_{ij,xx'} c^\dagger_{i,x} c_{j,x'}$$

$$\overline{|U_{ijkl}|^2} = \frac{2U^2}{N^3} \qquad \overline{|t_{ij,x,x'}|^2} = t_0^2/N$$

See also A. Georges and O. Parcollet PRB **59**, 5341 (1999)

arXiv:1705.00117

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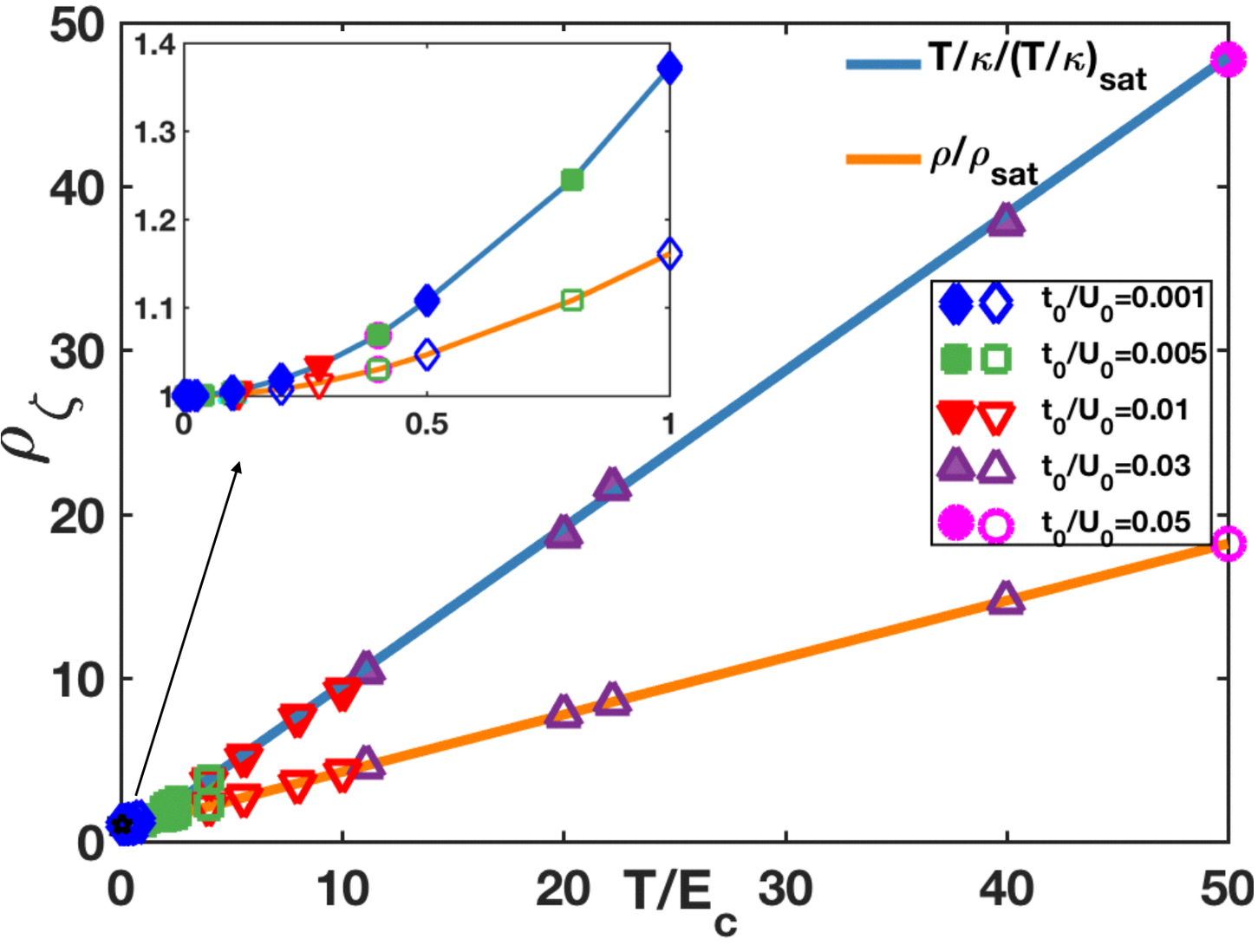
Low ‘coherence’ scale

$$E_c \sim \frac{t_0^2}{U}$$



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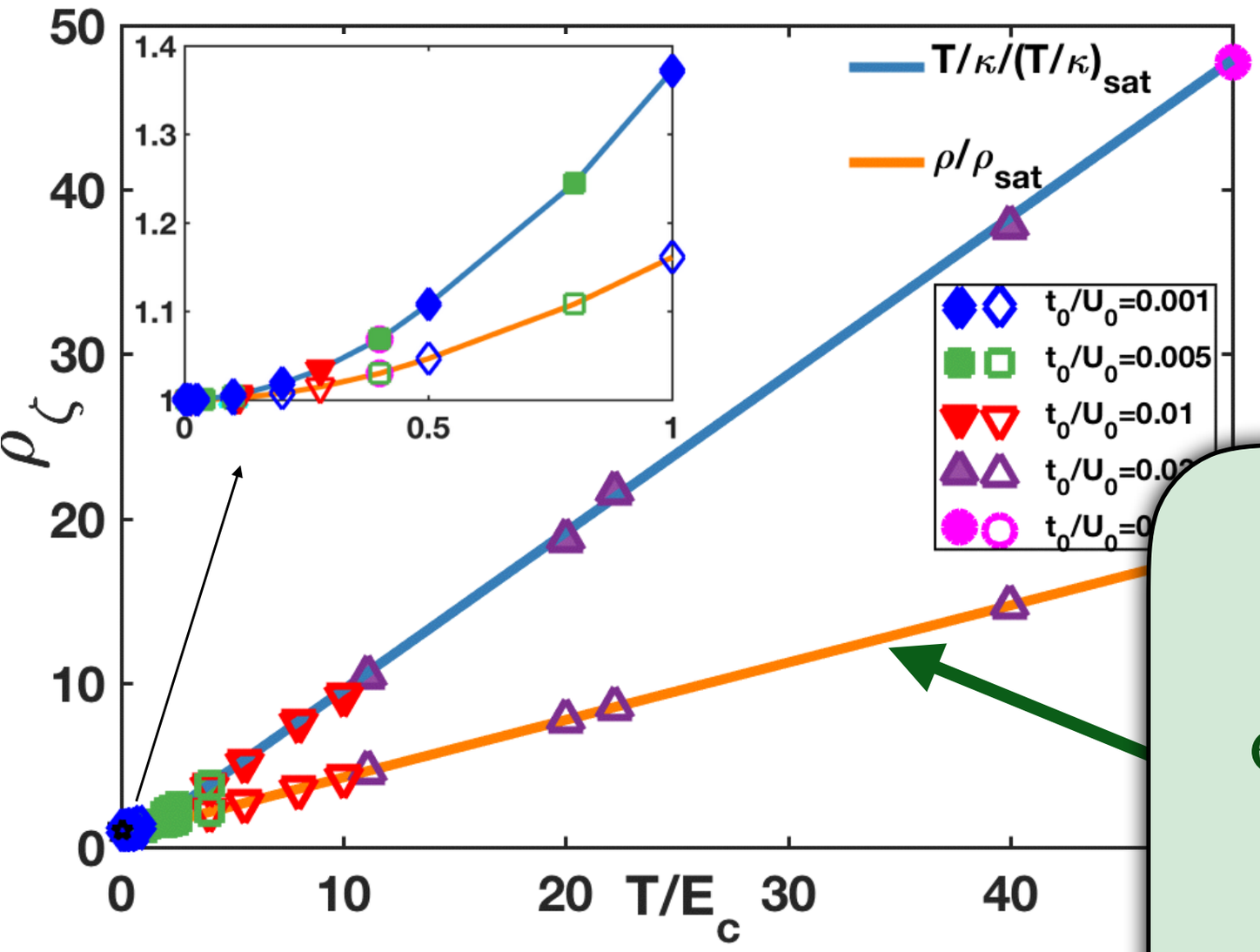

