

The SYK model

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PHYSICS



HARVARD

What are quasiparticles ?

- **Quasiparticles are additive excitations:**

The low-lying excitations of the many-body system can be identified as a set $\{n_\alpha\}$ of quasiparticles with energy ε_α

$$E = \sum_\alpha n_\alpha \varepsilon_\alpha + \sum_{\alpha,\beta} F_{\alpha\beta} n_\alpha n_\beta + \dots$$

In a lattice system of N sites, this parameterizes the energy of $\sim e^{\alpha N}$ states in terms of poly(N) numbers.

What are quasiparticles ?

- Quasiparticles eventually collide with each other. Such collisions eventually leads to thermal equilibration in a chaotic quantum state, but the equilibration takes a long time. In a Fermi liquid, this time diverges as

$$\tau_{\text{eq}} \sim \frac{\hbar E_F}{(k_B T)^2} \quad , \quad \text{as } T \rightarrow 0,$$

where E_F is the Fermi energy.

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where E_F is the Fermi energy.

- This time is much longer than the ‘Planckian time’ $\hbar/(k_B T)$, which we will find in systems without quasiparticle excitations.

$$\tau_{\text{eq}} \gg \frac{\hbar}{k_B T} \quad , \quad \text{as } T \rightarrow 0.$$

1. Random matrix quasiparticle model

$q=2$, complex SYK

2. Matter without quasiparticles

$q=4$, complex SYK

3. The Schwarzian theory

4. Connections to black holes
with AdS_2 horizons

1. Random matrix quasiparticle model

$q=2$, complex SYK

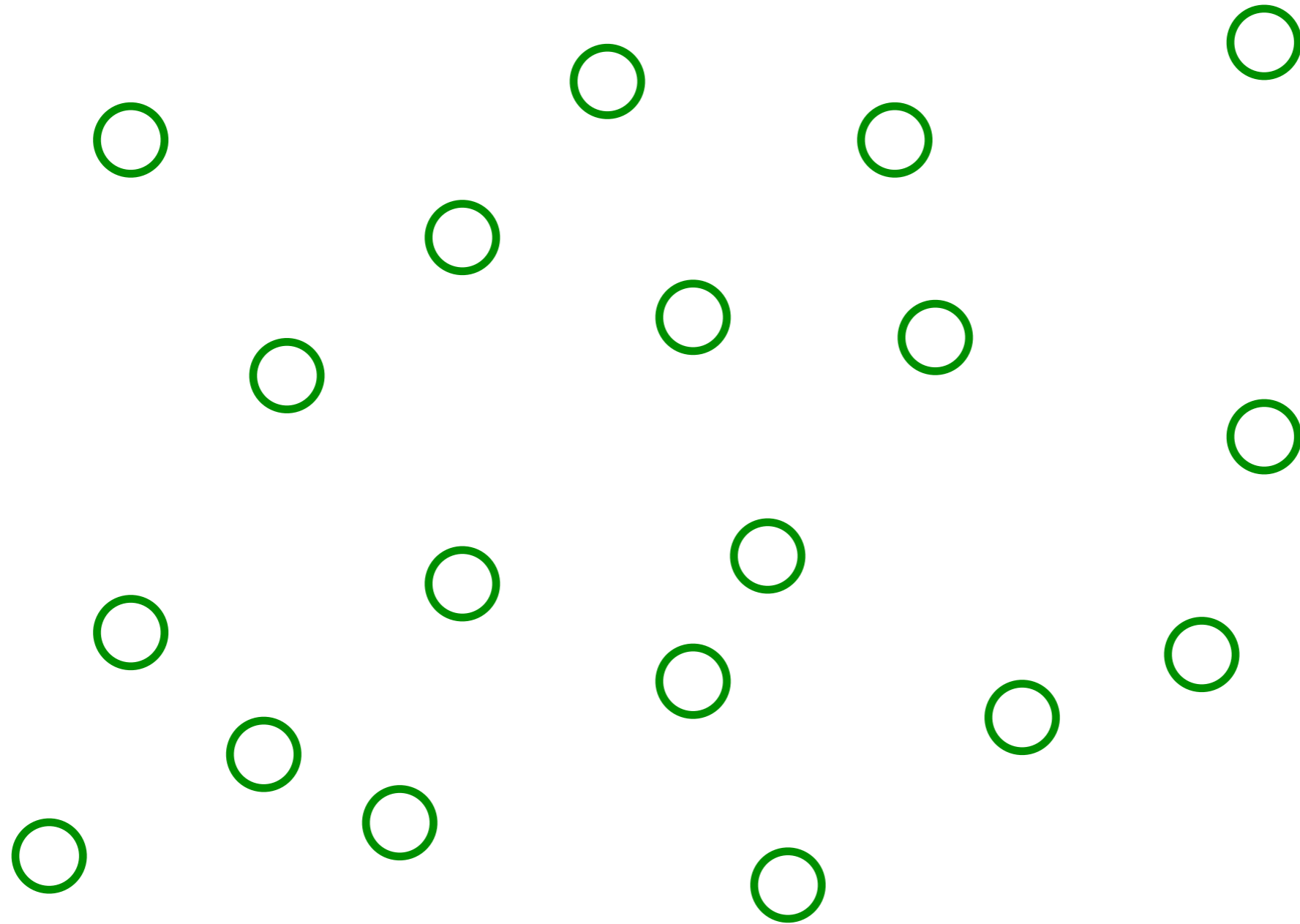
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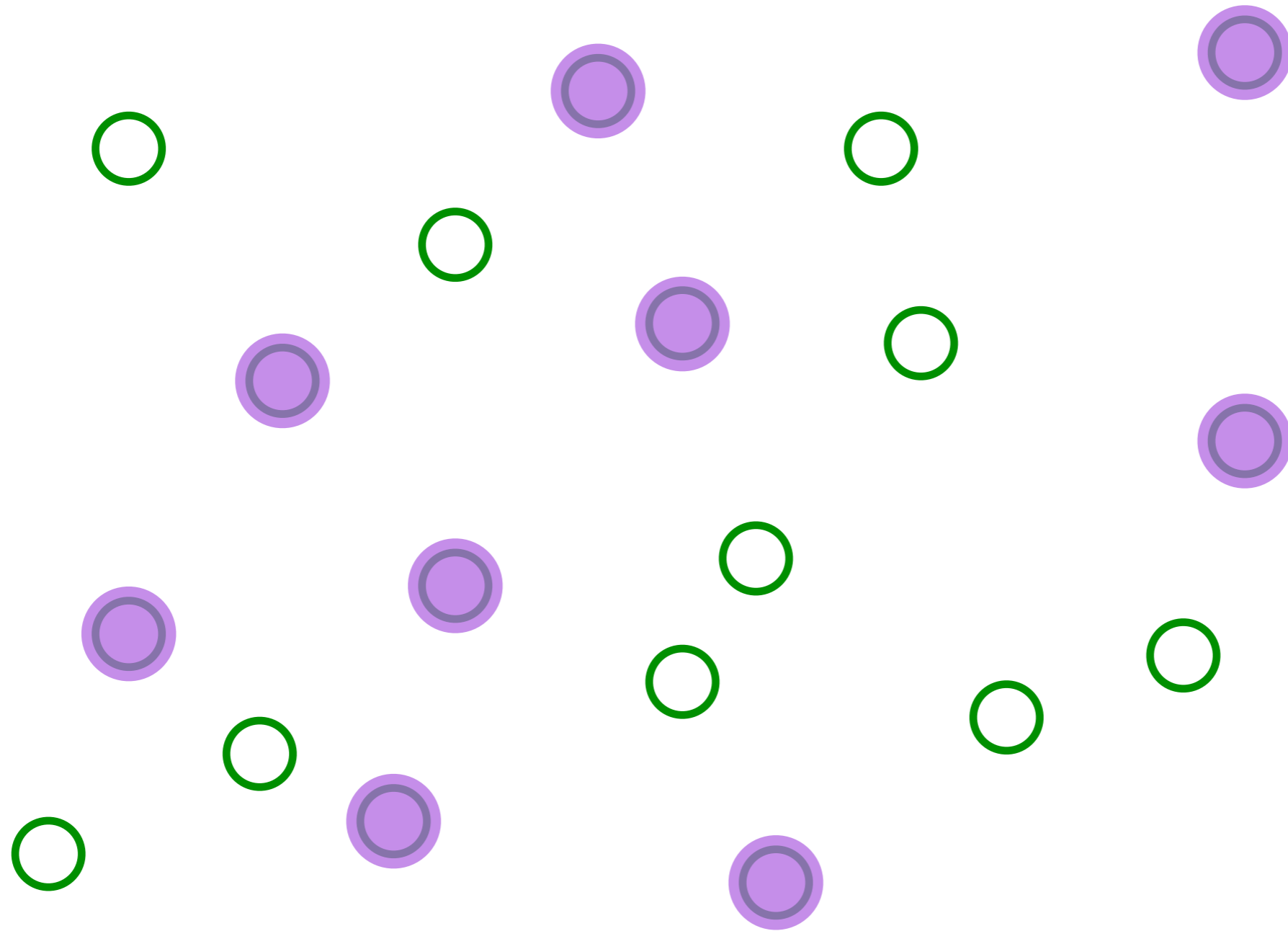
4. Connections to black holes
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A simple model of a metal with quasiparticles



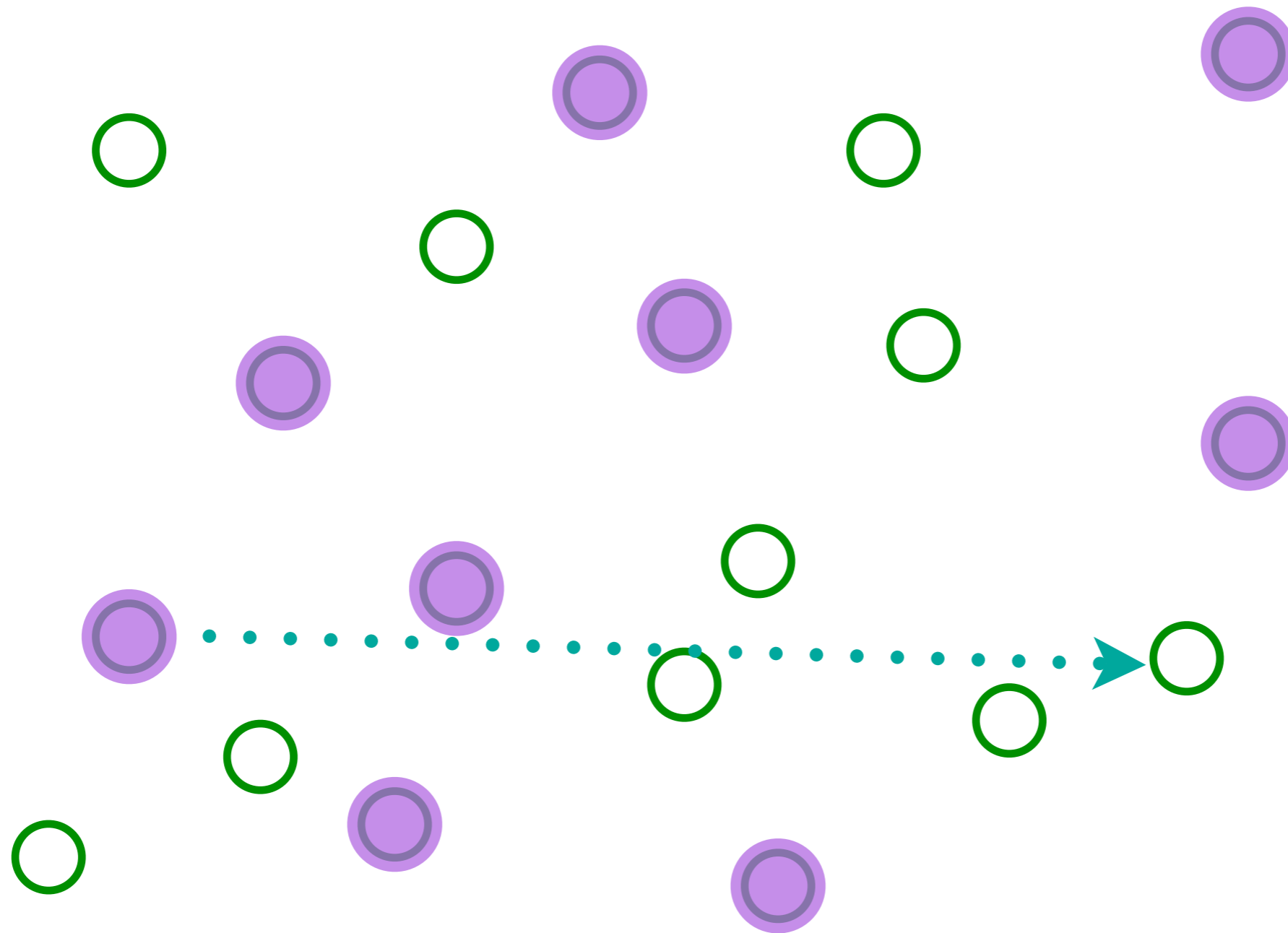
Pick a set of random positions

A simple model of a metal with quasiparticles



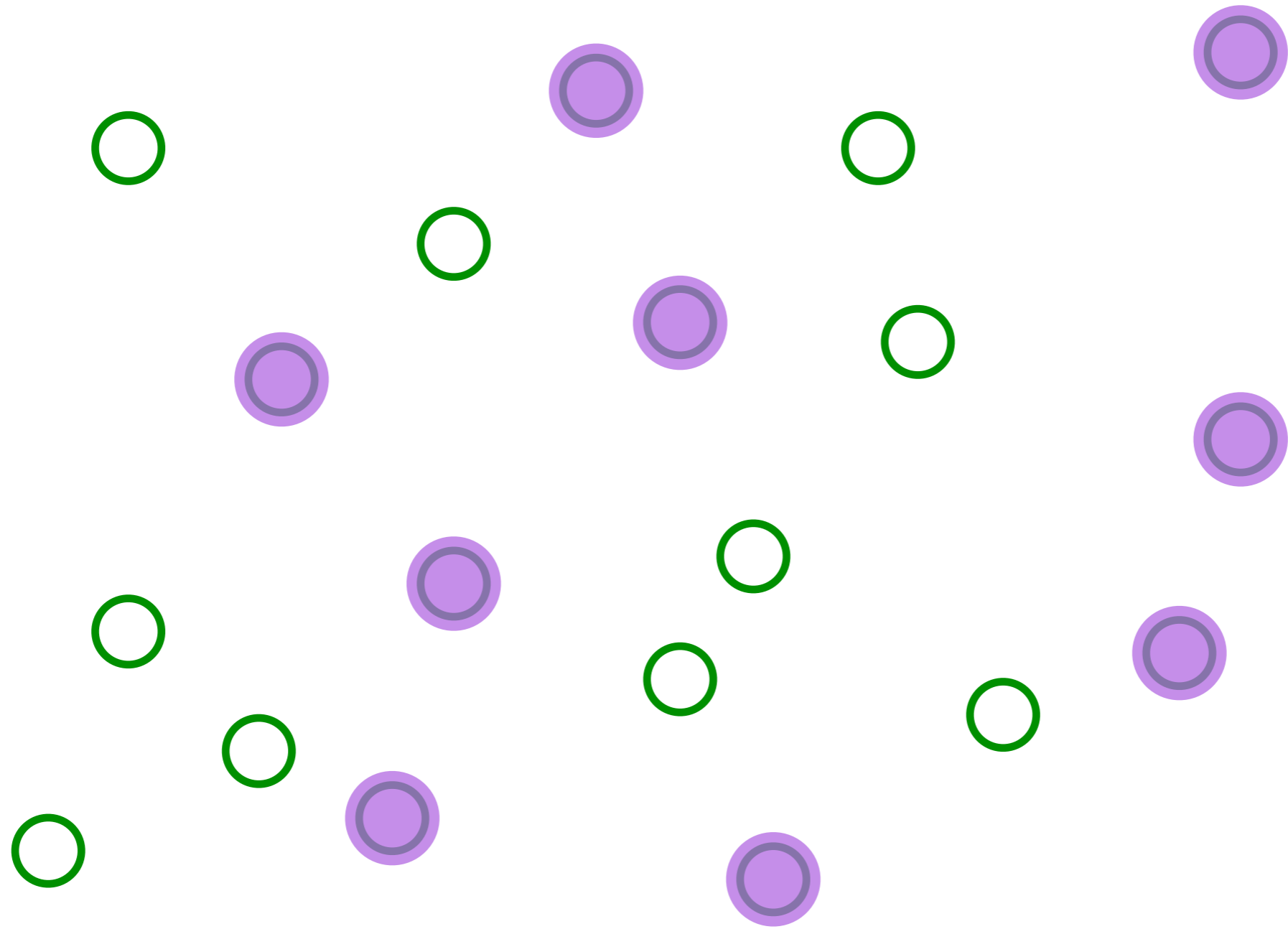
Place electrons randomly on some sites

A simple model of a metal with quasiparticles



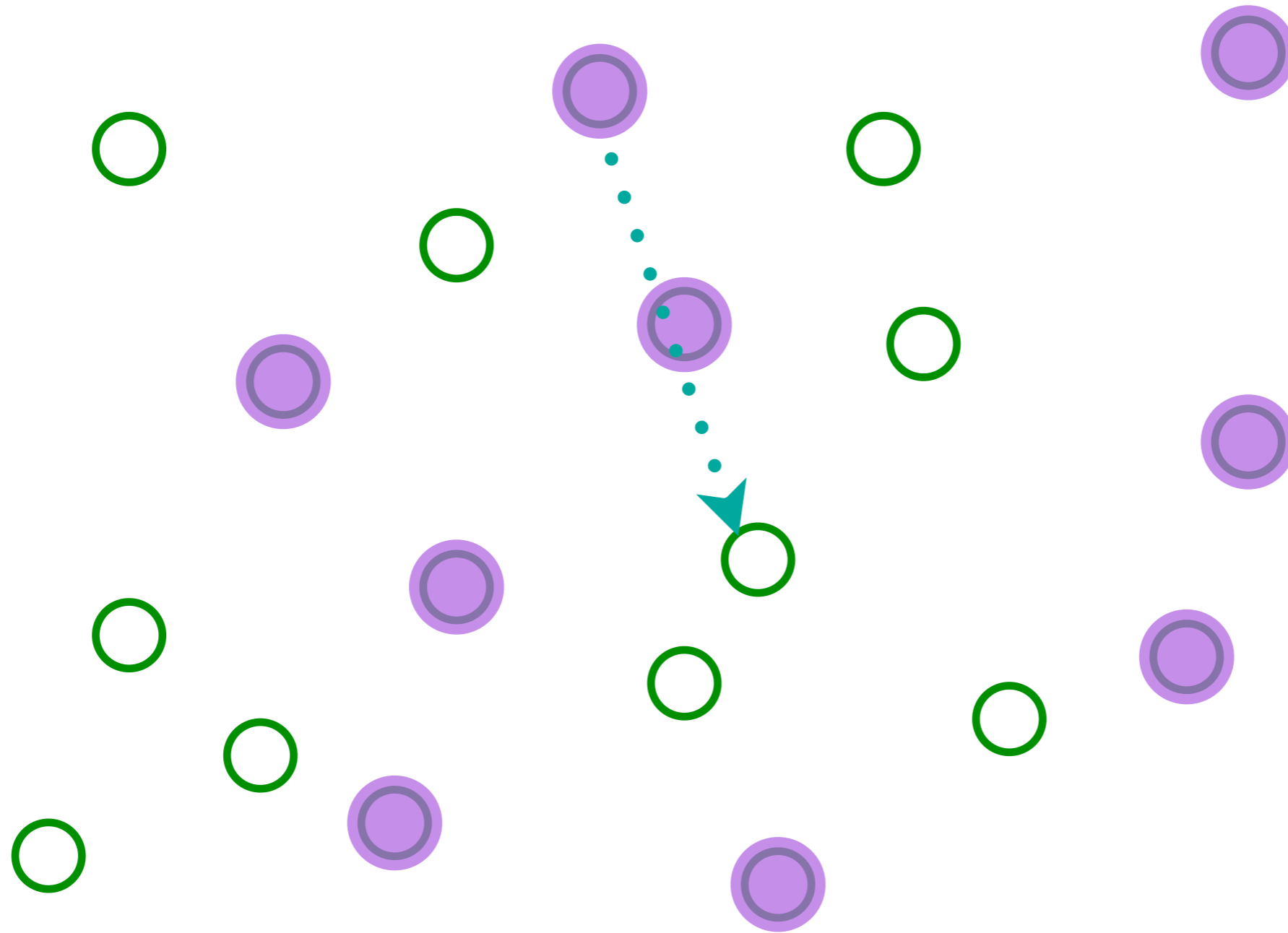
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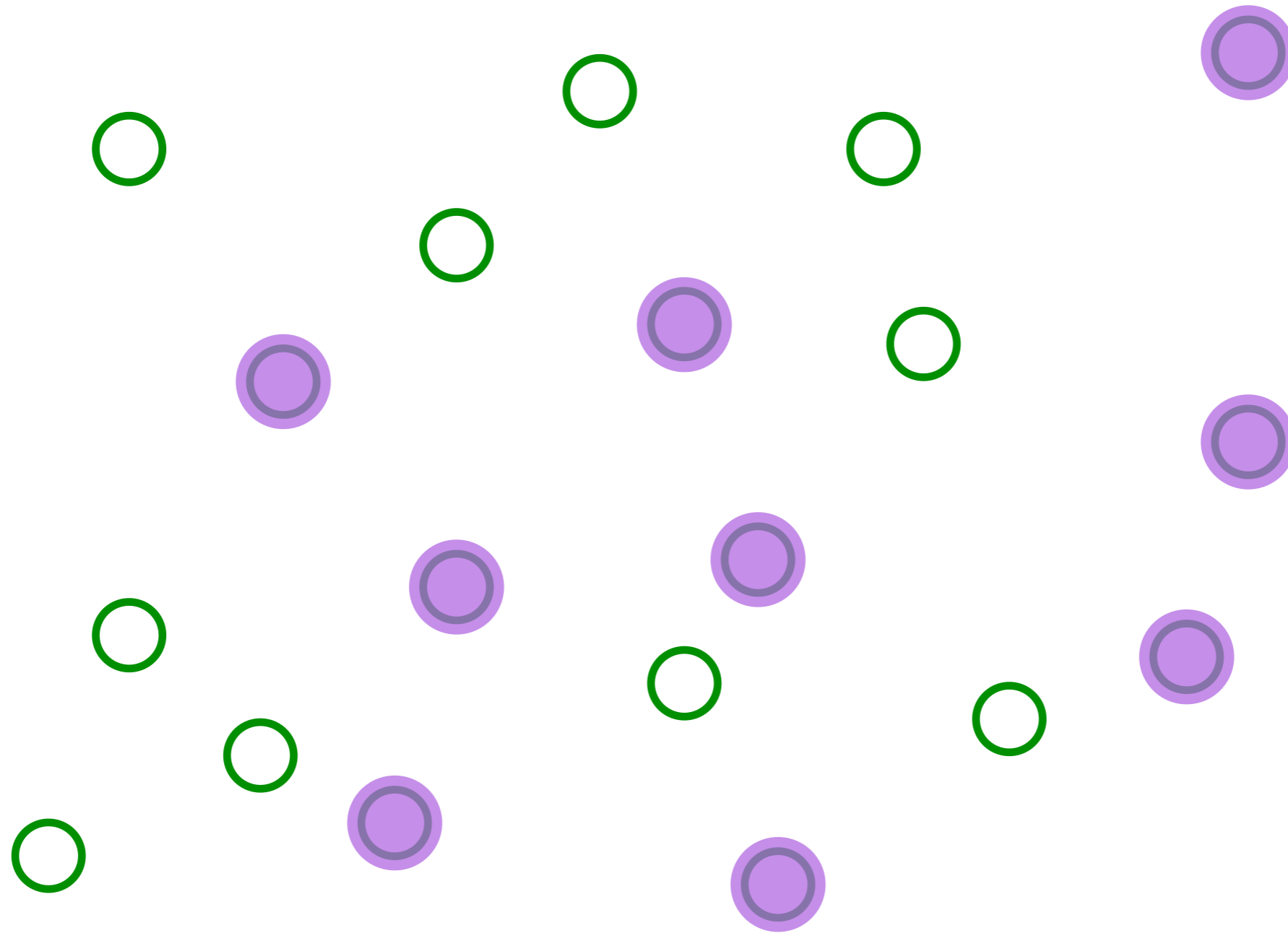
Electrons move one-by-one randomly