$$E_c \sim \frac{t_0^2}{U}$$

For $E_c < T < U$, the resistivity, ρ, and entropy density, s, are

$$\rho \sim \frac{h}{e^2}\left(\frac{T}{E_c}\right), \quad s = s_0$$

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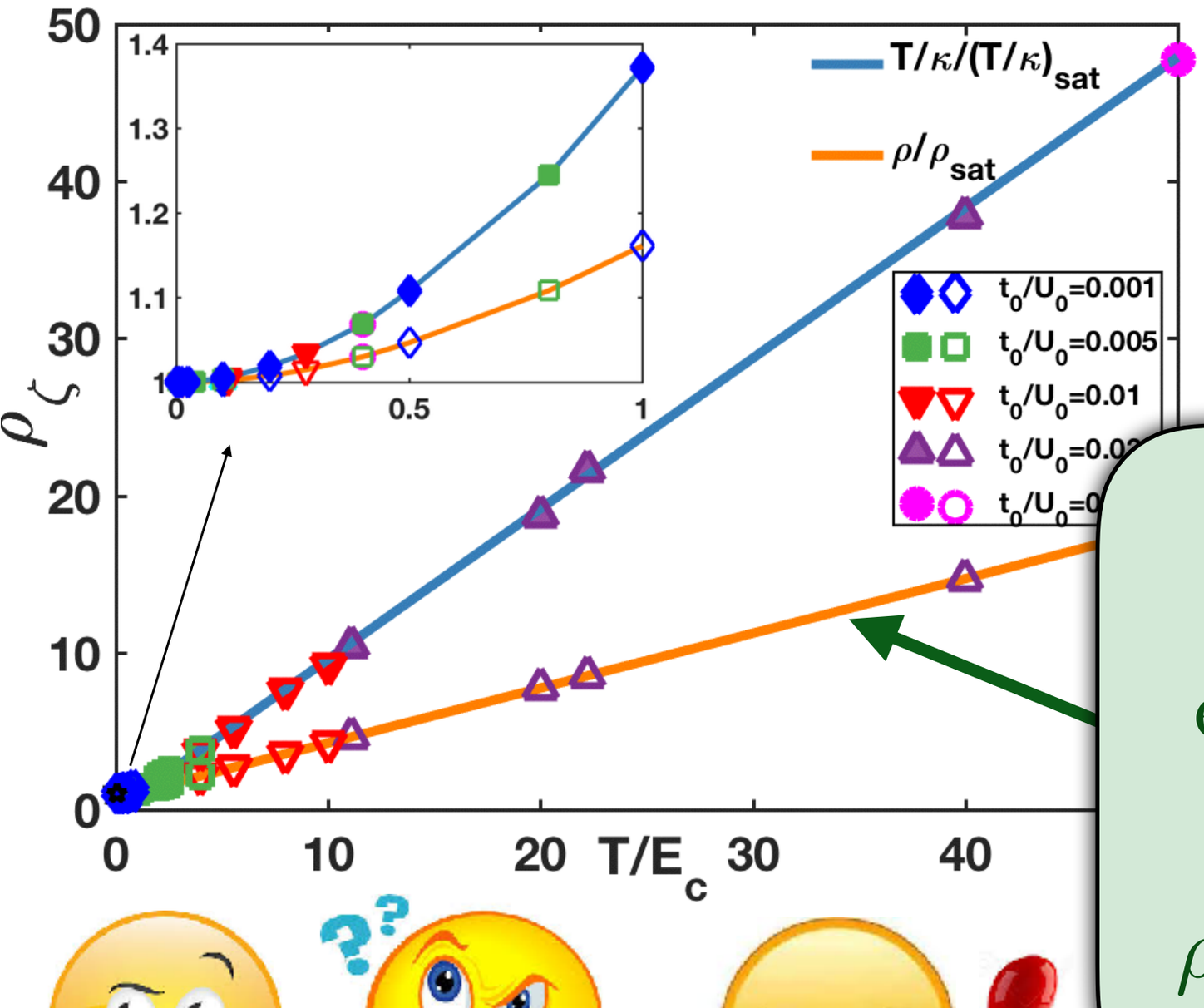

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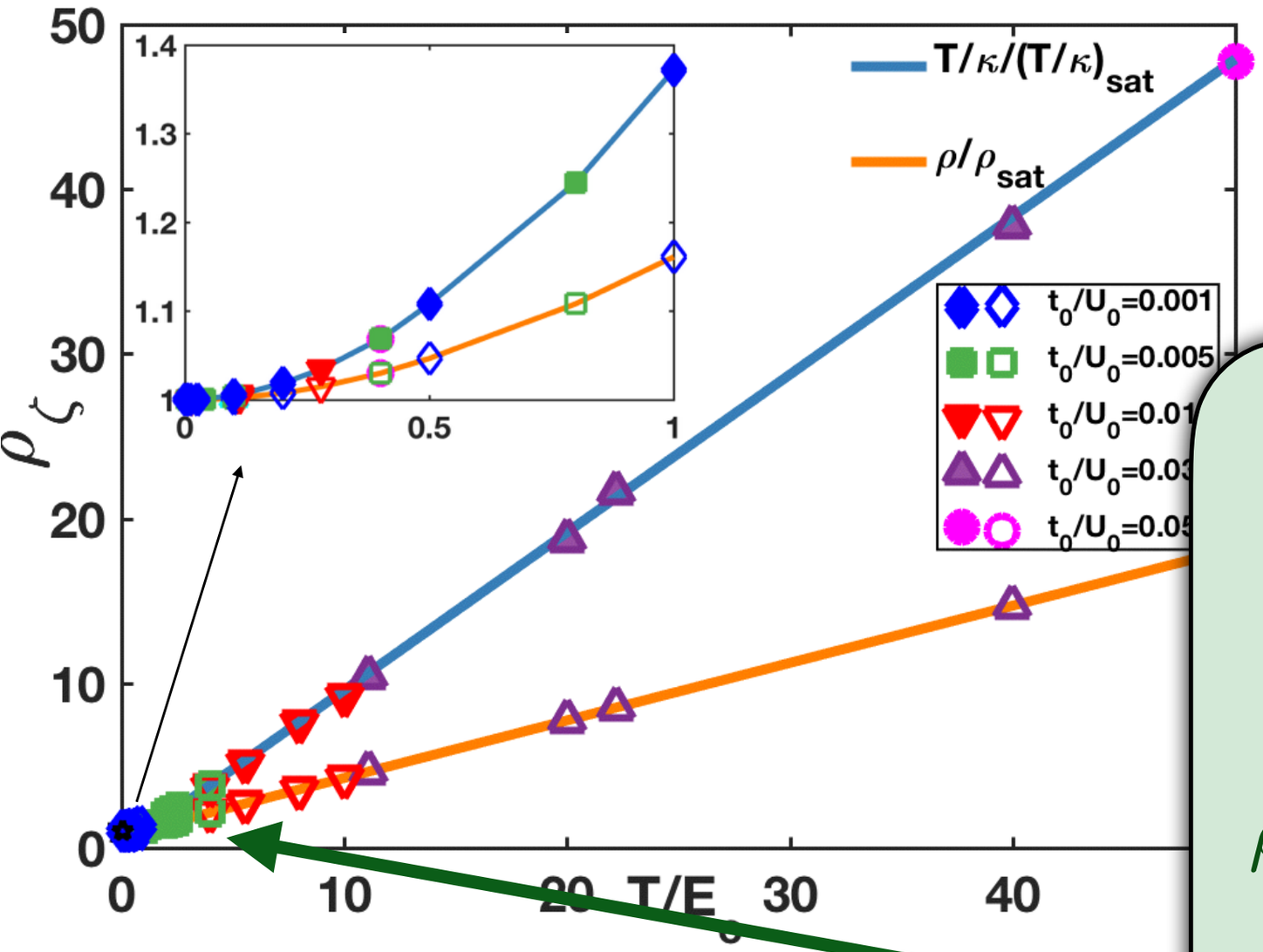


$$E_c \sim \frac{t_0^2}{U}$$

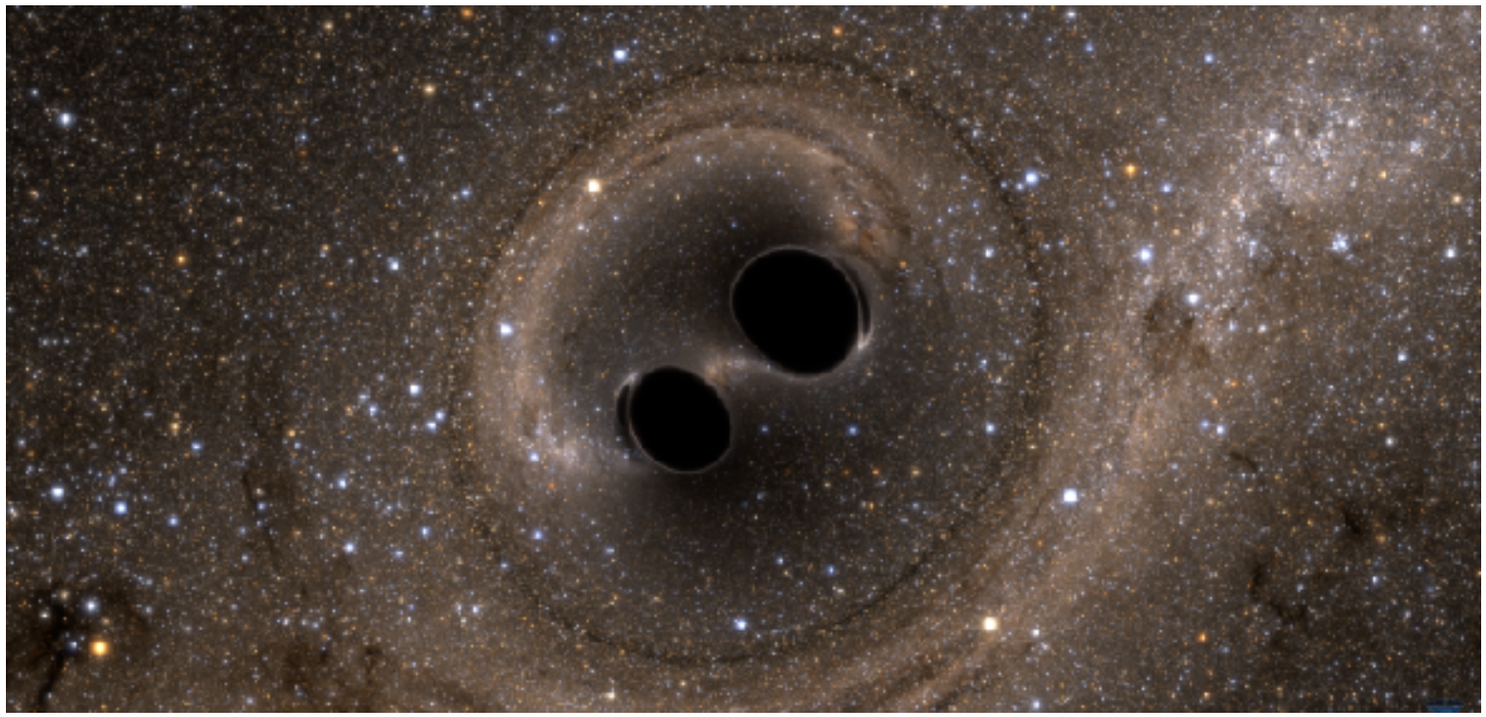
For $T < E_c$, the resistivity, ρ, and entropy density, s, are

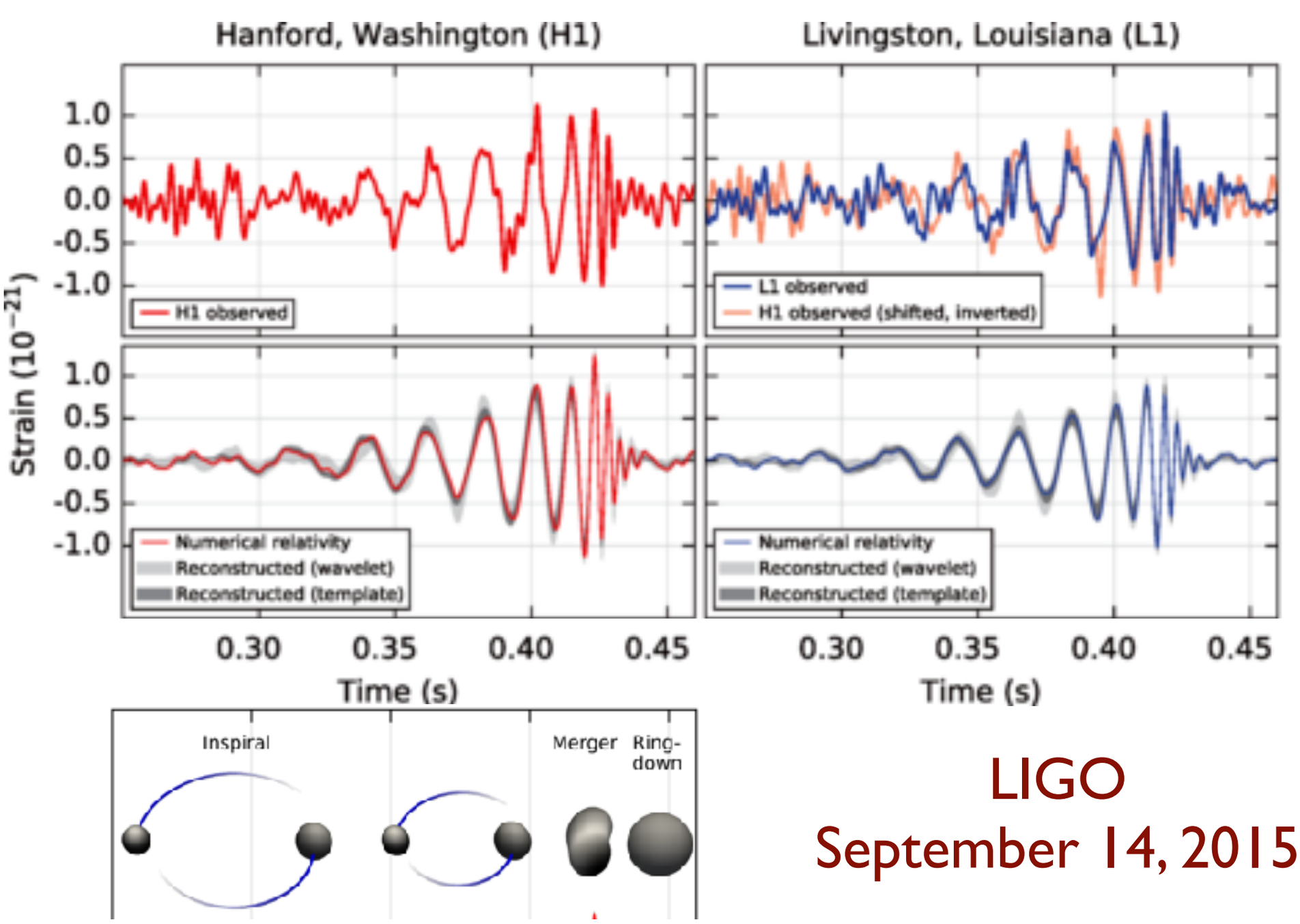
$$\rho = \frac{h}{e^2}\left[c_1 + c_2\left(\frac{T}{E_c}\right)^2\right]$$

$$s \sim s_0\left(\frac{T}{E_c}\right)$$

Black holes

- Black holes have an entropy and a temperature, T_H.
- The entropy is proportional to their surface area.





LIGO
September 14, 2015

- The Hawking temperature, T_H influences the radiation from the black hole at the very last stages of the ring-down (not observed so far). The ring-down (approach to thermal equilibrium) happens very rapidly in a time $\sim \dfrac{\hbar}{k_B T_H} = \dfrac{8\pi G M}{c^3} \sim 8$ milliseconds.