A simple model of a metal with quasiparticles



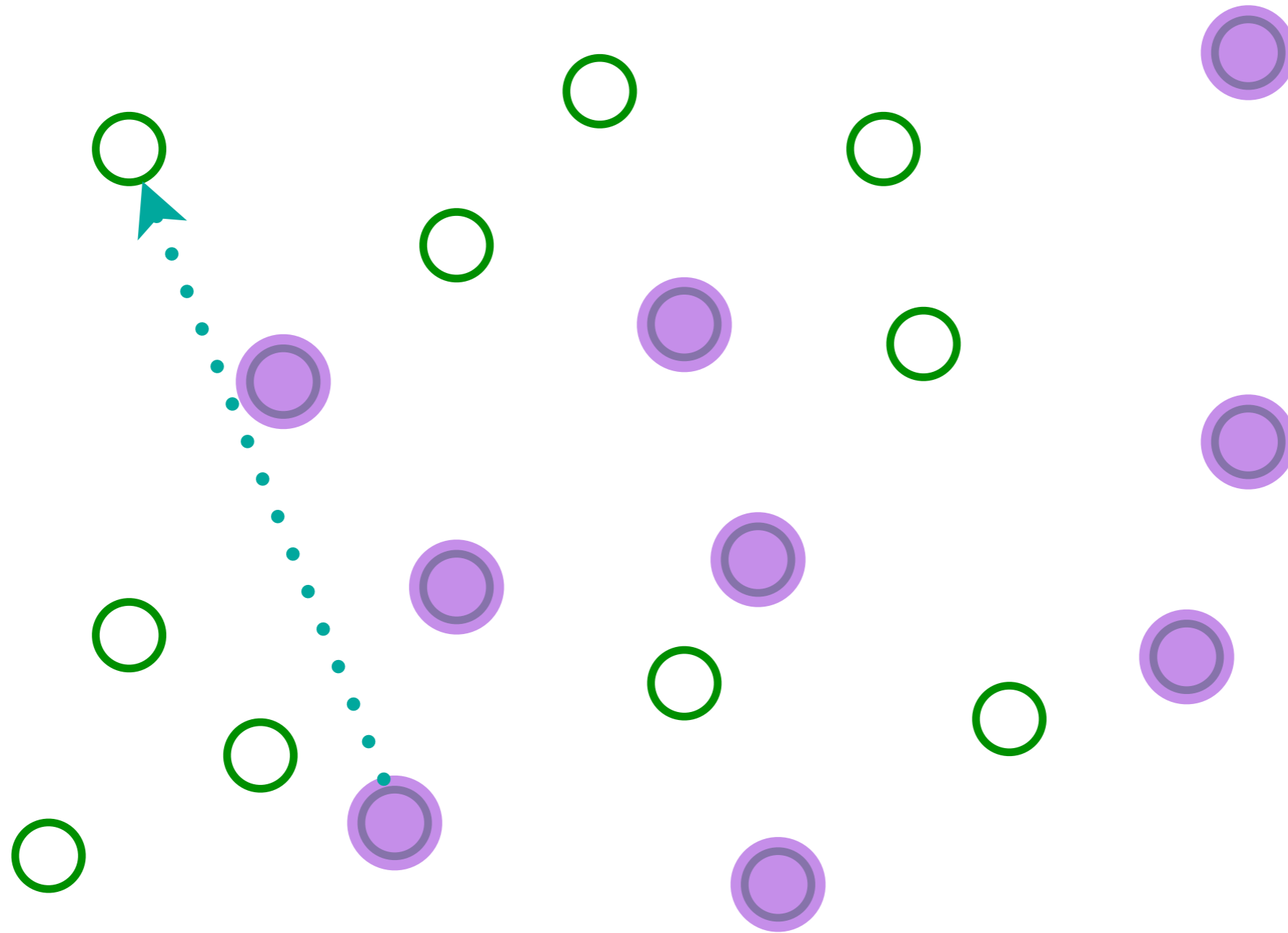
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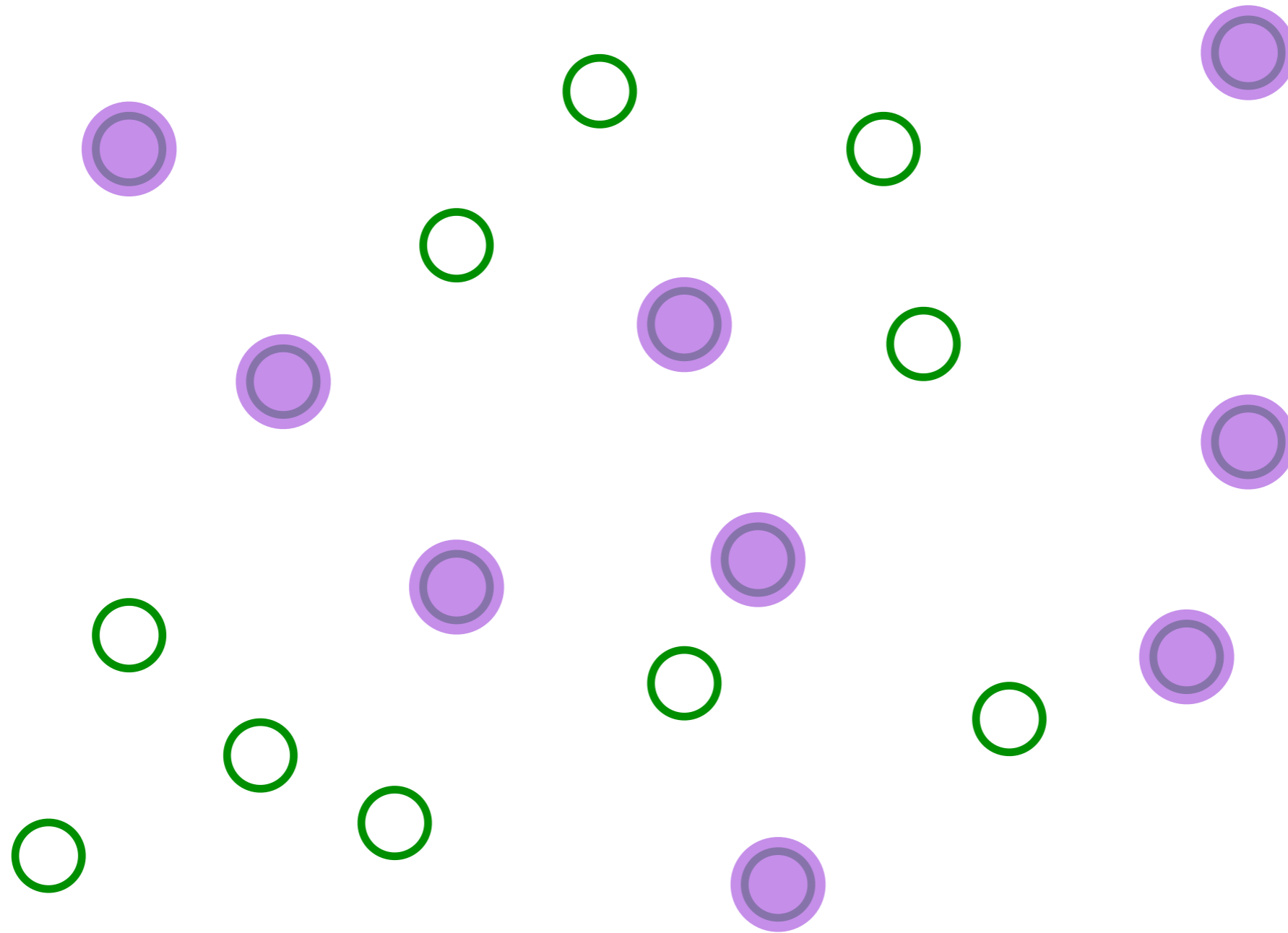
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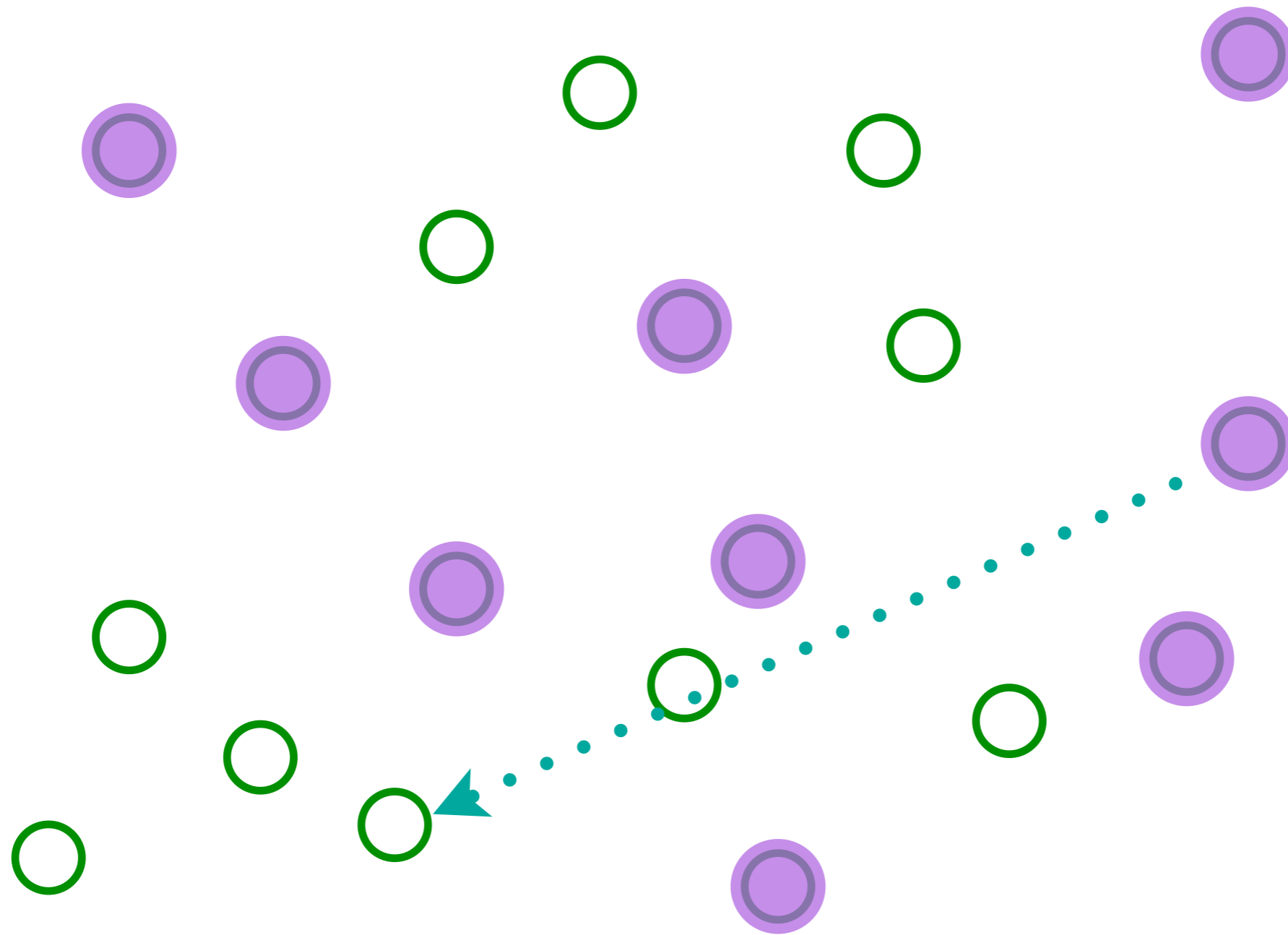
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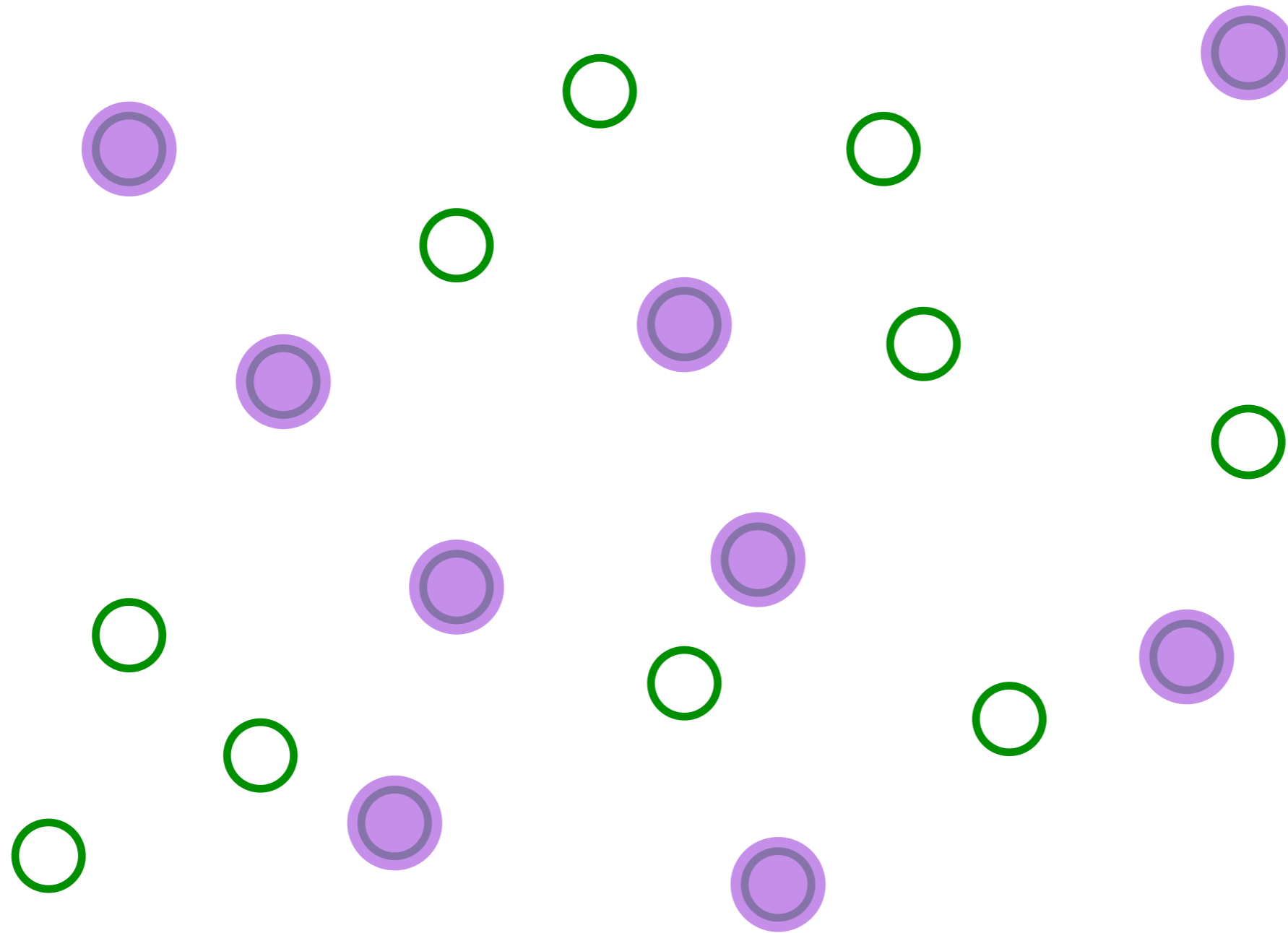
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A simple model of a metal with quasiparticles

$$H = \frac{1}{(N)^{1/2}} \sum_{i,j=1}^N t_{ij} c_i^\dagger c_j - \mu \sum_i c_i^\dagger c_i$$

$$c_i c_j + c_j c_i = 0 \quad , \quad c_i c_j^\dagger + c_j^\dagger c_i = \delta_{ij}$$

$$\frac{1}{N} \sum_i c_i^\dagger c_i = Q$$

t_{ij} are independent random variables with $\overline{t_{ij}} = 0$ and $\overline{|t_{ij}|^2} = t^2$

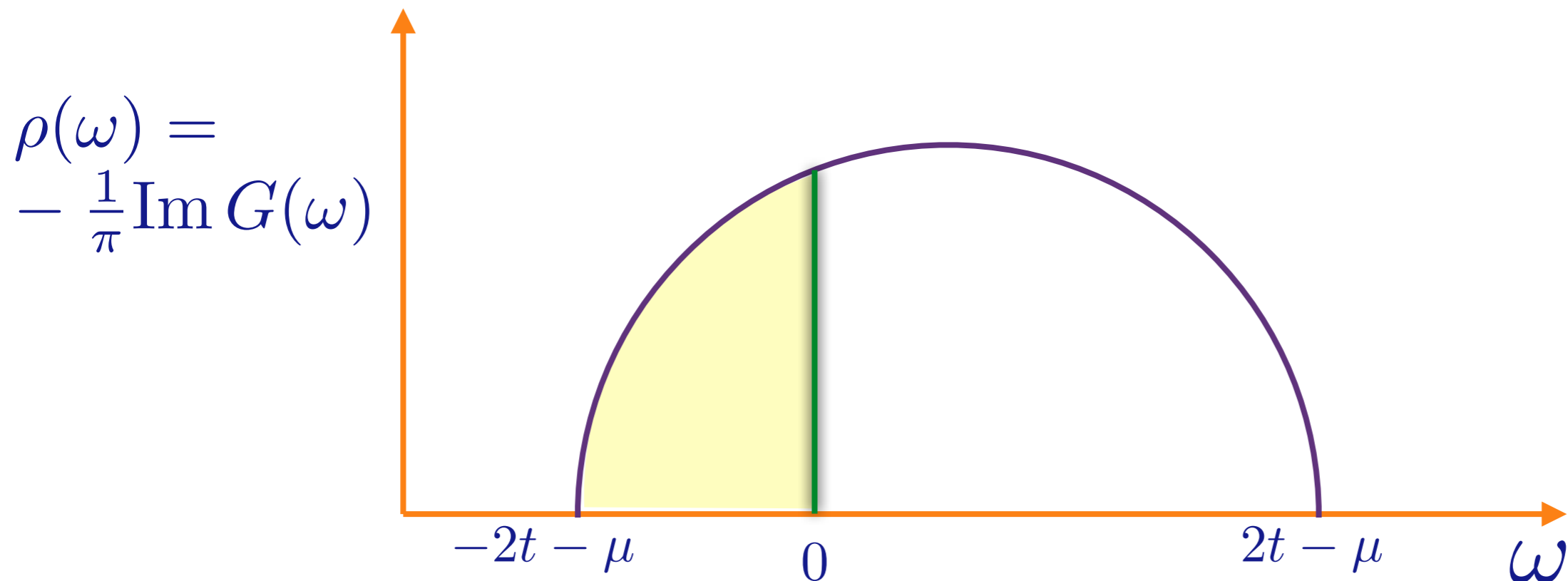
**Fermions occupying the eigenstates of a
 $N \times N$ random matrix**