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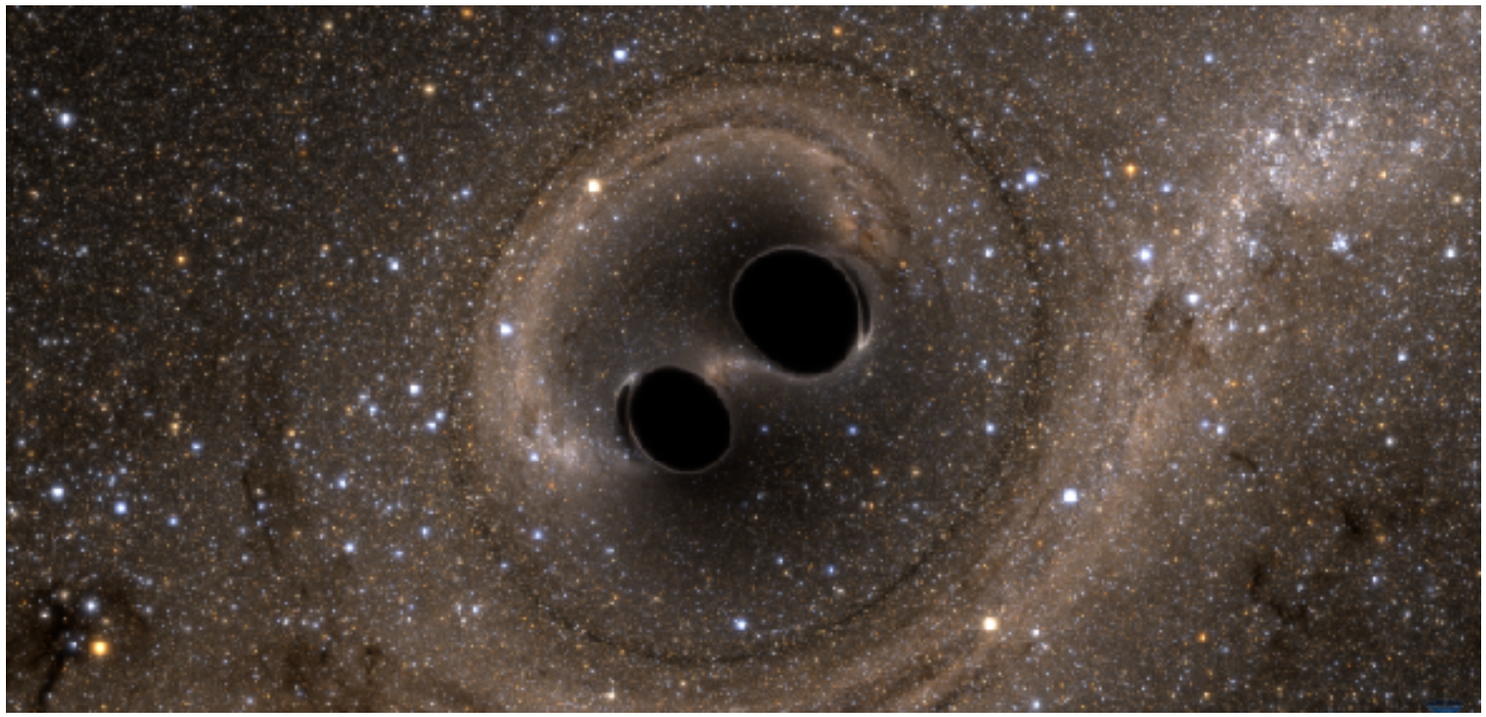


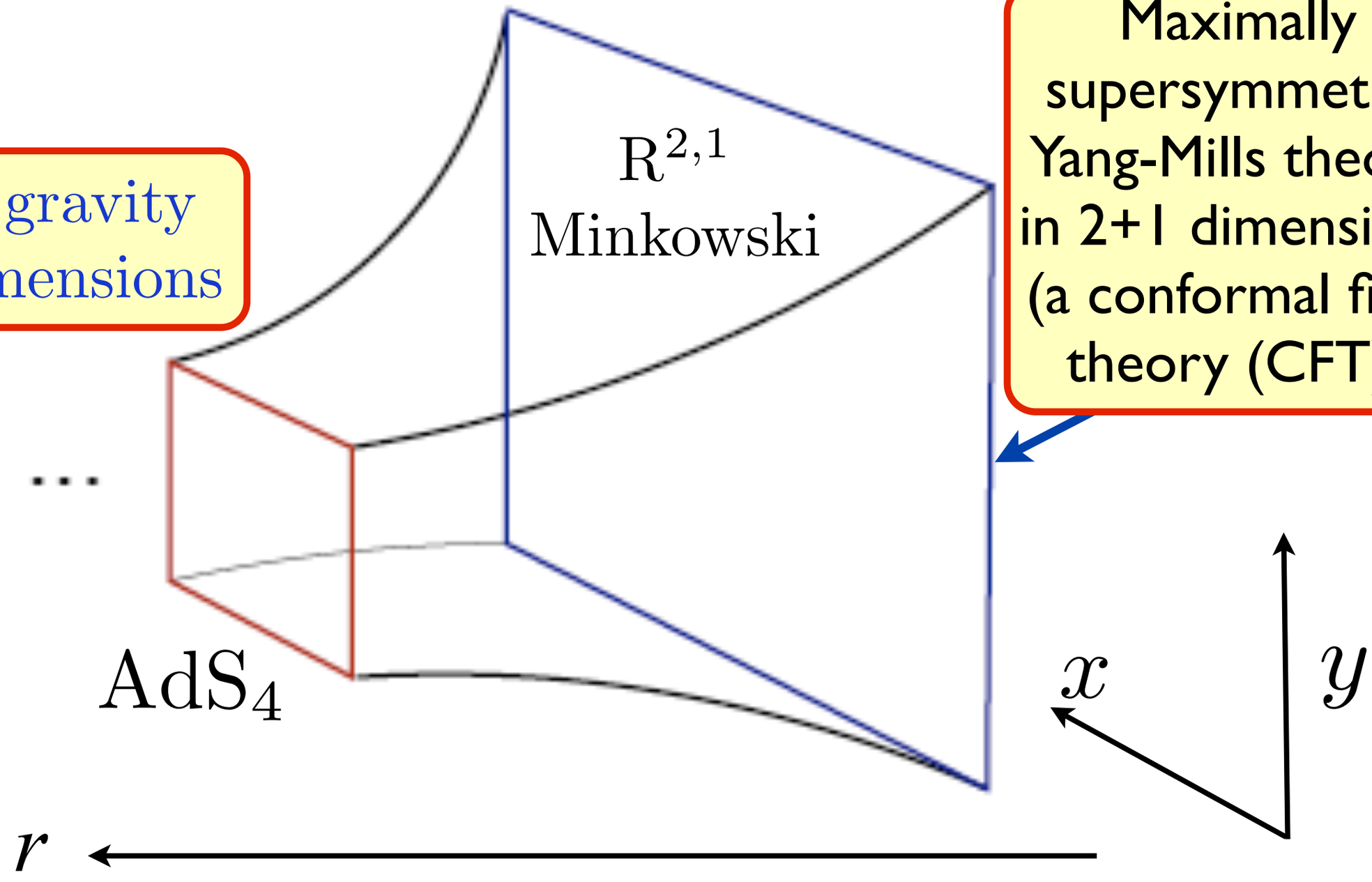
Black holes

- Black holes have an entropy and a temperature, T_H.
- The entropy is proportional to their surface area.
- They relax to thermal equilibrium in a time $\sim \hbar/(k_B T_H)$.

AdS/CFT correspondence at zero temperature

Quantum gravity in 3+1 dimensions

Maximally supersymmetric Yang-Mills theory in 2+1 dimensions (a conformal field theory (CFT))