A simple model of a metal with quasiparticles

Feynman graph expansion in $t_{ij..}$, and graph-by-graph average, yields exact equations in the large N limit:

$$G(\tau) \equiv -T_\tau \left\langle c_i(\tau) c_i^\dagger(0) \right\rangle$$
$$G(i\omega) = \frac{1}{i\omega + \mu - \Sigma(i\omega)} \quad , \quad \Sigma(\tau) = t^2 G(\tau)$$
$$G(\tau = 0^-) = Q.$$

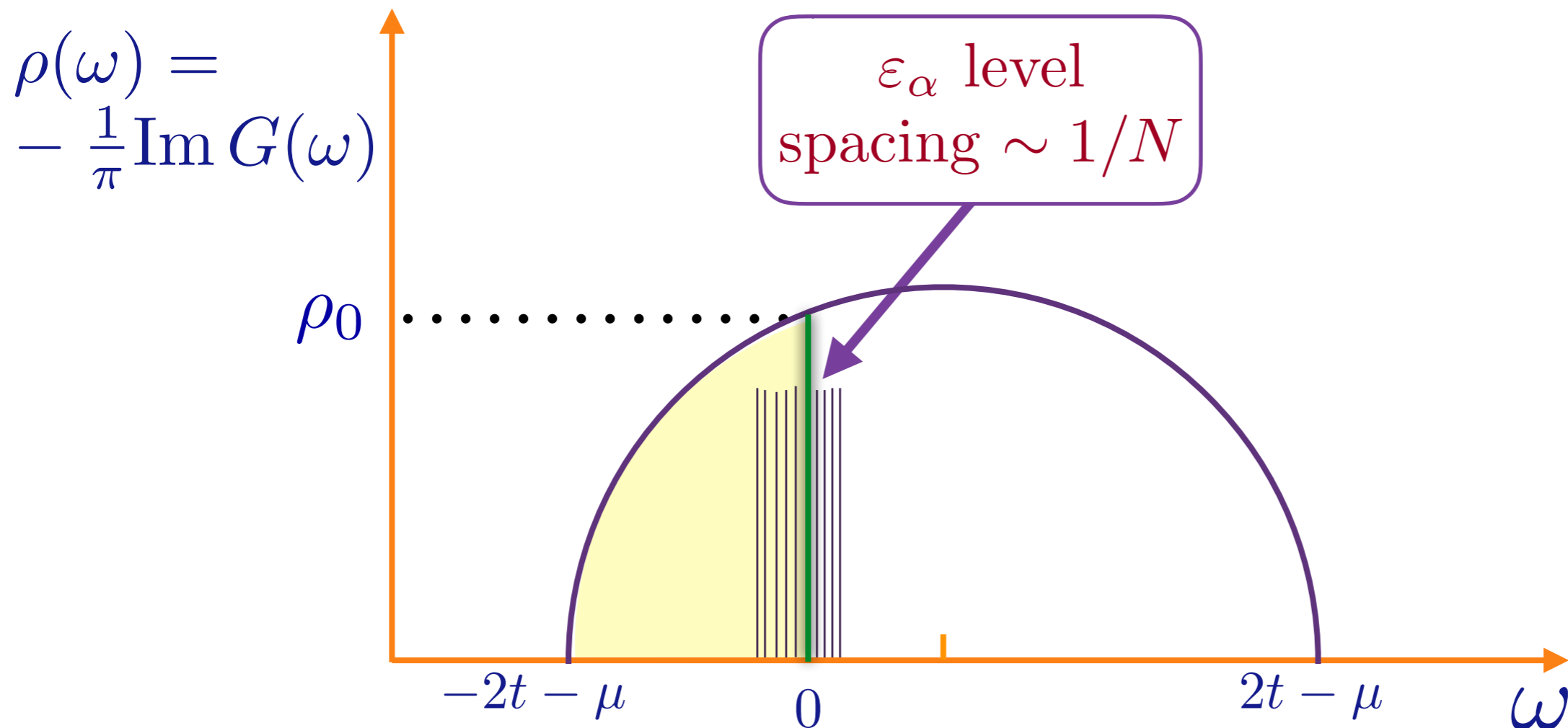
$G(\omega)$ can be determined by solving a quadratic equation.



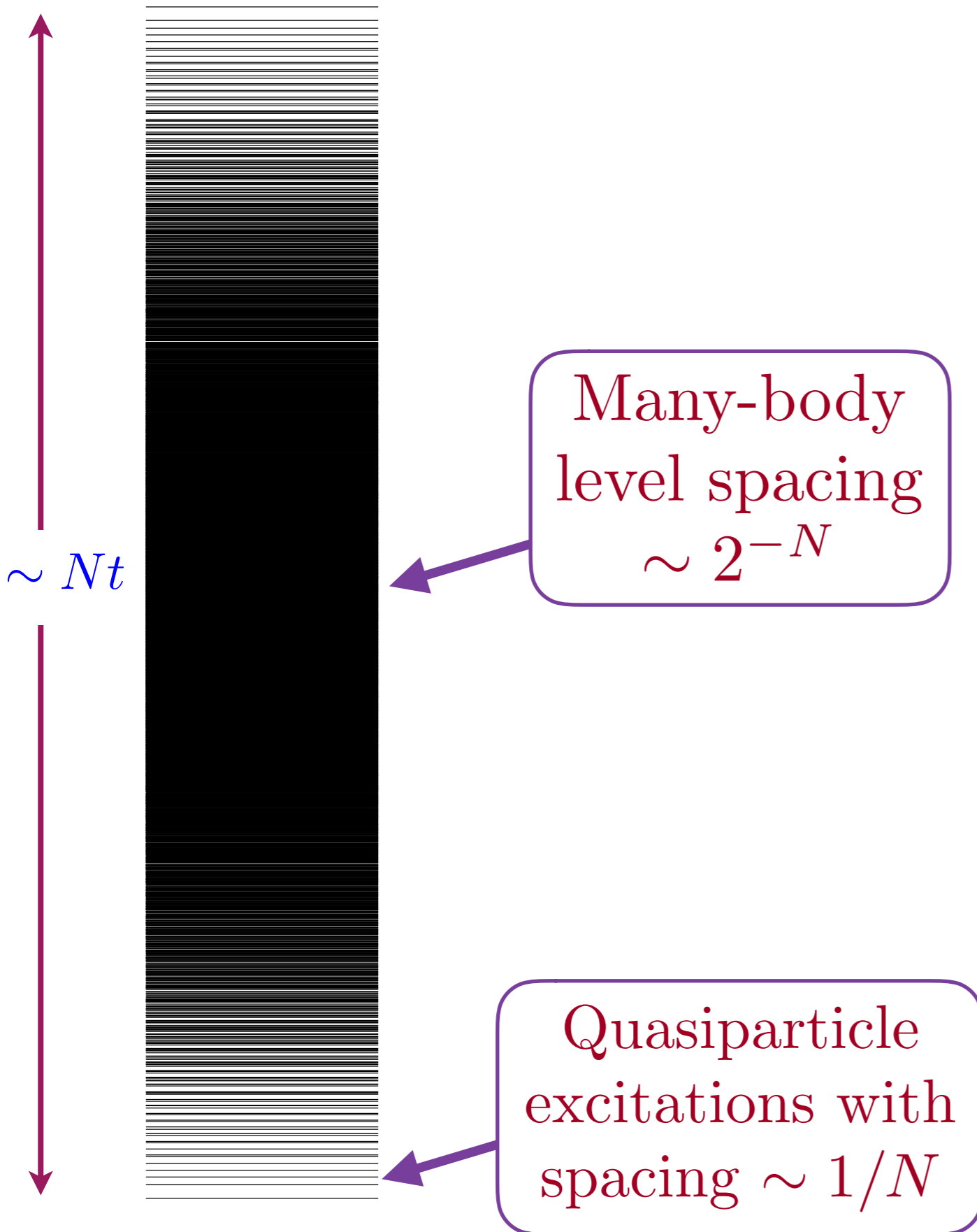
A simple model of a metal with quasiparticles

Let ε_α be the eigenvalues of the matrix t_{ij}/\sqrt{N} . The fermions will occupy the lowest NQ eigenvalues, upto the Fermi energy E_F . The single-particle density of states is

$$\rho(\omega) = (1/N) \sum_\alpha \delta(\omega - \varepsilon_\alpha), \text{ and } \rho_0 \equiv \rho(\omega = 0).$$



A simple model of a metal with quasiparticles



There are 2^N many body levels with energy

$$E = \sum_{\alpha=1}^N n_{\alpha} \varepsilon_{\alpha},$$

where $n_{\alpha} = 0, 1$. Shown are all values of E for a single cluster of size $N = 12$. The ε_{α} have a level spacing $\sim 1/N$.

A simple model of a metal with quasiparticles

The grand potential $\Omega(T)$ at low T is (from the Sommerfeld expansion)

$$\Omega(T) - E_0 = N \left(-\frac{\pi^2}{6} \rho_0 T^2 + \mathcal{O}(T^4) \right) + \dots$$

where $\rho_0 \equiv \rho(0)$ is the *single* particle density of states at the Fermi level.

We can also define the *many* body density of states, $D(E)$, via

$$Z = e^{-\Omega(T)/T} = \int_{-\infty}^{\infty} dE D(E) e^{-E/T}$$

The inversion from $\Omega(T)$ to $D(E)$ has to be performed with care (it does not commute with the $1/N$ expansion), and we obtain

$$D(E) \sim \exp \left(\pi \sqrt{\frac{2N\rho_0(E - E_0)}{3}} \right), \quad E > E_0, \quad \frac{1}{N} \ll \rho_0(E - E_0) \ll N$$

and $D(E) = 0$ for $E < E_0$. This is related to the asymptotic growth of the partitions of an integer, $p(n) \sim \exp(\pi\sqrt{2n/3})$. Near the lower bound, there are large sample-to-sample fluctuations due to variations in the lowest quasiparticle energies.

A simple model of a metal with quasiparticles

Now add weak interactions

$$H = \frac{1}{(N)^{1/2}} \sum_{i,j=1}^N t_{ij} c_i^\dagger c_j - \mu \sum_i c_i^\dagger c_i + \frac{1}{(2N)^{3/2}} \sum_{i,j,k,\ell=1}^N U_{ij;k\ell} c_i^\dagger c_j^\dagger c_k c_\ell$$

$U_{ij;k\ell}$ are independent random variables with $\overline{U_{ij;k\ell}} = 0$ and $|\overline{U_{ij;k\ell}}|^2 = U^2$. We compute the lifetime of a quasiparticle, τ_α , in an exact eigenstate $\psi_\alpha(i)$ of the free particle Hamiltonian with energy ε_α . By Fermi's Golden rule, for ε_α at the Fermi energy

$$\begin{aligned} \frac{1}{\tau_\alpha} &= \pi U^2 \rho_0^2 \int d\varepsilon_\beta d\varepsilon_\gamma d\varepsilon_\delta f(\varepsilon_\beta)(1 - f(\varepsilon_\gamma))(1 - f(\varepsilon_\delta)) \delta(\varepsilon_\alpha + \varepsilon_\beta - \varepsilon_\gamma - \varepsilon_\delta) \\ &= \frac{\pi^3 U^2 \rho_0^2}{4} T^2 \end{aligned}$$

where ρ_0 is the density of states at the Fermi energy, and $f(\varepsilon) = 1/(e^{\varepsilon/T} + 1)$ is the Fermi function.

Fermi liquid state: Two-body interactions lead to a scattering time of quasiparticle excitations from in (random) single-particle eigenstates which diverges as $\sim T^{-2}$ at the Fermi level.

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$q=2$, complex SYK

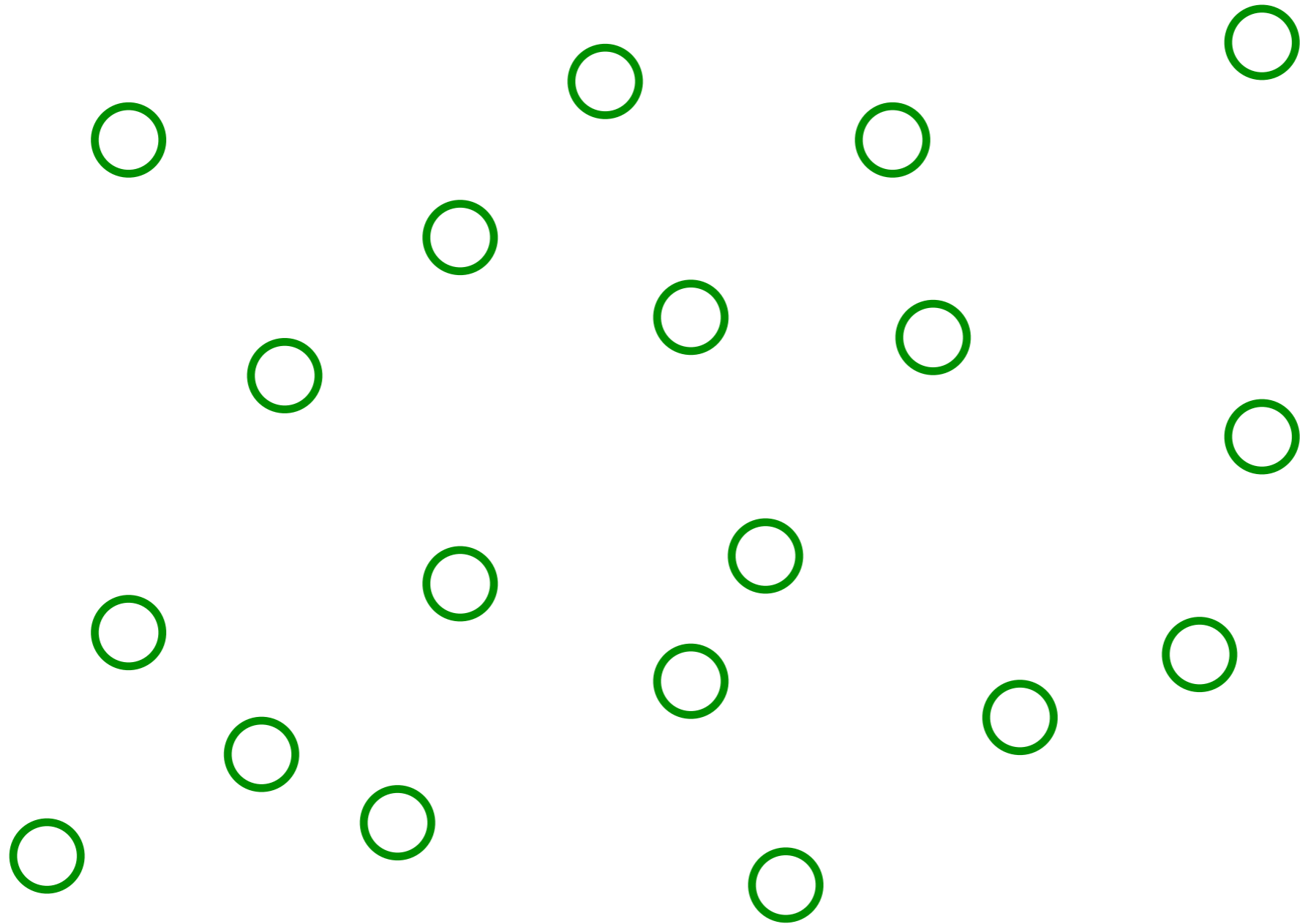
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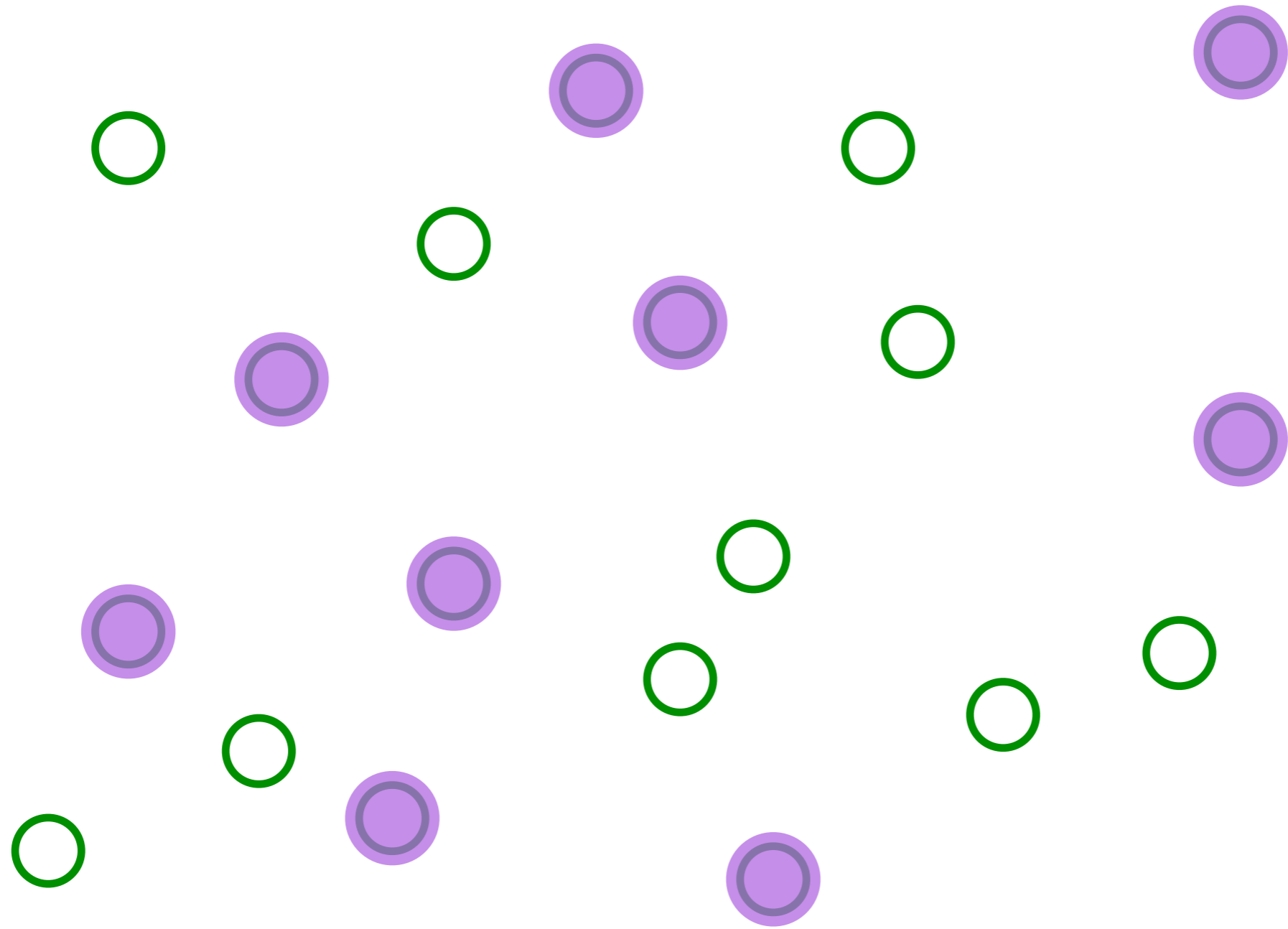
4. Connections to black holes
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The Sachdev-Ye-Kitaev (SYK) model



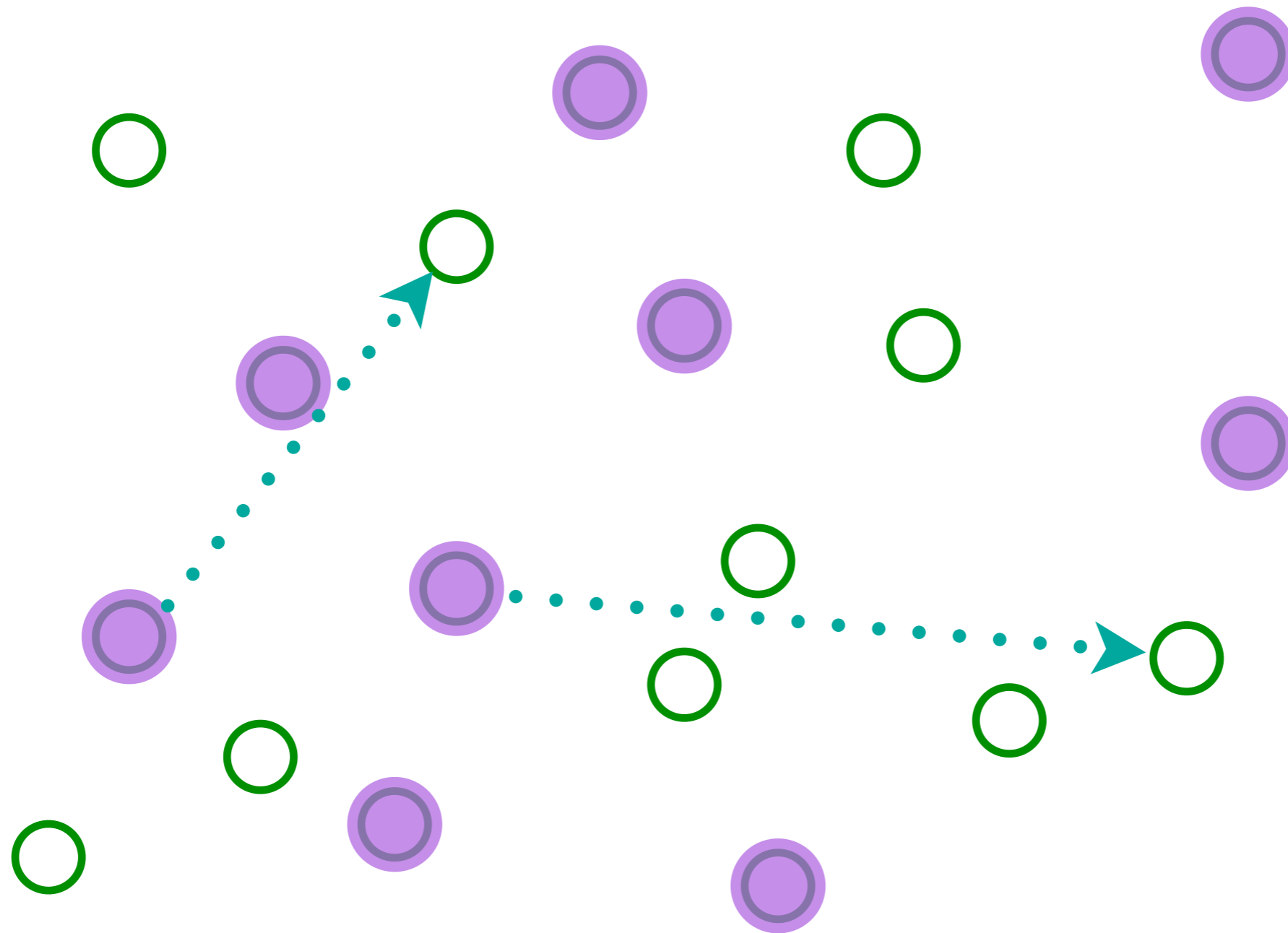
Pick a set of random positions

The SYK model



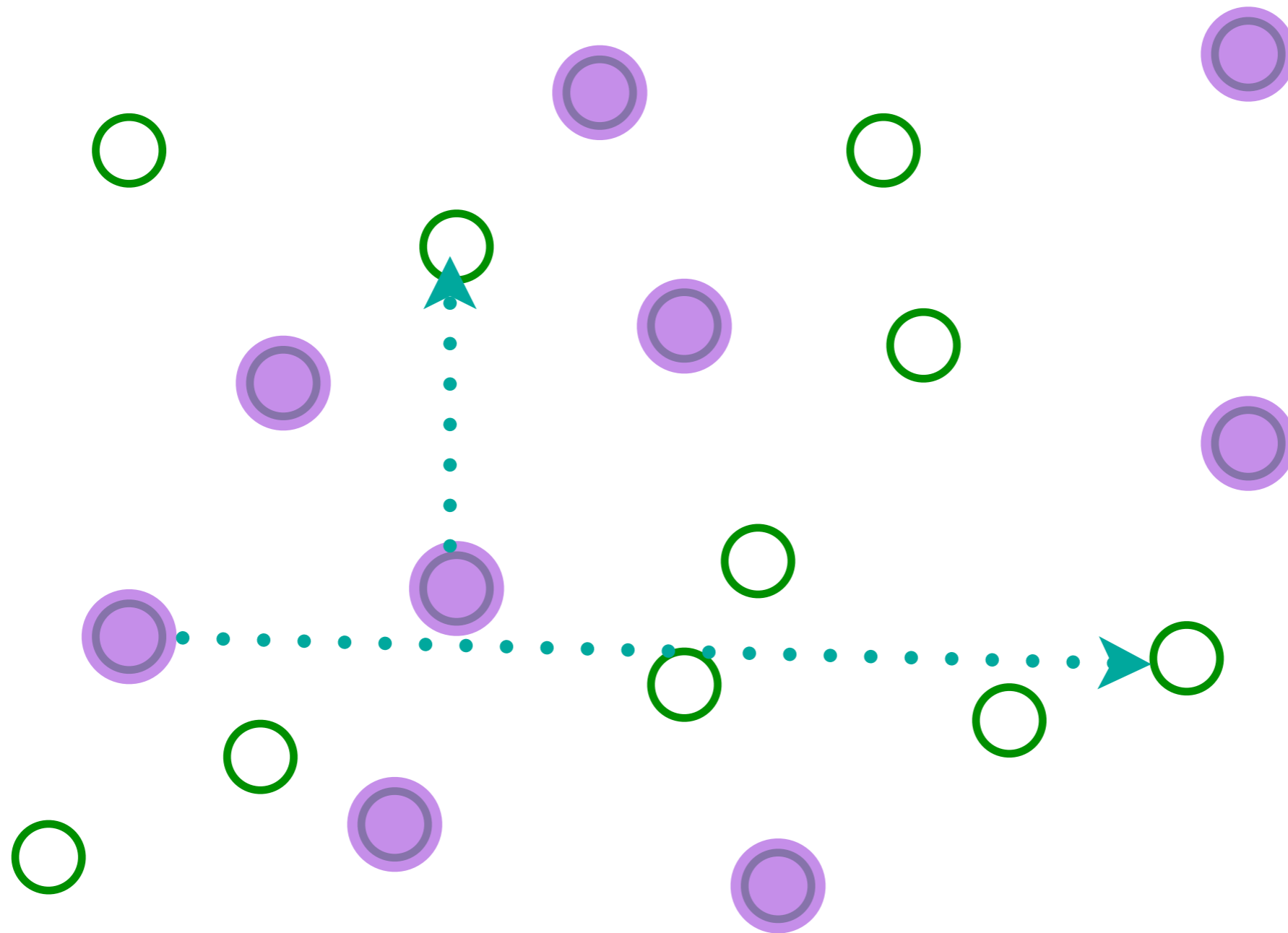
Place electrons randomly on some sites

The SYK model



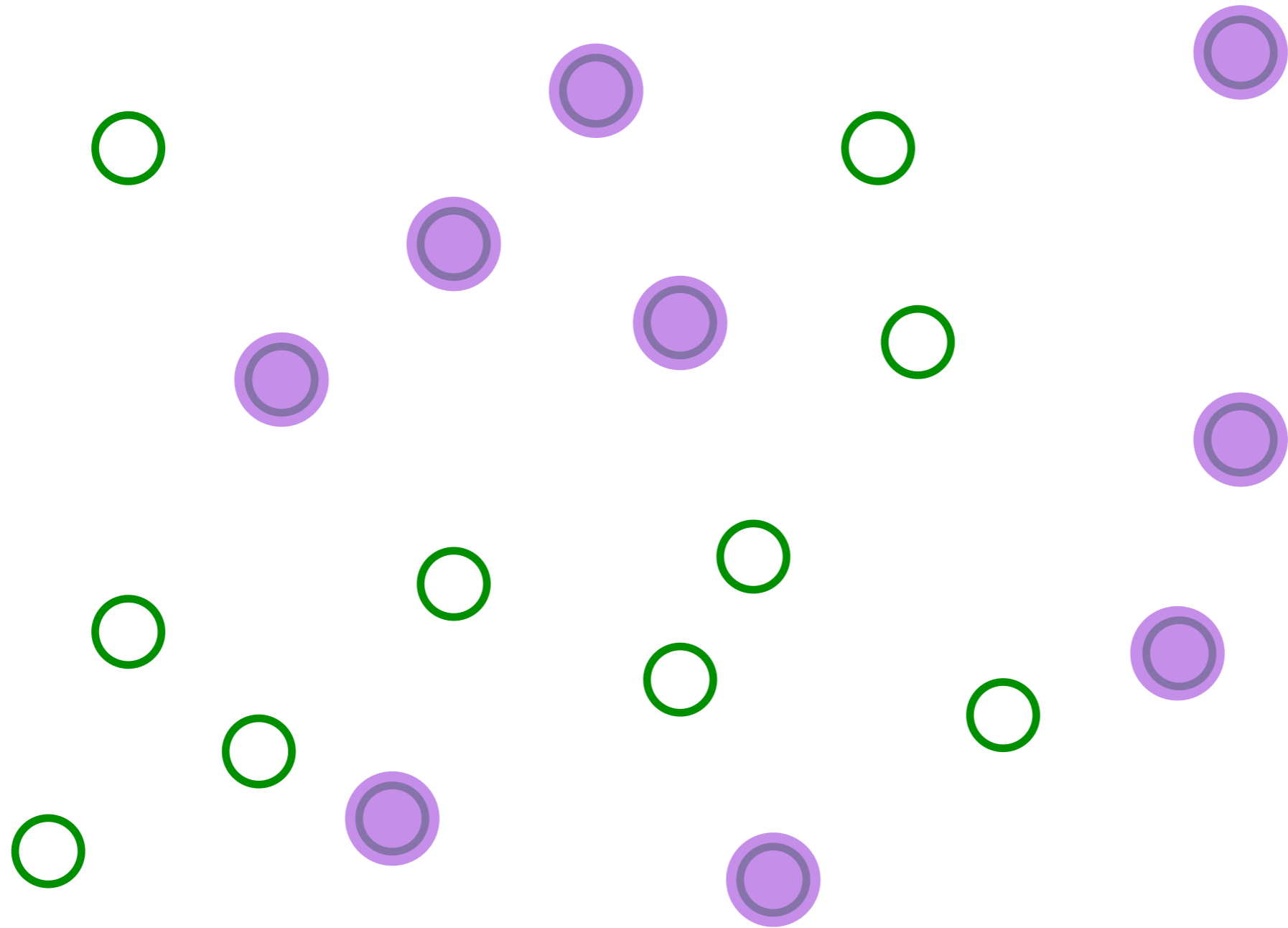
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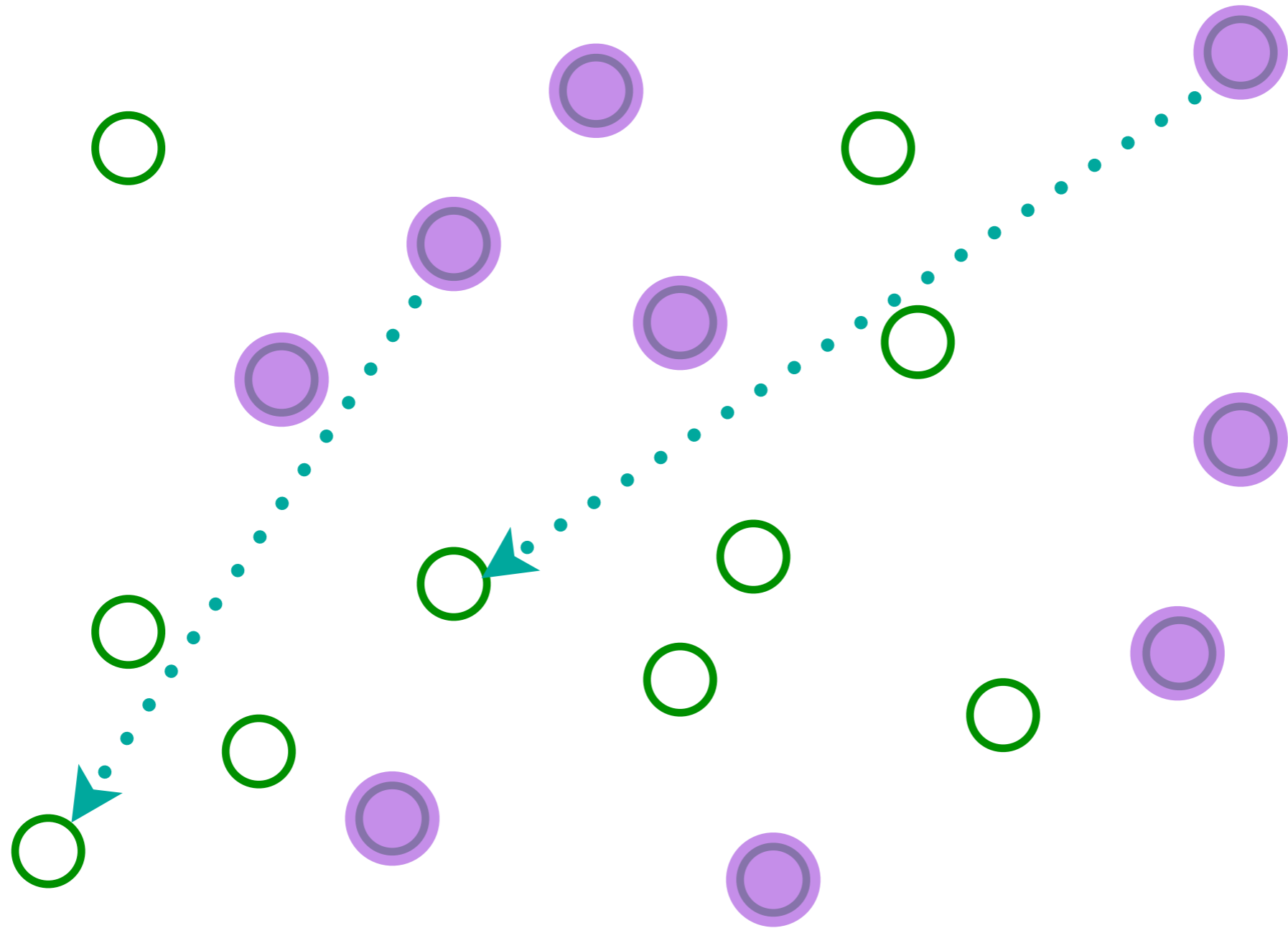
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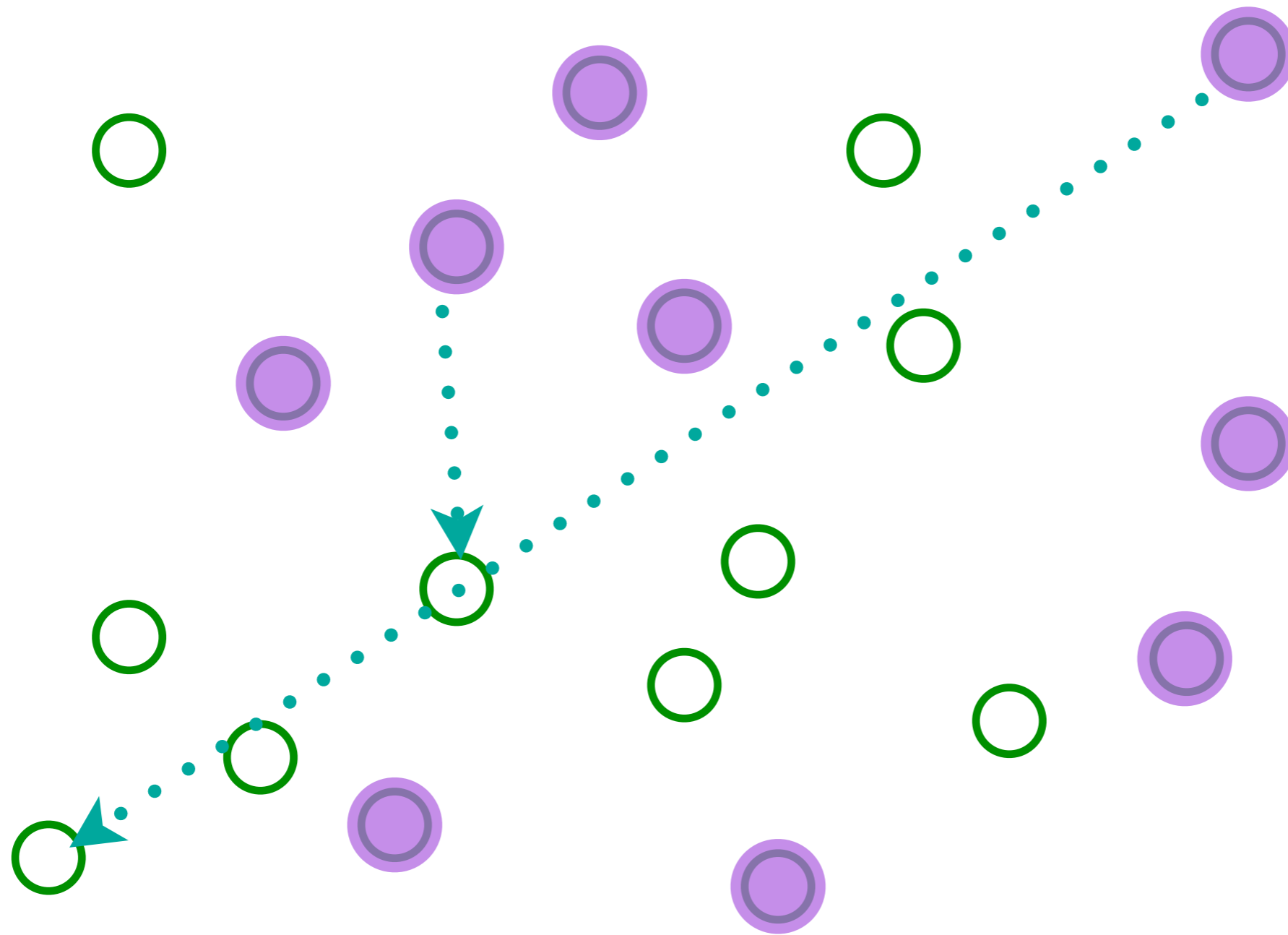
Entangle electrons pairwise randomly