$$ds^2 = \left(\frac{L}{r}\right)^2 \left[dr^2 - dt^2 + dx^2 + dy^2\right]$$

Maldacena, Gubser, Klebanov, Polyakov, Witten

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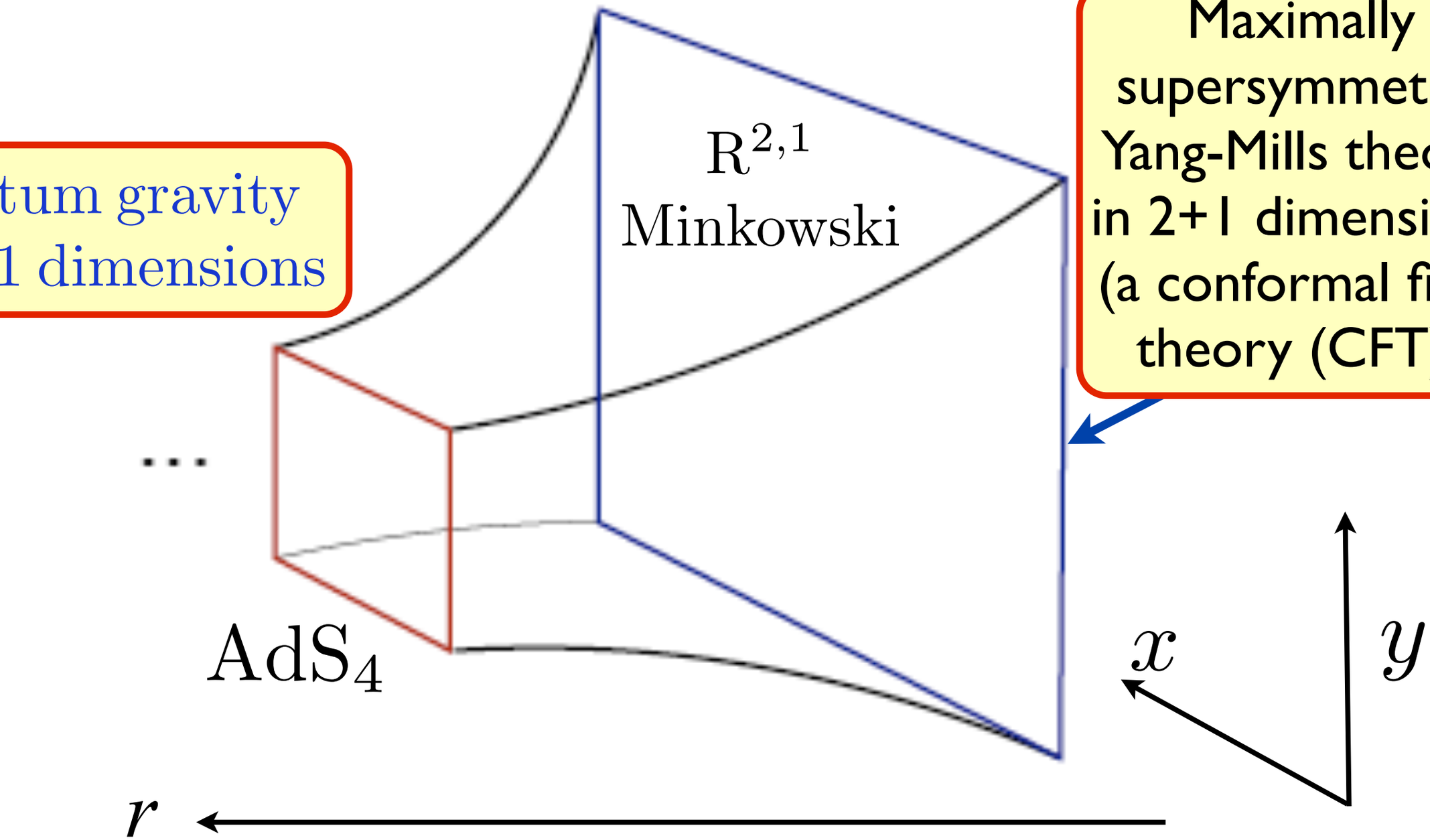


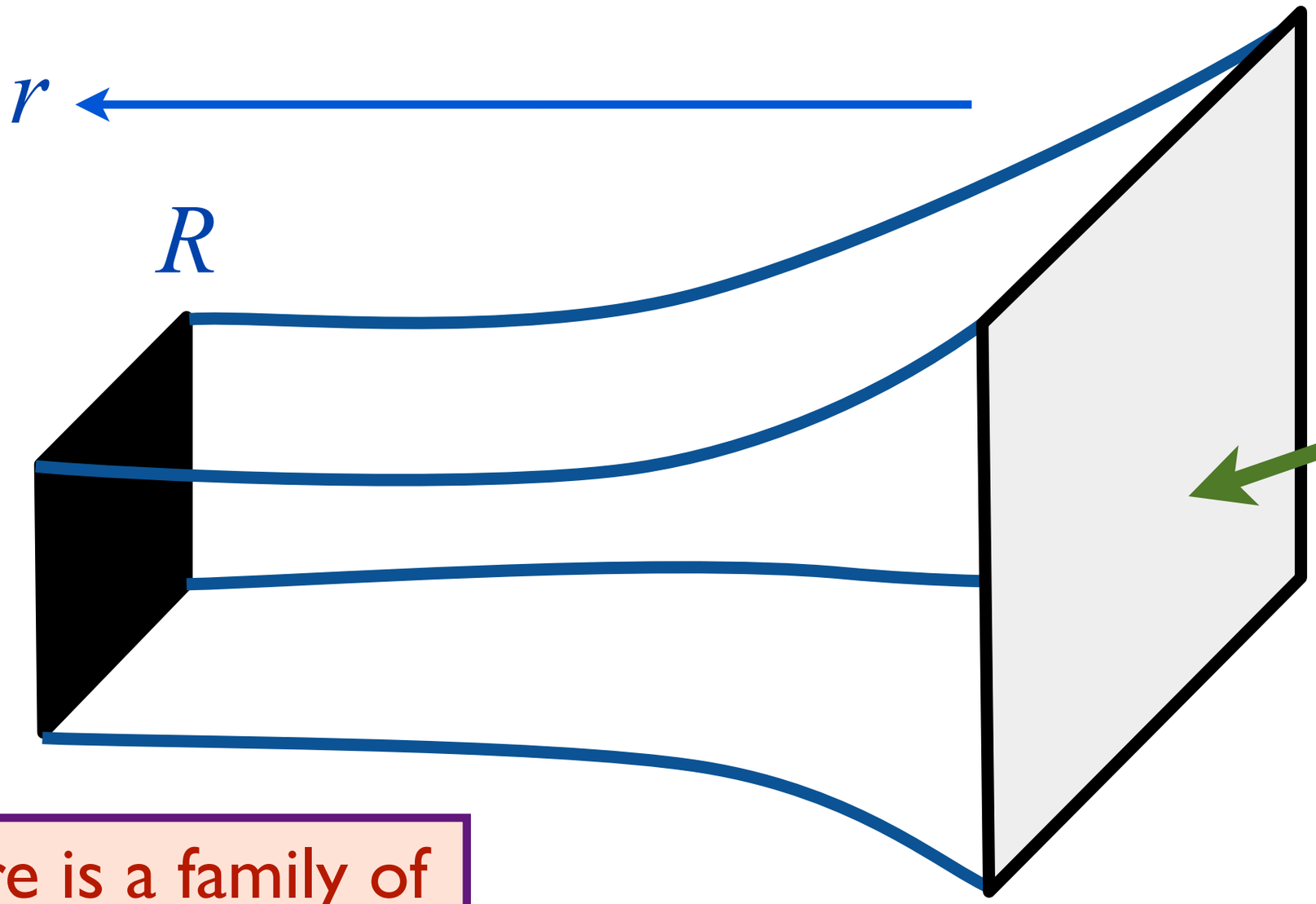
This spacetime is a solution of Einstein gravity with a negative cosmological constant

$$\mathcal{S}_E = \int d^4x\sqrt{-g}\left[\frac{1}{2\kappa^2}\left(R + \frac{6}{L^2}\right)\right]$$

Maldacena; Gubser; Klebanov; Polyakov, Witten

AdS/CFT correspondence at non-zero temperatures

AdS_4-Schwarzschild black-brane



Maximally supersymmetric Yang-Mills at a temperature

$$k_B T = \frac{3\hbar}{4\pi R}.$$

There is a family of solutions of Einstein gravity which describe non-zero temperatures