The SYK model



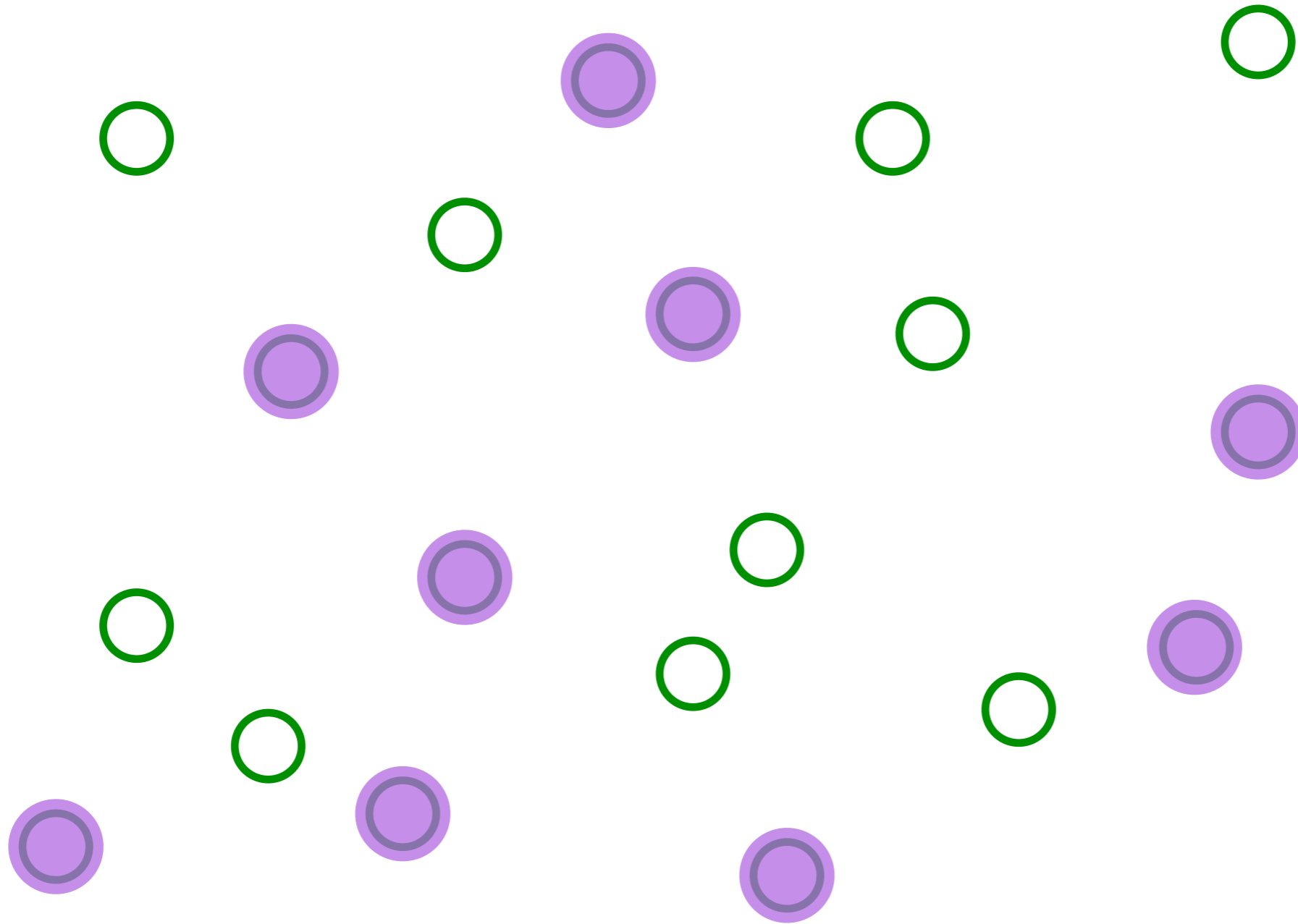
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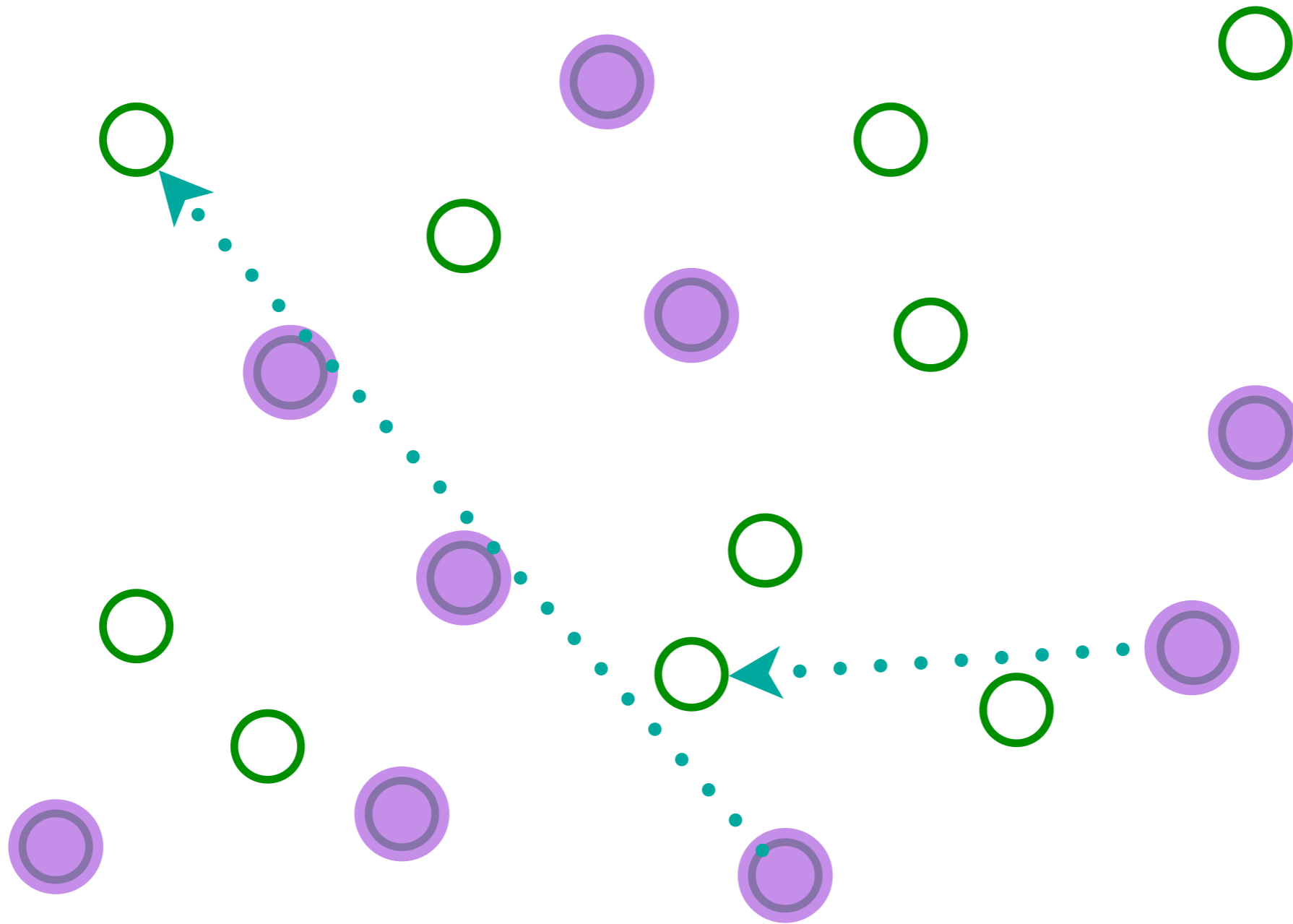
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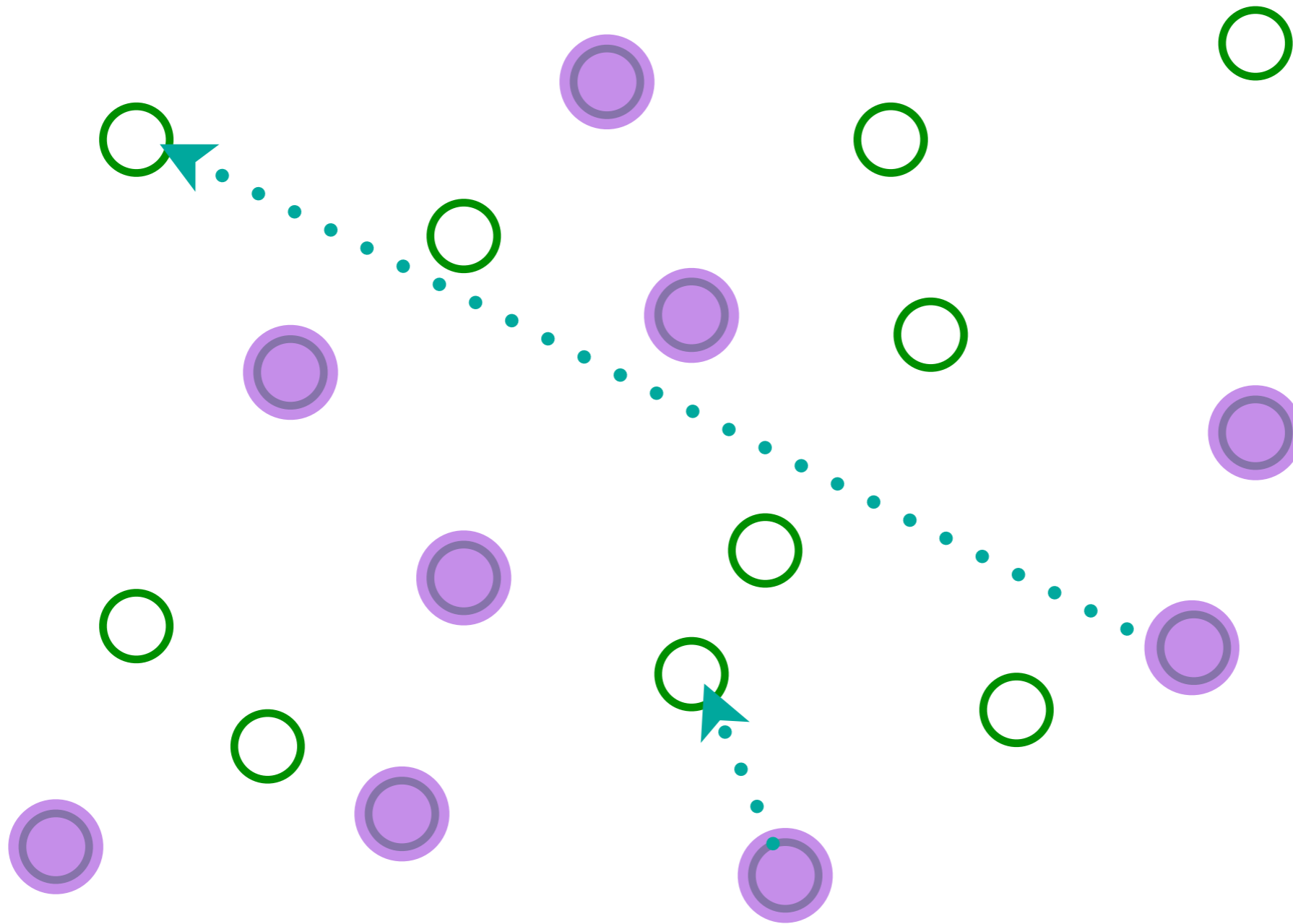
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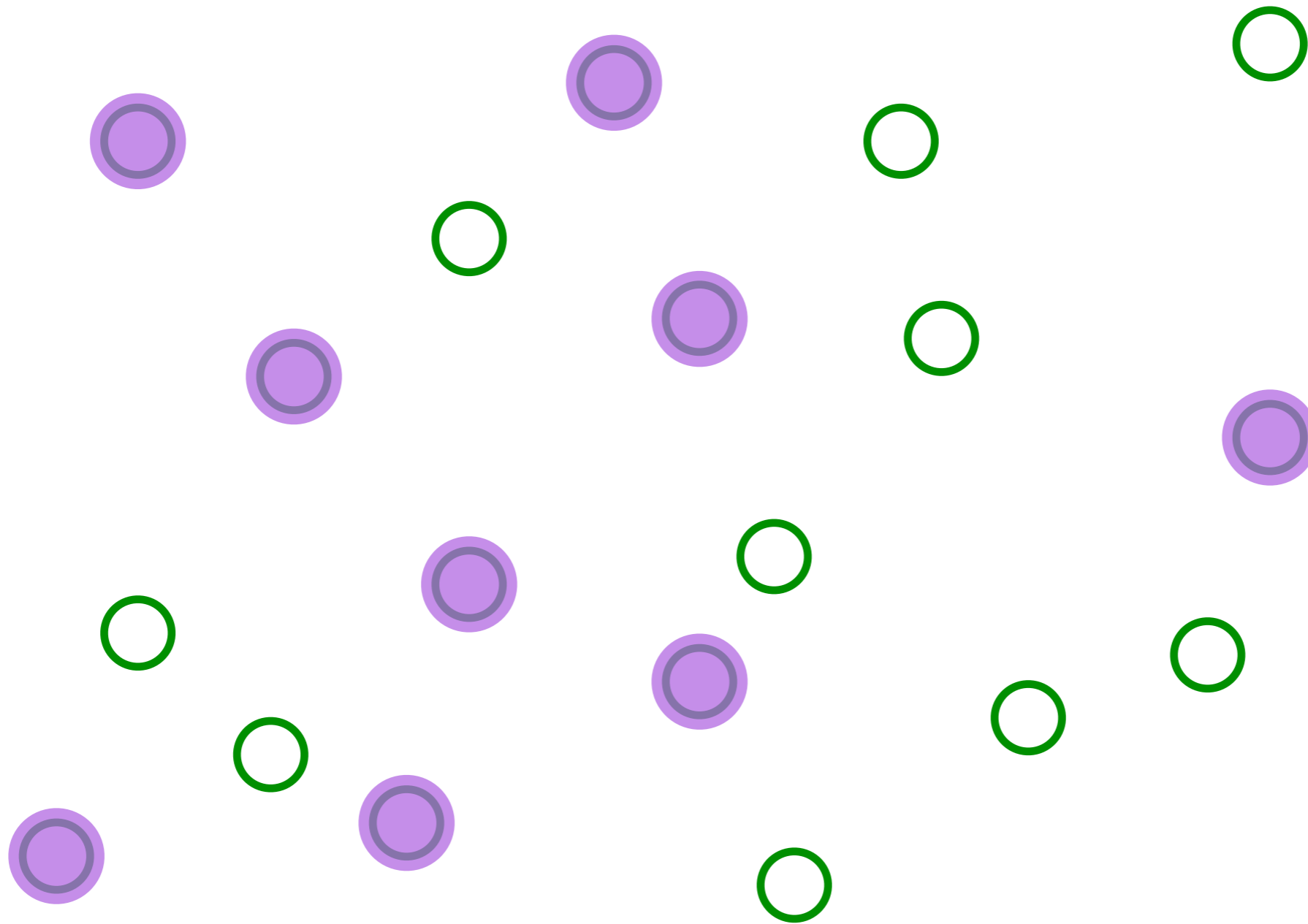
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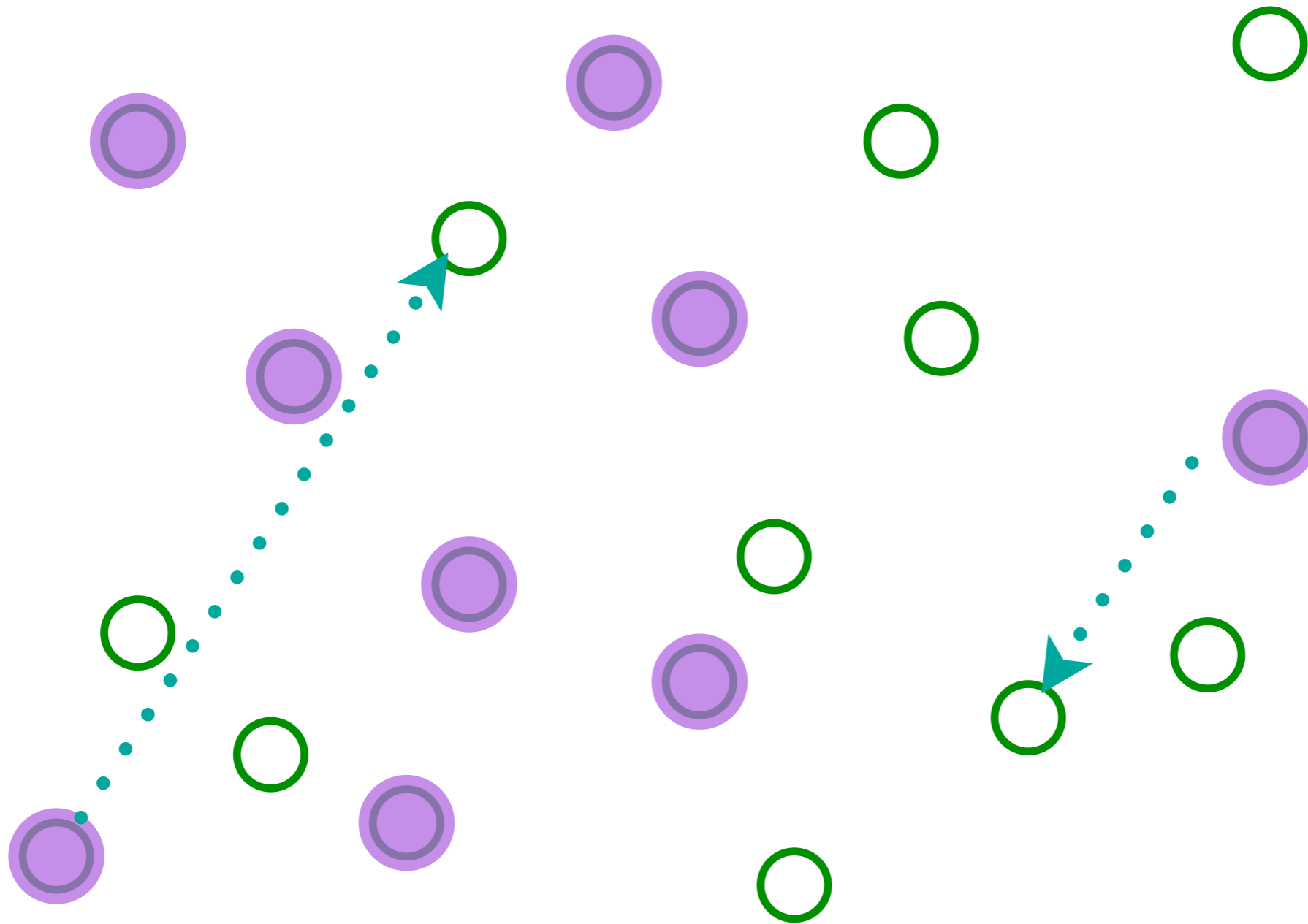
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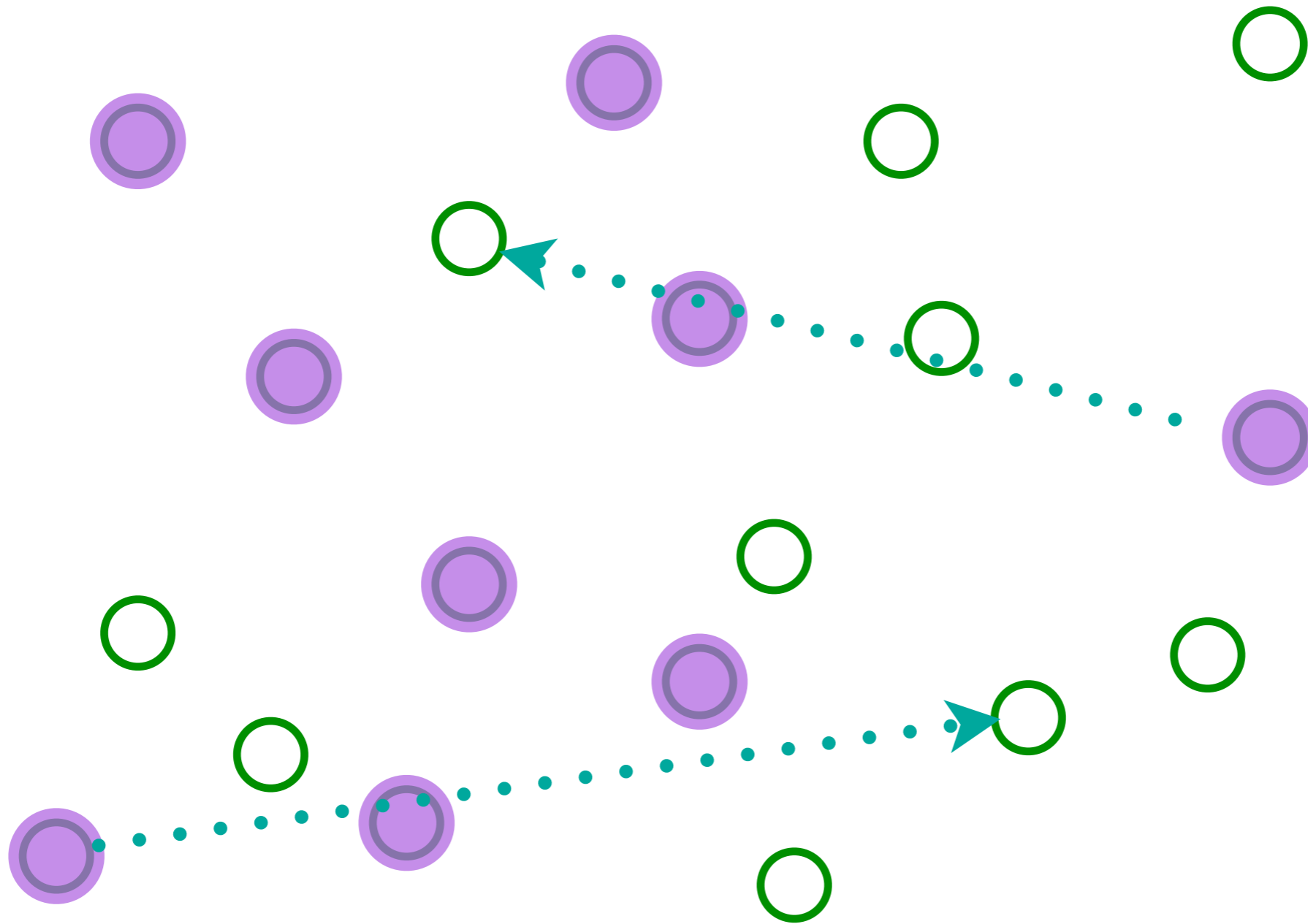
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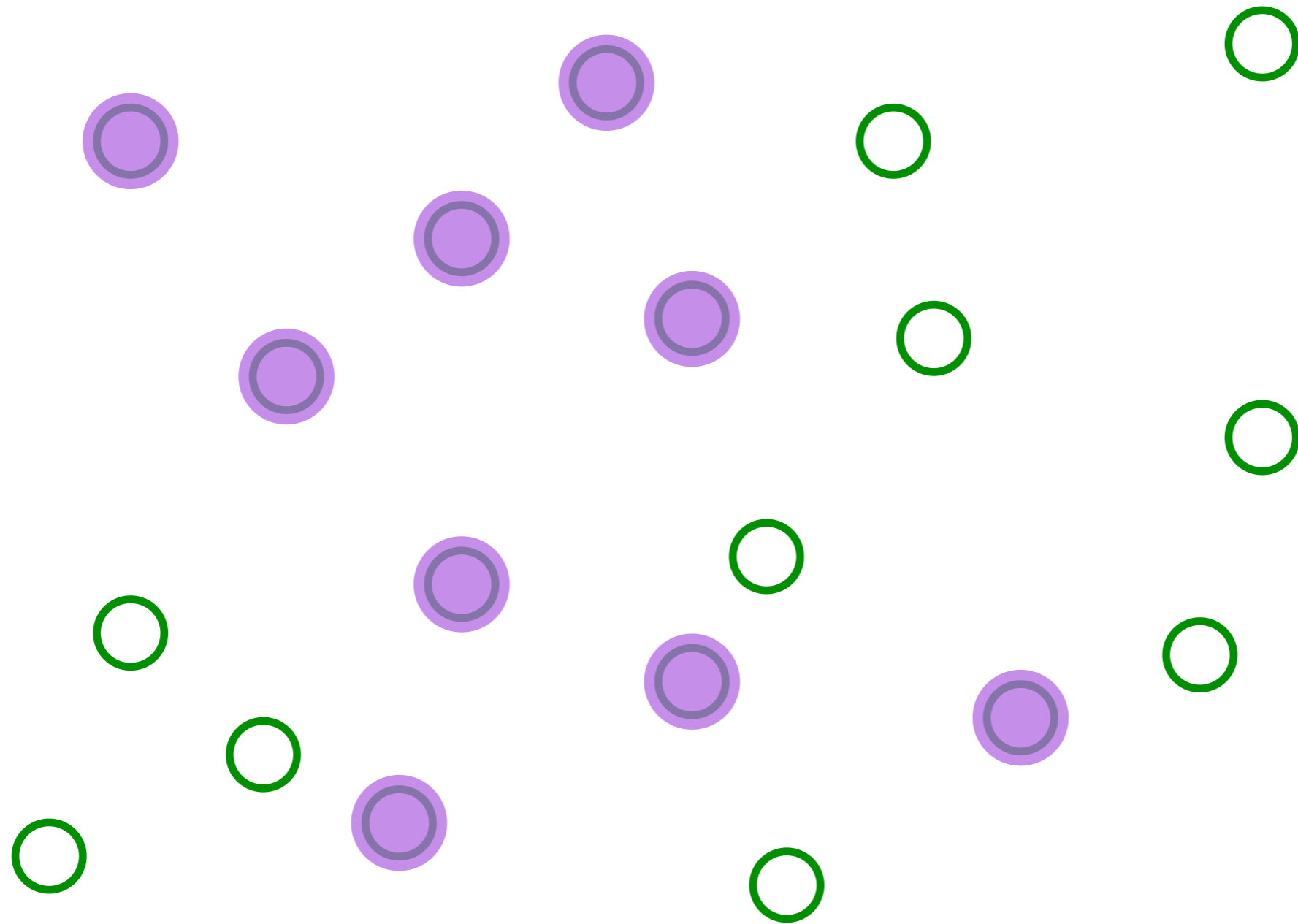
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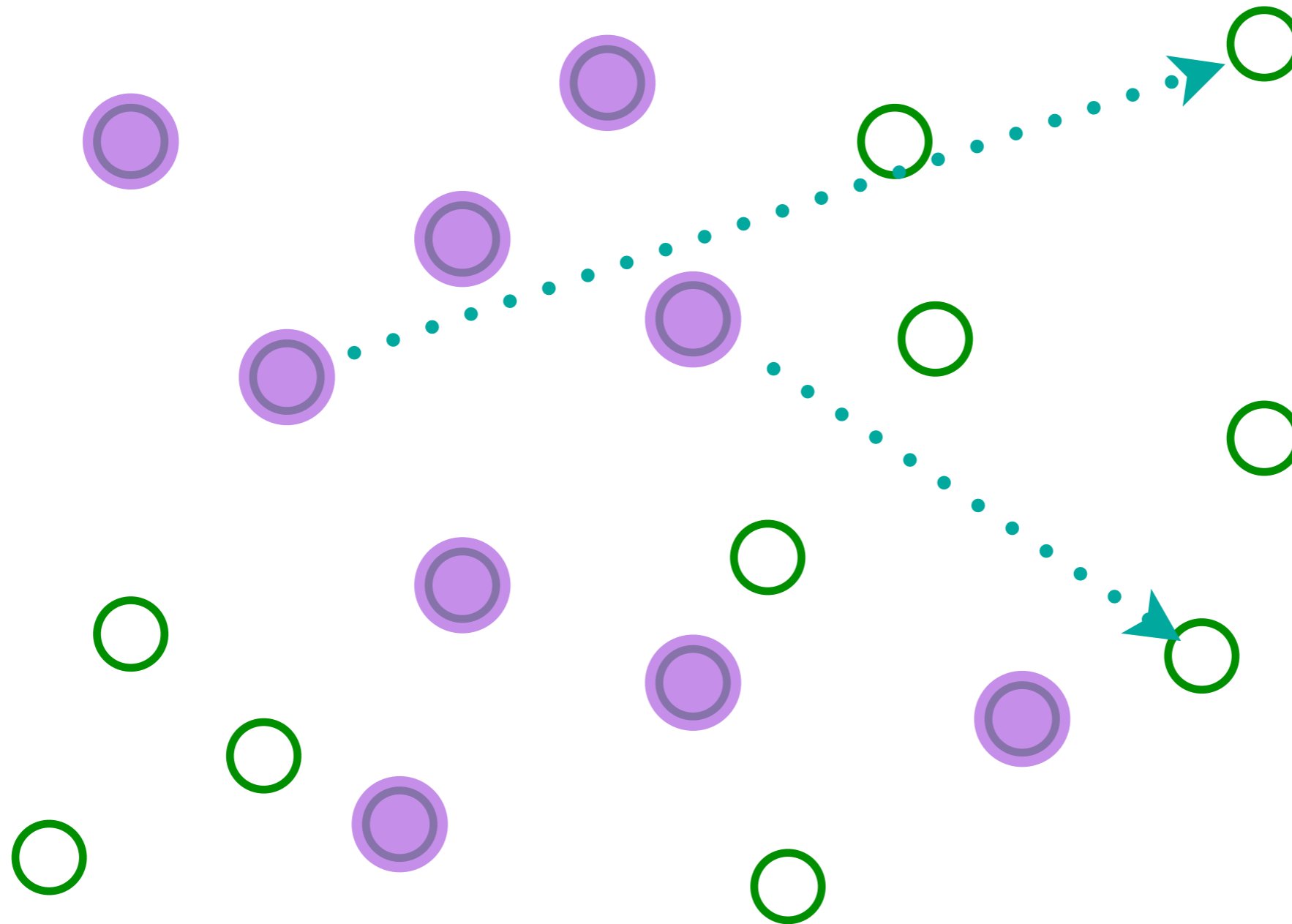
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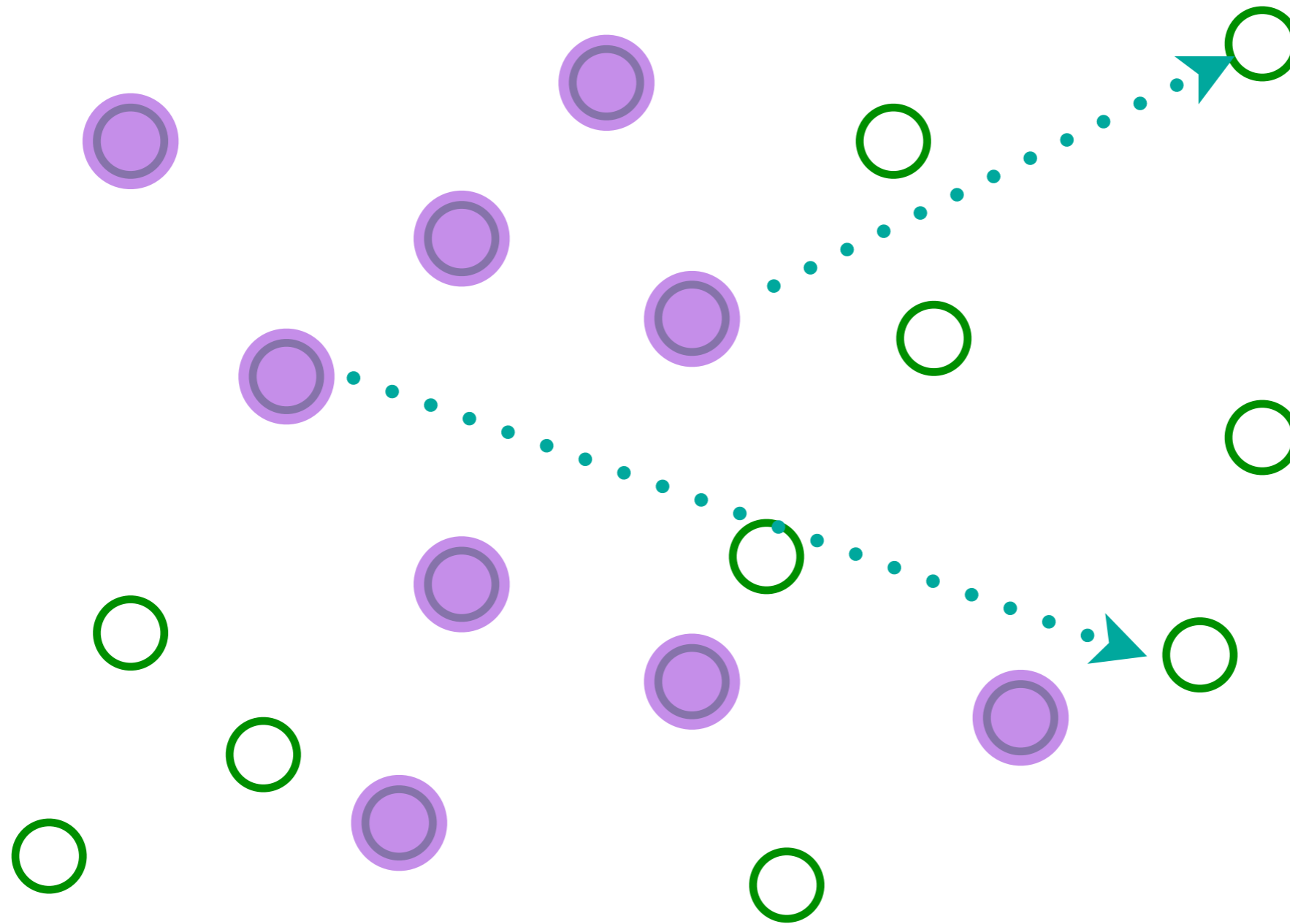
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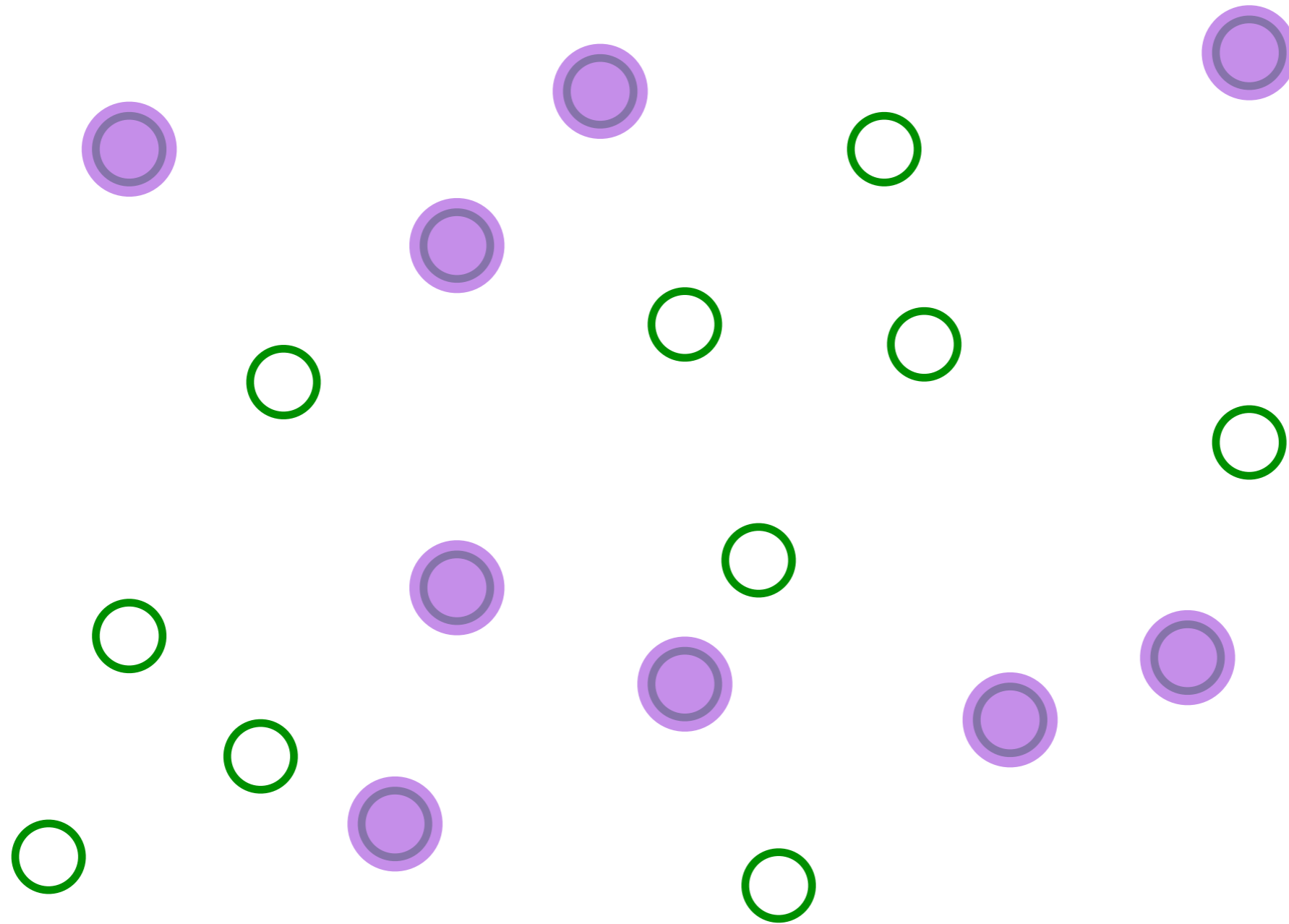
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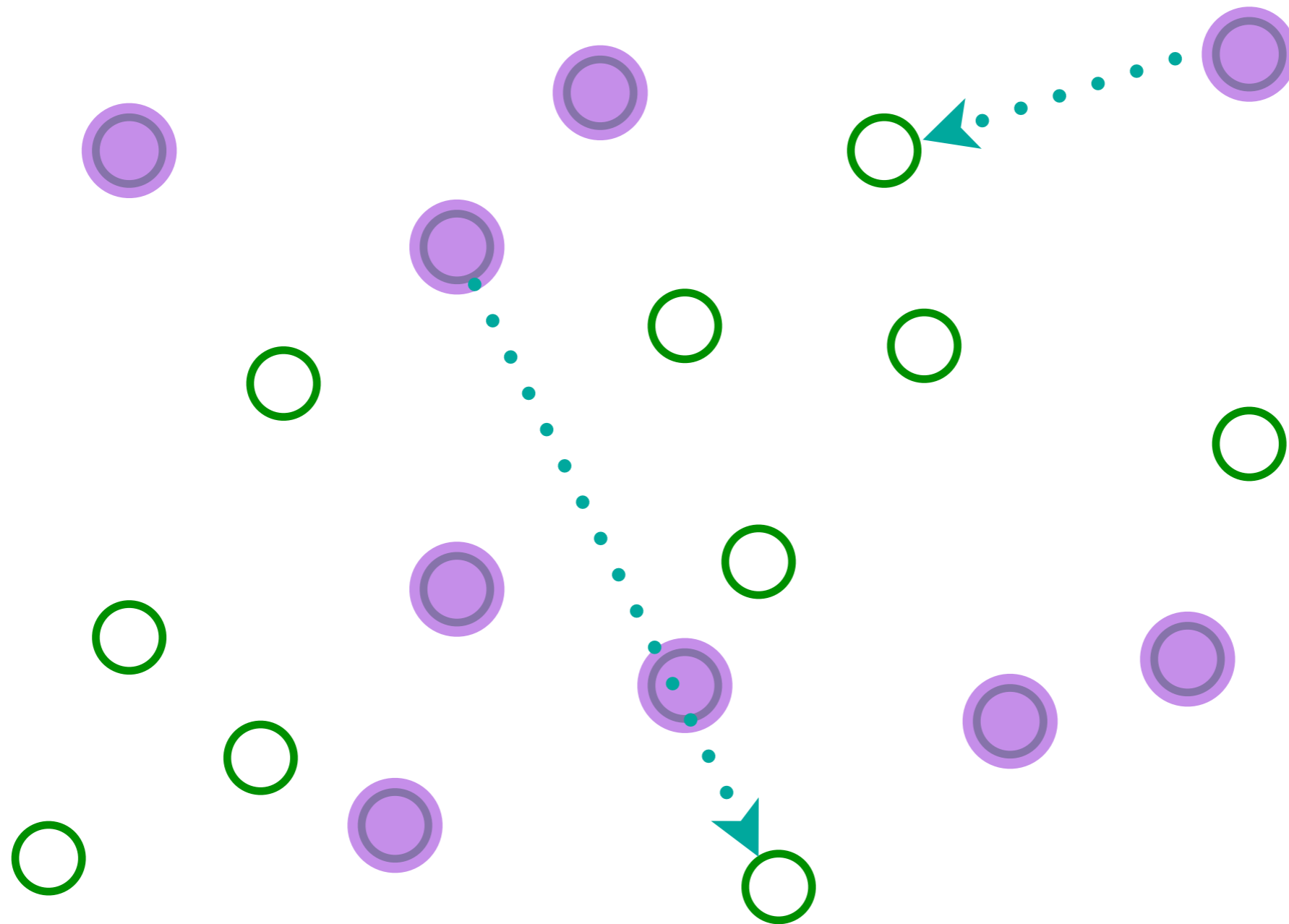
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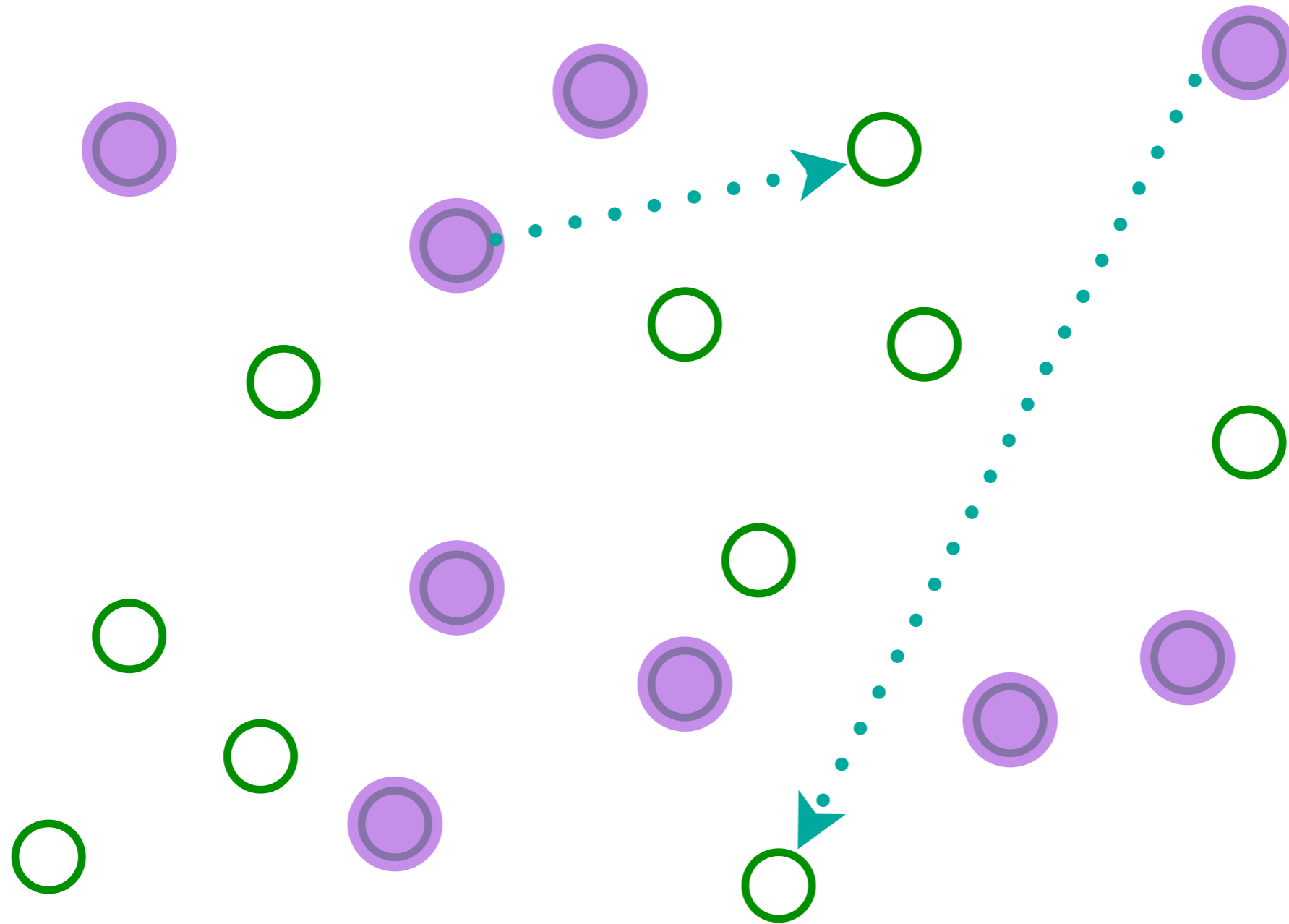
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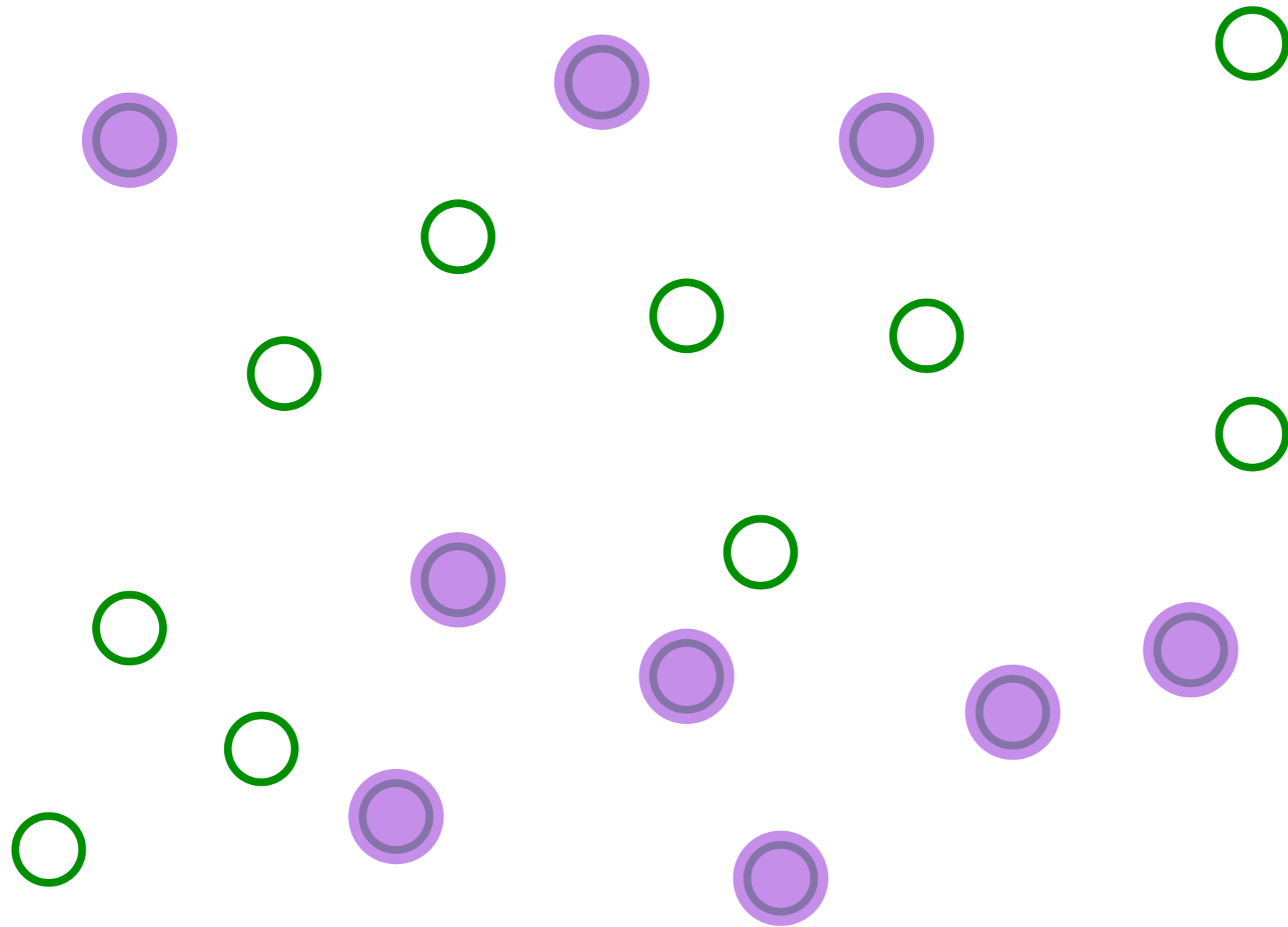
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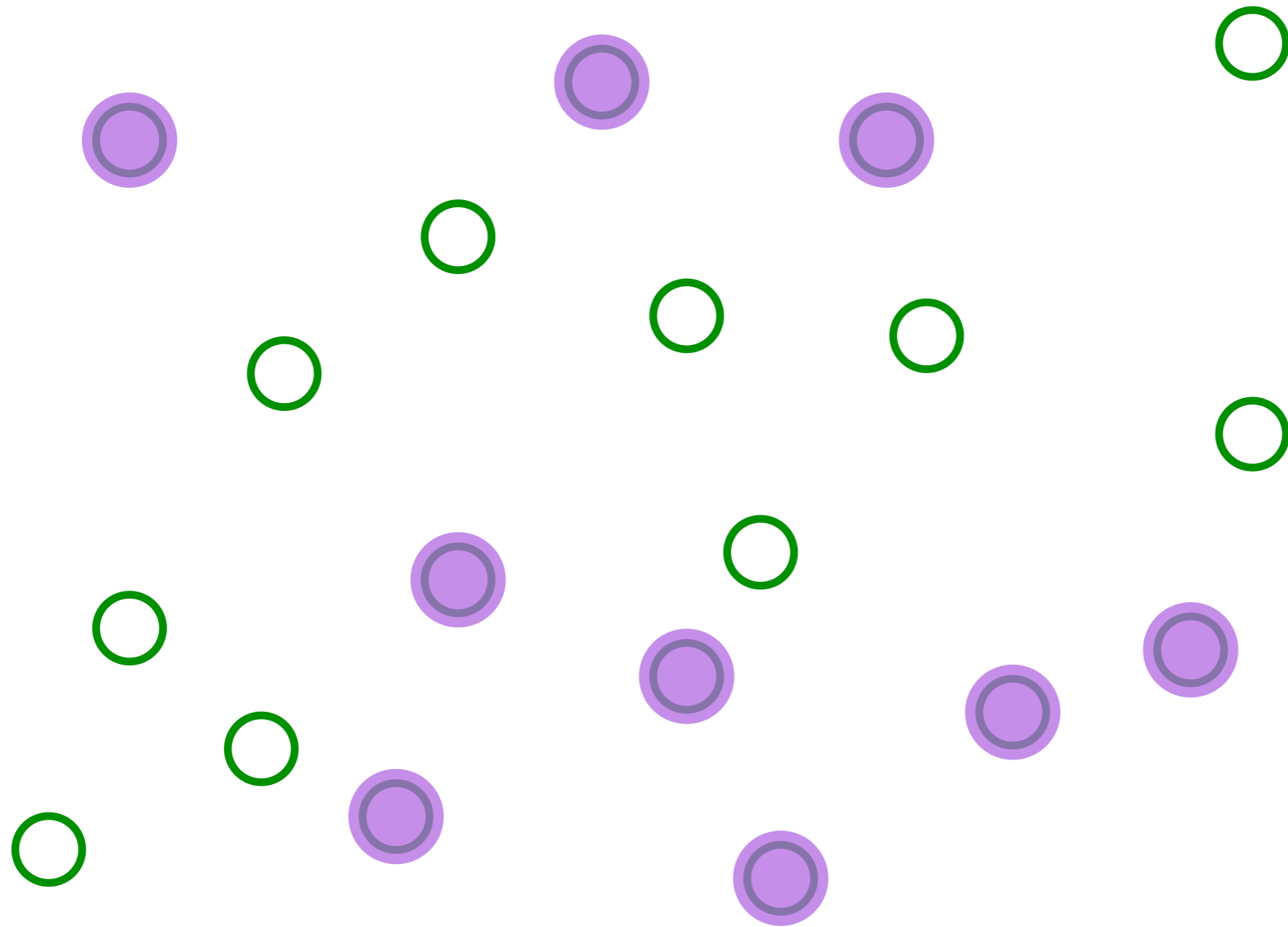
Entangle electrons pairwise randomly