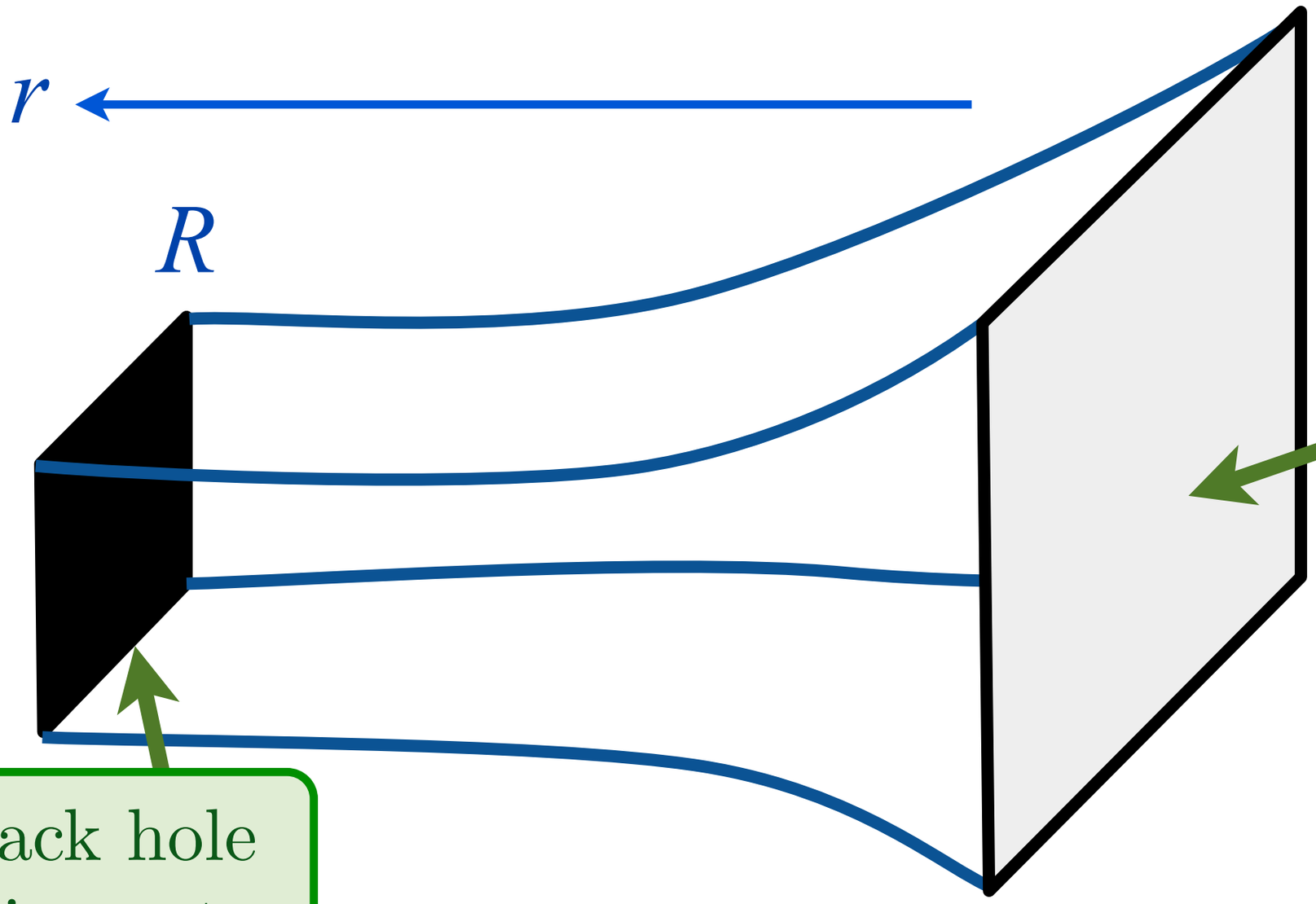
$$ds^2 = \left(\frac{L}{r}\right)^2 \left[\frac{dr^2}{f(r)} - f(r)dt^2 + dx^2 + dy^2\right]$$

with $f(r) = 1 - (r/R)^3$

Maldacena, Gubser, Klebanov, Polyakov, Witten

AdS/CFT correspondence at non-zero temperatures

AdS_4-Schwarzschild black-brane



Maximally supersymmetric Yang-Mills at a temperature

$$k_B T = \frac{3\hbar}{4\pi R}.$$

Black hole horizon at a Hawking temperature

$$T_H = T$$

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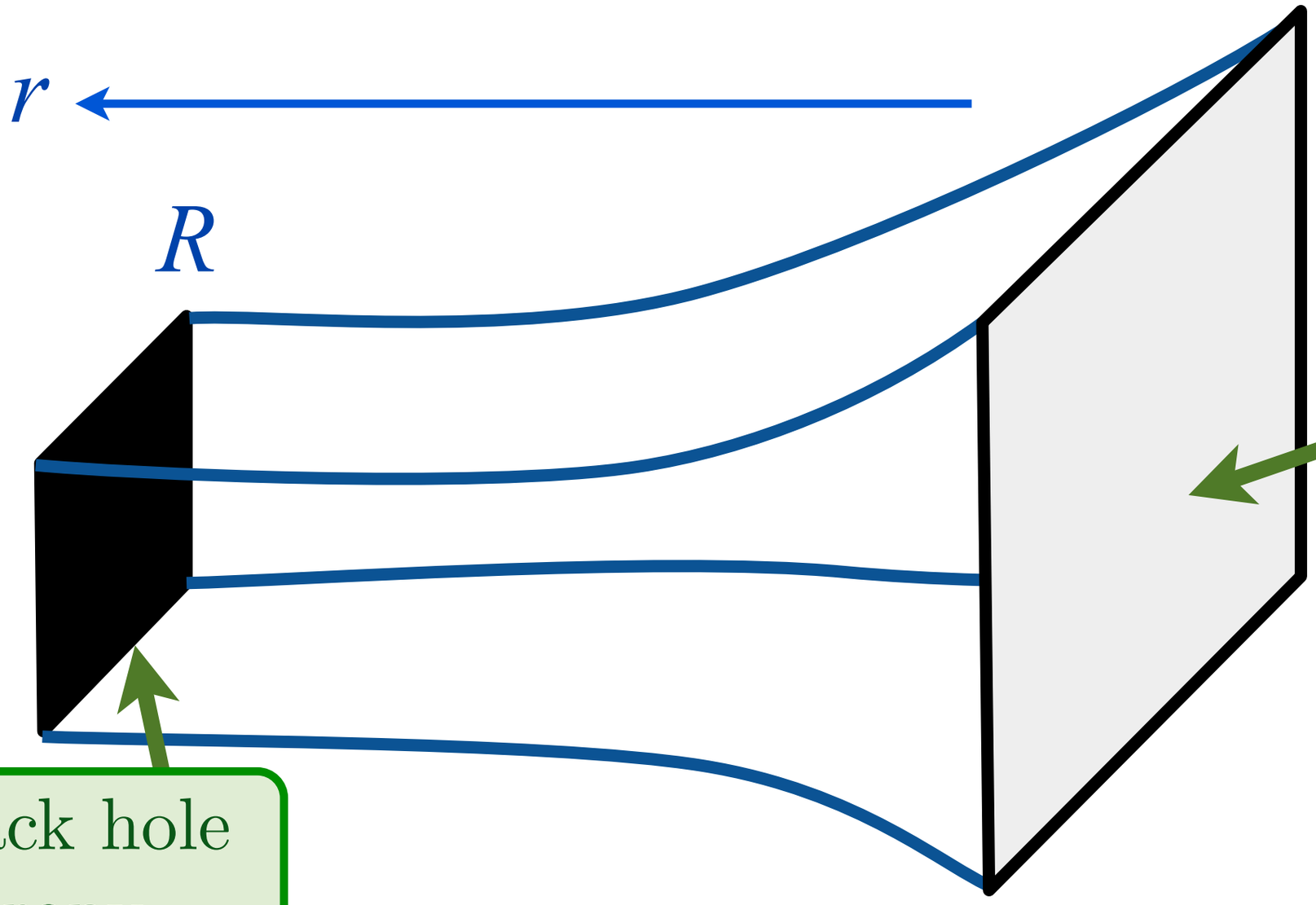

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Black hole entropy = Entropy of Yang-Mills theory

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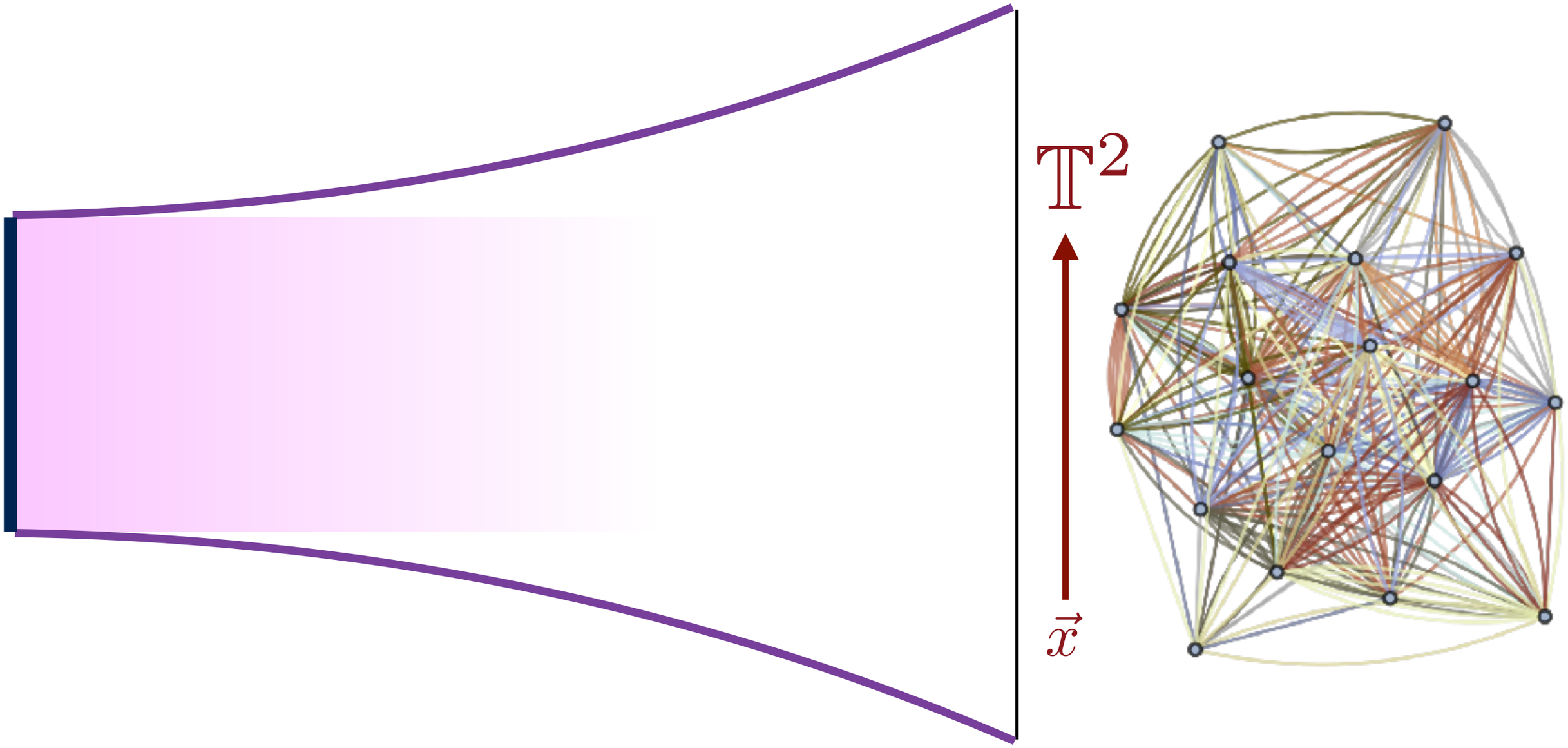
SYK and black holes



Is there a holographic quantum gravity dual of the SYK model ?

The SYK model

- Low energy, many-body density of states
  $\rho(E) \sim e^{N s_0} \sinh(\sqrt{2(E-E_0)N\gamma})$

- Low temperature entropy $S = N s_0 + N\gamma T + \ldots$.

- $T = 0$ fermion Green's function $G(\tau) \sim \tau^{-1/2}$ at large τ. (Fermi liquids with quasiparticles have $G(\tau) \sim 1/\tau$)

- $T > 0$ Green's function has conformal invariance
  $G \sim (T/\sin(\pi k_B T \tau/\hbar))^{1/2}$

- The last property indicates $\tau_{\rm eq} \sim \hbar/(k_B T)$, and this has been found in a recent numerical study.

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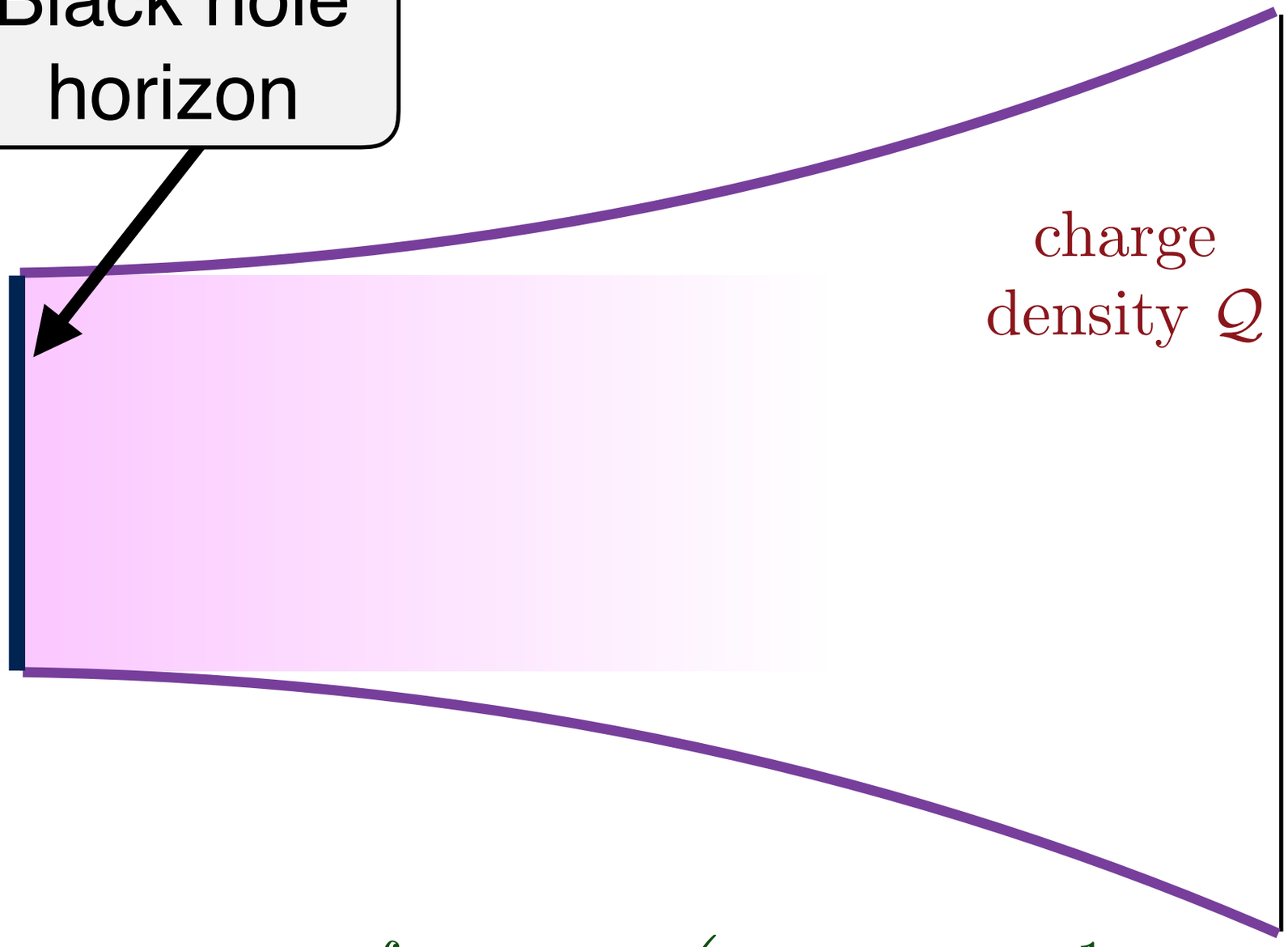
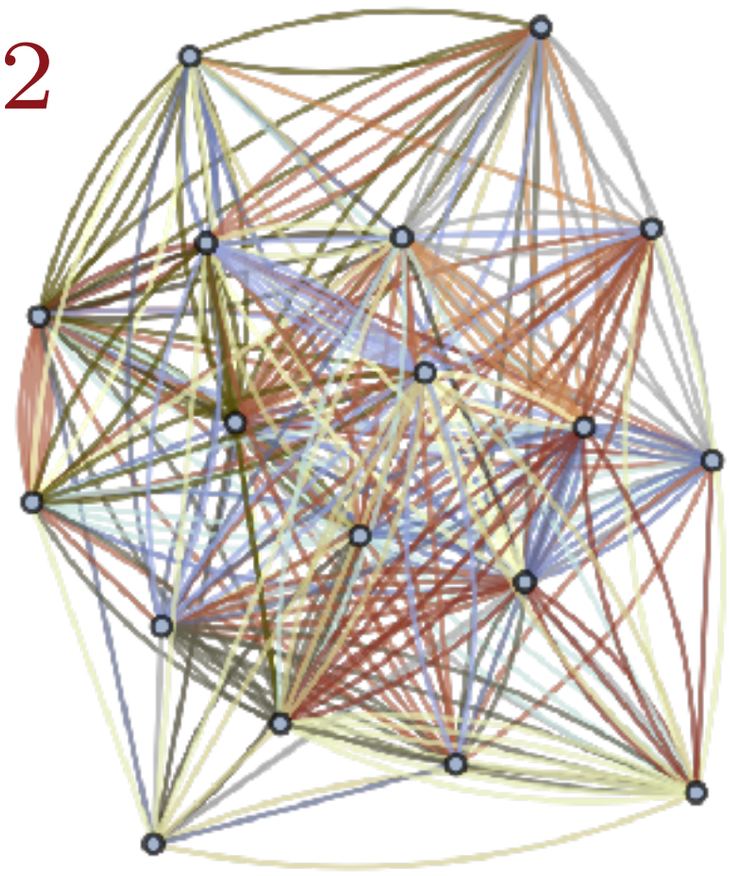
- Th
ha

Is there any black hole which holographically matches these properties ?