The SYK model



Entangle electrons pairwise randomly

The SYK model



This describes both a strange metal and a black hole!

The SYK model

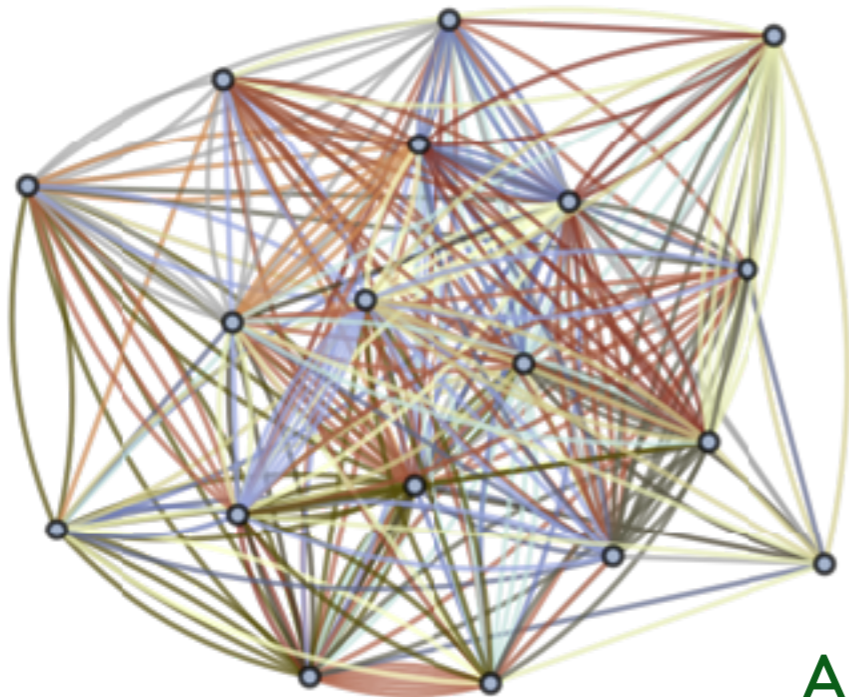
(See also: the “2-Body Random Ensemble” in nuclear physics; did not obtain the large N limit; T.A. Brody, J. Flores, J.B. French, P.A. Mello, A. Pandey, and S.S.M. Wong, Rev. Mod. Phys. **53**, 385 (1981))

$$H = \frac{1}{(2N)^{3/2}} \sum_{i,j,k,\ell=1}^N U_{ij;k\ell} c_i^\dagger c_j^\dagger c_k c_\ell - \mu \sum_i c_i^\dagger c_i$$

$$c_i c_j + c_j c_i = 0 \quad , \quad c_i c_j^\dagger + c_j^\dagger c_i = \delta_{ij}$$

$$Q = \frac{1}{N} \sum_i c_i^\dagger c_i$$

$U_{ij;k\ell}$ are independent random variables with $\overline{U_{ij;k\ell}} = 0$ and $\overline{|U_{ij;k\ell}|^2} = U^2$
 $N \rightarrow \infty$ yields critical strange metal.



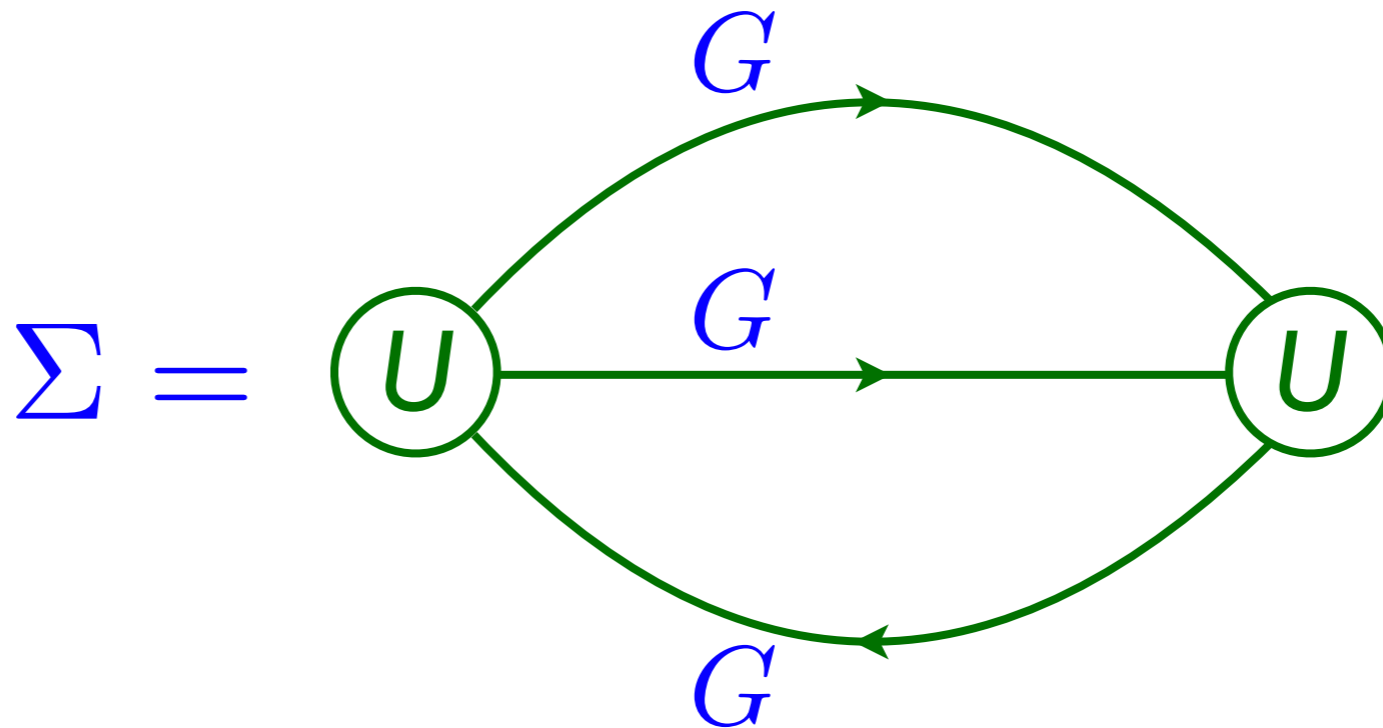
S. Sachdev and J. Ye, PRL **70**, 3339 (1993)

A. Kitaev, unpublished; S. Sachdev, PRX **5**, 041025 (2015)

The SYK model

Feynman graph expansion in U_{ijkl} , and graph-by-graph average, yields exact equations in the large N limit:

$$G(i\omega) = \frac{1}{i\omega + \mu - \Sigma(i\omega)} \quad , \quad \Sigma(\tau) = -U^2 G^2(\tau) G(-\tau)$$
$$G(\tau = 0^-) = \mathcal{Q}.$$



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$$G(\tau = 0^-) = \mathcal{Q}.$$

Low frequency analysis shows that the solutions must be gapless and obey

$$\Sigma(z) = \mu - \frac{e^{i(\pi/4+\theta)}}{A} \sqrt{z} + \dots \quad , \quad G(z) = \frac{A e^{-i(\pi/4+\theta)}}{\sqrt{z}}$$

where $A = (\pi/U^2 \cos(2\theta))^{1/4}$. The value of θ is universally related to \mathcal{Q} by a Luttinger-Ward functional analysis similar to that used to establish the Luttinger theorem of Fermi liquid theory:

$$\mathcal{Q} = \frac{1}{2} - \frac{\theta}{\pi} - \frac{\sin(2\theta)}{4}$$

S. Sachdev and J. Ye, Phys. Rev. Lett. **70**, 3339 (1993)

A. Georges, O. Parcollet, and S. Sachdev, PRB **63**, 134406 (2001)

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$$G(i\omega) = \frac{1}{i\omega + \mu - \Sigma(i\omega)} \quad , \quad \Sigma(\tau) = -U^2 G^2(\tau) G(-\tau)$$
$$G(\tau = 0^-) = Q.$$

At $T > 0$, we obtain a solution with a conformal structure

$$G(\tau) = -A \frac{e^{-2\pi\mathcal{E}T\tau}}{\sqrt{1 + e^{-4\pi\mathcal{E}}}} \left(\frac{T}{\sin(\pi T\tau)} \right)^{1/2} \quad , \quad 0 < \tau < 1/T \quad ,$$

where the ‘particle-hole asymmetry’ is determined by \mathcal{E}

$$e^{2\pi\mathcal{E}} = \frac{\sin(\pi/4 + \theta)}{\sin(\pi/4 - \theta)} .$$

The SYK model

There are 2^N many body levels with energy E . Shown are all values of E for a single cluster of size $N = 12$. The $T \rightarrow 0$ state has an entropy $S_{GPS} = N s_0$, where $s_0 < \ln 2$ is determined by integrating

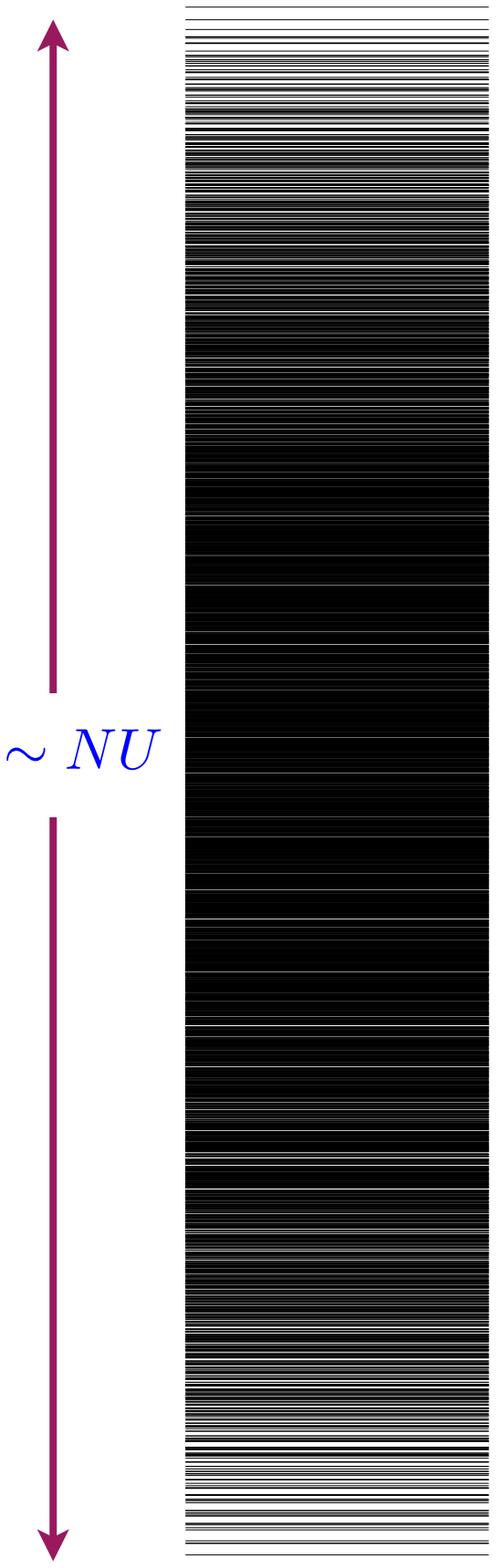
$$\frac{ds_0}{dQ} = 2\pi\mathcal{E}.$$

At $Q = 1/2$,

$$s_0 = \frac{G}{\pi} + \frac{\ln(2)}{4} = 0.464848\dots$$

where G is Catalan's constant.

GPS: A. Georges, O. Parcollet, and S. Sachdev, PRB **63**, 134406 (2001)



Many-body level spacing $\sim 2^{-N} = e^{-N \ln 2}$

Non-quasiparticle excitations with spacing $\sim e^{-N s_0}$

The SYK model

$$\Omega(T) - E_0 = N \left[-s_0 T - \frac{1}{2}(\gamma + 4\pi^2 \mathcal{E}^2 K) T^2 + \mathcal{O}(T^3) \right] + 2T \ln \left(\frac{U}{T} \right) \dots$$

is the grand potential, where $K = d\mathcal{Q}/d\mu \sim 1/U$ is the compressibility/ N , $\gamma \sim 1/U$ will appear later in the co-efficient of the Schwarzian, and the N^0 term arises from fluctuations about the large N theory described by the Schwarzian.

The inversion from $\Omega(T)$ to the *many*-body density of states, $D(E)$,

$$Z = e^{-\Omega(T)/T} = \int_{-\infty}^{\infty} dE D(E) e^{-E/T}$$

requires terms in $\Omega(T)$ which are exponentially small in N (not shown above) from the Schwarzian action, yielding terms which are not small in $D(E)$. We obtain

J. S. Cotler, G. Gur-Ari, M. Hanada, J. Polchinski, P. Saad, S. H. Shenker, D. Stanford, A. Streicher, and M. Tezuka, arXiv:1611.04650;
R. Davison, Wenbo Fu, A. Georges, Yingfei Gu, K. Jensen, S. Sachdev, arXiv:1612.00849 ;
A.M. Garcia-Garcia and J.J.M. Verbaarschot, arXiv:1701.06593; D. Bagrets, A. Altland, and A. Kamenev, arXiv:1702.08902;
D. Stanford and E. Witten, arXiv:1703.04612; A. Kitaev and S.J. Suh, arXiv:1711.08467; Yingfei Gu and S. Sachdev, unpublished.