SYK and black holes

Black hole horizon





$$S = \int d^4x \sqrt{-\hat{g}} \left(\hat{\mathcal{R}} + 6/L^2 - \frac{1}{4} \hat{F}_{\mu\nu} \hat{F}^{\mu\nu} \right)$$

Yes, a charged black hole in Einstein-Maxwell theory

S. Sachdev, PRL **105**, 151602 (2010)

SYK and black holes

Black hole horizon

$$\mathrm{AdS}_2 \times \mathbb{T}^2$$
$$ds^2 = (d\zeta^2 - dt^2)/\zeta^2 + d\vec{x}^2$$
Gauge field: $A = (\mathcal{E}/\zeta)dt$

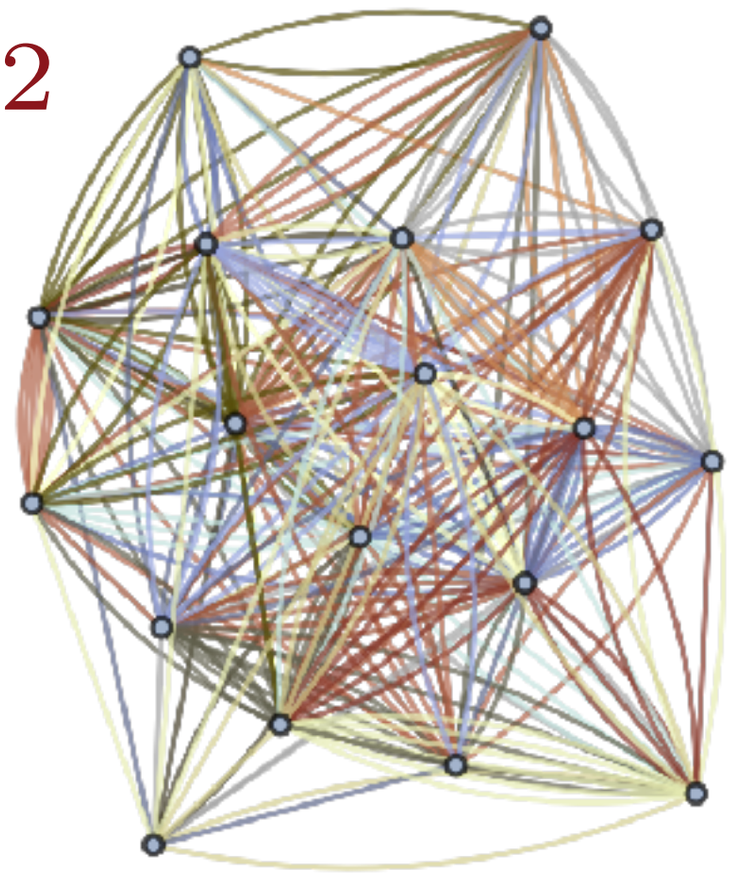
charge density $\mathcal{Q}$

$\zeta = \infty$

ζ

$\mathbb{T}^2$

$\vec{x}$



The leading low temperature properties of an AdS-Reissner-Nordstrom black hole (as computed by T. Faulkner, H. Liu, J. McGreevy, and D. Vegh, arXiv:0907.2694) match those of the SYK model.
The mapping applies when temperature $\ll 1/(\text{size of } \mathbb{T}^2)$.

S. Sachdev, PRL **105**, 151602 (2010)

SYK and black holes

Black hole horizon

charge density $\mathcal{Q}$

$\mathrm{AdS}_2 \times \mathbb{T}^2$

$$\boxed{ds^2 = (d\zeta^2 - dt^2)/\zeta^2} + d\vec{x}^2$$

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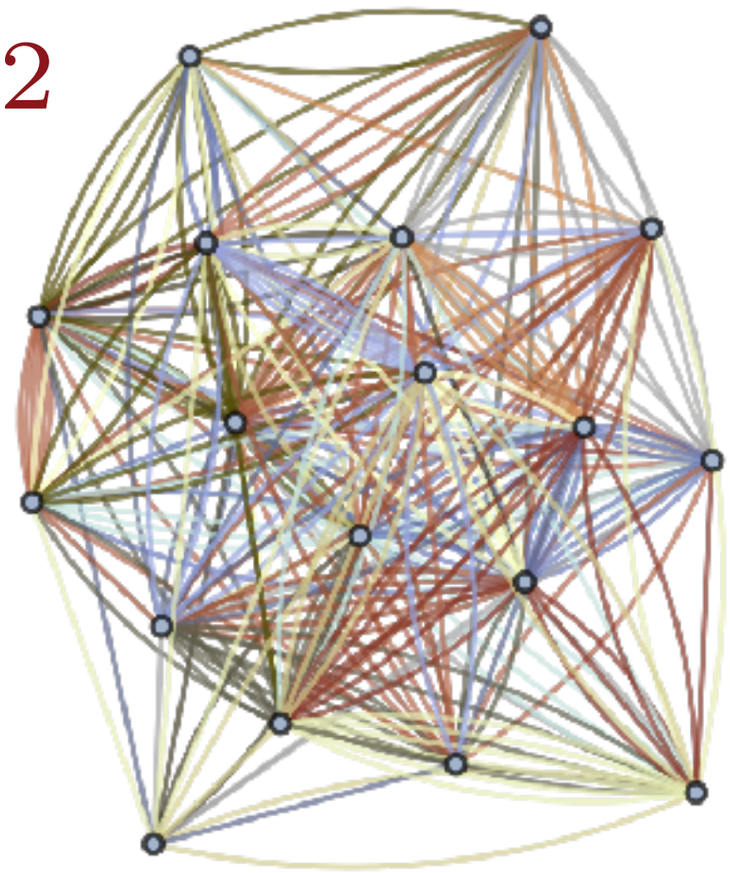

Quantum gravity on the 1+1 dimensional spacetime AdS_2 (when embedded in AdS_4) is holographically matched to the 0+1 dimensional SYK model

S. Sachdev, PRL **105**, 151602 (2010); A. Kitaev (unpublished); J. Maldacena, D. Stanford, and Zhenbin Yang, arXiv:1606.01857

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- $T > 0$ Green's function has conformal invariance

- [illegible]

- [illegible]

All these properties of the SYK model match those of the AdS_2 horizon in Einstein-Maxwell theory

S. Sachdev, PRL **105**, 151602 (2010)

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-

Schwarzian theory of quantum gravity fluctuations also matches these corrections

A Kitaev, unpublished, J. Maldacena, D. Stanford, and Zhenbin Yang, arXiv:1606.01857;
D. Stanford and E. Witten, arXiv:1703.04612

Many-body quantum chaos

- Using holographic analogies, Shenker and Stanford introduced the "Lyapunov time", τ_L, the time over which a generic many-body quantum system loses memory of its initial state.

S. Shenker and D. Stanford, arXiv:1306.0622

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S. Shenker and D. Stanford, arXiv:1306.0622

- A shortest-possible time to reach quantum chaos was established

$$\tau_L \geq \frac{\hbar}{2\pi k_B T}$$

J. Maldacena, S. H. Shenker and D. Stanford, arXiv:1503.01409

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J. Maldacena, S. H. Shenker and D. Stanford, arXiv:1503.01409

- The SYK model, and black holes in Einstein gravity, saturate the bound on the Lyapunov time

$$\tau_L = \frac{\hbar}{2\pi k_B T}$$

A. Kitaev, unpublished
J. Maldacena and D. Stanford, arXiv:1604.07818

Quantum matter <u>without</u> quasiparticles:

- No quasiparticle decomposition of low-lying states:
  $E \neq \sum_\alpha n_\alpha \varepsilon_\alpha$
  $+ \sum_{\alpha,\beta} F_{\alpha\beta} n_\alpha n_\beta + \ldots$

- Thermalization and many-body chaos in the shortest possible time of order $\hbar/(k_B T)$.

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- These are also characteristics of black holes in quantum gravity.