The SYK model

$$D(E) = \sum_{p=-\infty}^{\infty} e^{2\pi p \mathcal{E}} d\left(E - \frac{p^2}{2NK}\right)$$

where $N\mathcal{Q} + p$ is the integer fermion number, $d(E) = 0$ for $E < E_0$, and

$$d(E) \sim \exp(Ns_0) \sinh\left(\sqrt{2N\gamma(E - E_0)}\right), \quad E > E_0, \quad e^{-cN} \ll \gamma(E - E_0) \ll N$$

There are exponentially more low energy states than for the quasiparticle case, and $D(E)$ self-averages down to energies exponentially small in N .

We can understand the dependence on the integer charge p by the relationship $ds_0/d\mathcal{Q} = 2\pi\mathcal{E}$, and hence $Ns_0(\mathcal{Q} + p/N) \approx Ns_0 + 2\pi p\mathcal{E}$.

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where $\rho_0 \equiv \rho(0)$ is the *single* particle density of states at the Fermi level.

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$$D(E) \sim \exp \left(\pi \sqrt{\frac{2N\rho_0(E - E_0)}{3}} \right), \quad E > E_0, \quad \frac{1}{N} \ll \rho_0(E - E_0) \ll N$$

and $D(E) = 0$ for $E < E_0$. This is related to the asymptotic growth of the partitions of an integer, $p(n) \sim \exp(\pi\sqrt{2n/3})$. Near the lower bound, there are large sample-to-sample fluctuations due to variations in the lowest quasiparticle energies.

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$$d(E) \sim \exp(Ns_0) \sinh\left(\sqrt{2N\gamma(E - E_0)}\right), \quad E > E_0, \quad e^{-cN} \ll \gamma(E - E_0) \ll N$$

There are exponentially more low energy states than for the quasiparticle case, and $D(E)$ self-averages down to energies exponentially small in N .

We can understand the dependence on the integer charge p by the relationship $ds_0/d\mathcal{Q} = 2\pi\mathcal{E}$, and hence $Ns_0(\mathcal{Q} + p/N) \approx Ns_0 + 2\pi p\mathcal{E}$.

The SYK model

No quasiparticles

- Rapid local thermal equilibration (of fermion correlators) in a ‘Planckian’ time

$$\tau_{\text{eq}} \sim \frac{\hbar}{k_B T} \quad , \quad \text{as } T \rightarrow 0.$$

A. Georges and O. Parcollet
PRB **59**, 5341 (1999)

A. Eberlein, V. Kasper, S. Sachdev, and
J. Steinberg, PRB **96**, 205123 (2017)

Established by solution of Schwinger-Keldysh equations for a quench.

- Presence of quasiparticles should slow down thermalization, so *all* quantum systems obey

$$\tau_{\text{eq}} > C \frac{\hbar}{k_B T} \quad , \quad \text{as } T \rightarrow 0.$$

S. Sachdev, *Quantum Phase Transitions*,
Cambridge (1999)

Absence of quasiparticles \Leftrightarrow Fastest possible thermalization

1. Random matrix quasiparticle model

$q=2$, complex SYK

2. Matter without quasiparticles

$q=4$, complex SYK

3. The Schwarzian theory

4. Connections to black holes
with AdS_2 horizons

The SYK model

Feynman graph expansion in $J_{ij..}$, and graph-by-graph average, yields exact equations in the large N limit:

$$G(i\omega) = \frac{1}{i\omega + \mu - \Sigma(i\omega)} \quad , \quad \Sigma(\tau) = -U^2 G^2(\tau) G(-\tau)$$
$$G(\tau = 0^-) = Q.$$

Low frequency analysis shows that the solutions must be gapless and obey

$$\Sigma(z) = \mu - \frac{1}{A} \sqrt{z} + \dots \quad , \quad G(z) = \frac{A}{\sqrt{z}}$$

where $A = e^{-i\pi/4} (\pi/U^2)^{1/4}$ at half-filling. The ground state is a non-Fermi liquid, with a continuously variable density Q .

The SYK model

The equations for the Green's function can also be solved at non-zero T . At half-filling, $Q = 1/2$, we “guess” the particle-hole symmetric solution

$$G(\tau) = B \operatorname{sgn}(\tau) \left| \frac{\pi T}{\sin(\pi T \tau)} \right|^\rho$$

Then the self-energy is

$$\Sigma(\tau) = U^2 B^3 \operatorname{sgn}(\tau) \left| \frac{\pi T}{\sin(\pi T \tau)} \right|^{3\rho}$$

A. Georges and O. Parcollet
PRB **59**, 5341 (1999)

The SYK model

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Taking Fourier transforms, we have as a function of the Matsubara frequency ω_n

$$G(i\omega_n) = [iB\Pi(\rho)] \frac{T^{\rho-1} \Gamma\left(\frac{\rho}{2} + \frac{\omega_n}{2\pi T}\right)}{\Gamma\left(1 - \frac{\rho}{2} + \frac{\omega_n}{2\pi T}\right)}$$
$$\Sigma_{\text{sing}}(i\omega_n) = [iU^2 B^3 \Pi(3\rho)] \frac{T^{3\rho-1} \Gamma\left(\frac{3\rho}{2} + \frac{\omega_n}{2\pi T}\right)}{\Gamma\left(1 - \frac{3\rho}{2} + \frac{\omega_n}{2\pi T}\right)},$$

A. Georges and O. Parcollet
PRB **59**, 5341 (1999)

The SYK model

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where we have dropped a less-singular term in Σ , and

$$\Pi(s) \equiv \pi^{s-1} 2^s \cos\left(\frac{\pi s}{2}\right) \Gamma(1-s).$$

Now the singular part of Dyson's equation is

$$G(i\omega_n) \Sigma_{\text{sing}}(i\omega_n) = -1$$

A. Georges and O. Parcollet
PRB **59**, 5341 (1999)

Remarkably, the Γ functions appear with just the right arguments, so that there is a solution of the Dyson equation at $\rho = 1/2$!

So the Green's functions display thermal 'damping' at a scale set by T alone, which is independent of U .

The SYK model

Away from half-filling, the $T = 0$ solution has the low frequency form

$$\Sigma(z) = \mu - \frac{e^{i(\pi/4+\theta)}}{A} \sqrt{z} + \dots, \quad G(z) = \frac{Ae^{-i(\pi/4+\theta)}}{\sqrt{z}}$$

where $A = (\pi/U^2 \cos(2\theta))^{1/4}$. The value of θ is universally related to Q by a Luttinger-Ward functional analysis similar to that used to establish the Luttinger theorem of Fermi liquid theory:

$$Q = \frac{1}{2} - \frac{\theta}{\pi} - \frac{\sin(2\theta)}{4}$$

At $T > 0$, we obtain a solution with a conformal structure

$$G(\tau) = -A \frac{e^{-2\pi\mathcal{E}T\tau}}{\sqrt{1 + e^{-4\pi\mathcal{E}}}} \left(\frac{T}{\sin(\pi T\tau)} \right)^{1/2}, \quad 0 < \tau < 1/T,$$

where the ‘particle-hole asymmetry’ is determined by \mathcal{E}

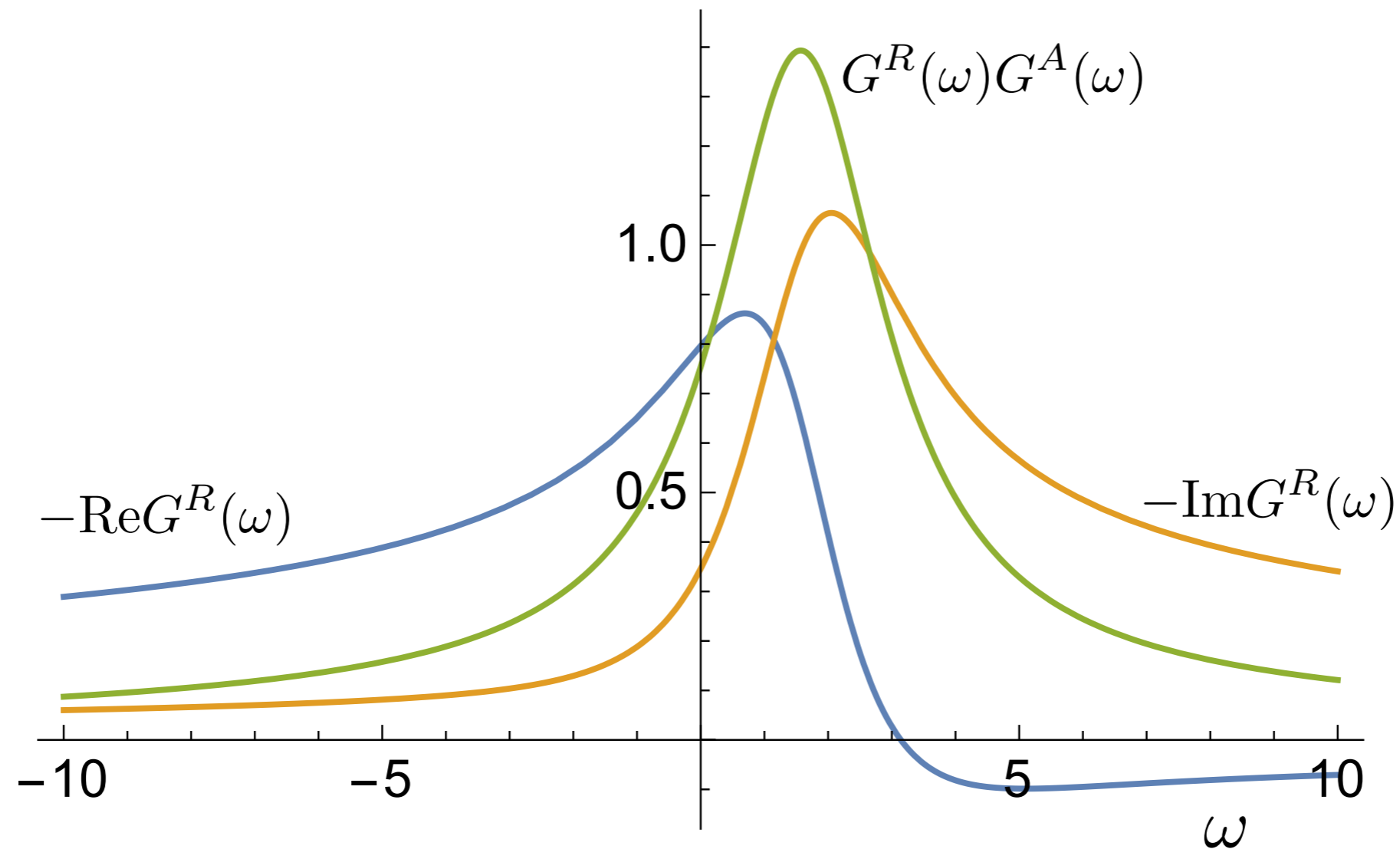
$$e^{2\pi\mathcal{E}} = \frac{\sin(\pi/4 + \theta)}{\sin(\pi/4 - \theta)}.$$

A. Georges and O. Parcollet
PRB **59**, 5341 (1999)

A. Georges, O. Parcollet, and S. Sachdev,
PRB **63**, 134406 (2001)

S. Sachdev, PRX **5**, 041025 (2015)

The SYK model



Green's functions away from half-filling

So the Green's functions display thermal 'damping' at a scale set by T alone, which is independent of U .

The SYK model

$$G(i\omega) = \frac{1}{i\omega + \mu - \Sigma(i\omega)} \quad , \quad \Sigma(\tau) = -U^2 G^2(\tau) G(-\tau)$$
$$\Sigma(z) = \mu - \frac{1}{A} \sqrt{z} + \dots \quad , \quad G(z) = \frac{A}{\sqrt{z}}$$

The SYK model

$$G(i\omega) = \frac{1}{i\omega + \mu - \Sigma(i\omega)} \quad , \quad \Sigma(\tau) = -U^2 G^2(\tau) G(-\tau)$$
$$\Sigma(z) = \mu - \frac{1}{A} \sqrt{z} + \dots \quad , \quad G(z) = \frac{A}{\sqrt{z}}$$

At frequencies $\ll U$, the $i\omega + \mu$ can be dropped, and without it equations are invariant under the reparametrization and gauge transformations.

The singular part of the self-energy and the Green's function obey

$$\int_0^\beta d\tau_2 \Sigma_{\text{sing}}(\tau_1, \tau_2) G(\tau_2, \tau_3) = -\delta(\tau_1 - \tau_3)$$

$$\Sigma_{\text{sing}}(\tau_1, \tau_2) = -U^2 G^2(\tau_1, \tau_2) G(\tau_2, \tau_1)$$

The SYK model

$$\int_0^\beta d\tau_2 \Sigma(\tau_1, \tau_2) G(\tau_2, \tau_3) = -\delta(\tau_1 - \tau_3)$$
$$\Sigma(\tau_1, \tau_2) = -U^2 G^2(\tau_1, \tau_2) G(\tau_2, \tau_1)$$

These equations are invariant under

$$\tau = f(\sigma)$$

$$G(\tau_1, \tau_2) = [f'(\sigma_1) f'(\sigma_2)]^{-1/4} \frac{g(\sigma_1)}{g(\sigma_2)} \tilde{G}(\sigma_1, \sigma_2)$$

$$\Sigma(\tau_1, \tau_2) = [f'(\sigma_1) f'(\sigma_2)]^{-3/4} \frac{g(\sigma_1)}{g(\sigma_2)} \tilde{\Sigma}(\sigma_1, \sigma_2)$$

where $f(\sigma)$ and $g(\sigma)$ are arbitrary functions.

By using $f(\sigma) = \tan(\pi T \sigma) / (\pi T)$ we can

now obtain the $T > 0$ solution from the $T = 0$ solution.

The SYK model

Let us write the large N saddle point solutions of S as

$$\begin{aligned} G_s(\tau_1 - \tau_2) &\sim (\tau_1 - \tau_2)^{-1/2} \\ \Sigma_s(\tau_1 - \tau_2) &\sim (\tau_1 - \tau_2)^{-3/2}. \end{aligned}$$

The saddle point will be invariant under a reparamaterization $f(\tau)$ when choosing $G(\tau_1, \tau_2) = G_s(\tau_1 - \tau_2)$ leads to a transformed $\tilde{G}(\sigma_1, \sigma_2) = G_s(\sigma_1 - \sigma_2)$ (and similarly for Σ). It turns out this is true only for the $SL(2, \mathbb{R})$ transformations under which

$$f(\tau) = \frac{a\tau + b}{c\tau + d}, \quad ad - bc = 1.$$

So the (approximate) reparametrization symmetry is spontaneously broken down to $SL(2, \mathbb{R})$ by the saddle point.

Basics of conformal field theory

In a space with metric tensor $g_{\mu\nu}$ and proper distance

$$ds^2 = g_{\mu\nu} dx_\mu dx_\nu$$

after the co-ordinate transformation $x_\mu \rightarrow x'_\mu$, the new metric tensor is

$$g'_{\mu\nu} = g_{\rho\lambda} \frac{\partial x_\rho}{\partial x'_\mu} \frac{\partial x_\lambda}{\partial x'_\nu}.$$

A conformal transformation is one which preserves all angles and so

$$g'_{\mu\nu}(x') = \Lambda(x) g_{\mu\nu}(x).$$

In a conformal field theory, two-point correlators of scalar fields transform as

$$\langle \phi(x_1) \phi(x_2) \rangle = \left| \det \left[\frac{\partial x'_1}{\partial x_1} \right] \right|^{\Delta/d} \left| \det \left[\frac{\partial x'_2}{\partial x_2} \right] \right|^{\Delta/d} \langle \phi(x'_1) \phi(x'_2) \rangle$$

Infinite-range (SYK) model without quasiparticles

After introducing replicas $a = 1 \dots n$, and integrating out the disorder, the partition function can be written as

$$Z = \int \mathcal{D}c_{ia}(\tau) \exp \left[- \sum_{ia} \int_0^\beta d\tau c_{ia}^\dagger \left(\frac{\partial}{\partial \tau} - \mu \right) c_{ia} - \frac{U^2}{4N^3} \sum_{ab} \int_0^\beta d\tau d\tau' \left| \sum_i c_{ia}^\dagger(\tau) c_{ib}(\tau') \right|^4 \right].$$

For simplicity, we neglect the replica indices, and introduce the identity

$$1 = \int \mathcal{D}\Sigma(\tau_1, \tau_2) \exp \left[-N \int_0^\beta d\tau_1 d\tau_2 \Sigma(\tau_1, \tau_2) \left(G(\tau_2, \tau_1) + \frac{1}{N} \sum_i c_i(\tau_2) c_i^\dagger(\tau_1) \right) \right].$$

Infinite-range (SYK) model without quasiparticles

Then the partition function can be written as a path integral with an action S analogous to a Luttinger-Ward functional

$$Z = \int \mathcal{D}G(\tau_1, \tau_2) \mathcal{D}\Sigma(\tau_1, \tau_2) \exp(-NS)$$
$$S = \ln \det [\delta(\tau_1 - \tau_2)(\partial_{\tau_1} + \mu) - \Sigma(\tau_1, \tau_2)]$$
$$+ \int d\tau_1 d\tau_2 \Sigma(\tau_1, \tau_2) [G(\tau_2, \tau_1) + (U^2/2)G^2(\tau_2, \tau_1)G^2(\tau_1, \tau_2)]$$

At frequencies $\ll U$, the time derivative in the determinant is less important, and without it the path integral is invariant under the reparametrization and gauge transformations

$$\tau = f(\sigma)$$

$$G(\tau_1, \tau_2) = [f'(\sigma_1)f'(\sigma_2)]^{-1/4} \frac{g(\sigma_1)}{g(\sigma_2)} G(\sigma_1, \sigma_2)$$

$$\Sigma(\tau_1, \tau_2) = [f'(\sigma_1)f'(\sigma_2)]^{-3/4} \frac{g(\sigma_1)}{g(\sigma_2)} \Sigma(\sigma_1, \sigma_2)$$

where $f(\sigma)$ and $g(\sigma)$ are arbitrary functions.

A. Georges and O. Parcollet
PRB **59**, 5341 (1999)

A. Kitaev, 2015

S. Sachdev, PRX **5**, 041025 (2015)

The SYK model

Reparametrization and phase zero modes

We can write the path integral for the SYK model as

$$\mathcal{Z} = \int \mathcal{D}G(\tau_1, \tau_1) \mathcal{D}\Sigma(\tau_1, \tau_2) e^{-NS[G, \Sigma]}$$

for a known action $S[G, \Sigma]$. We find the saddle point, G_s, Σ_s , and only focus on the “Nambu-Goldstone” modes associated with breaking reparameterization and U(1) gauge symmetries by writing

$$G(\tau_1, \tau_2) = [f'(\tau_1)f'(\tau_2)]^{1/4} G_s(f(\tau_1) - f(\tau_2)) e^{i\phi(\tau_1) - i\phi(\tau_2)}$$

(and similarly for Σ). Then the path integral is approximated by

$$\mathcal{Z} = \int \mathcal{D}f(\tau) \mathcal{D}\phi(\tau) e^{-NS_{\text{eff}}[f, \phi]}.$$

J. Maldacena and D. Stanford, arXiv:1604.07818;

R. Davison, Wenbo Fu, A. Georges, Yingfei Gu, K. Jensen, S. Sachdev, arXiv:1612.00849;

S. Sachdev, PRX **5**, 041025 (2015); J. Maldacena, D. Stanford, and Zhenbin Yang, arXiv:1606.01857;

K. Jensen, arXiv:1605.06098; J. Engelsoy, T.G. Mertens, and H. Verlinde, arXiv:1606.03438

The Schwarzian theory of the SYK model

Symmetry arguments, and explicit computations, show that the effective action is

$$S_{\text{eff}}[f, \phi] = \frac{K}{2} \int_0^{1/T} d\tau (\partial_\tau \phi + i(2\pi\mathcal{E}T)\partial_\tau f)^2 - \frac{\gamma}{4\pi^2} \int_0^{1/T} d\tau \{ \tan(\pi T f(\tau)), \tau \},$$

where $f(\tau)$ is a monotonic map from $[0, 1/T]$ to $[0, 1/T]$, the couplings K , γ , and \mathcal{E} can be related to thermodynamic derivatives and we have used the Schwarzian:

$$\{g, \tau\} \equiv \frac{g'''}{g'} - \frac{3}{2} \left(\frac{g''}{g'} \right)^2.$$

Specifically, an argument constraining the effective at $T = 0$ is

$$S_{\text{eff}} \left[f(\tau) = \frac{a\tau + b}{c\tau + d}, \phi(\tau) = 0 \right] = 0,$$

and this is origin of the Schwarzian.

J. Maldacena and D. Stanford, arXiv:1604.07818;
R. Davison, Wenbo Fu, A. Georges, Yingfei Gu, K. Jensen, S. Sachdev, arXiv:1612.00849;
A. Gaikwad, L.K. Joshi, G. Mandal, and S.R. Wadia, arXiv:1802.07746;
Yingfei Gu and S. Sachdev, unpublished

1. Random matrix quasiparticle model

$q=2$, complex SYK

2. Matter without quasiparticles

$q=4$, complex SYK

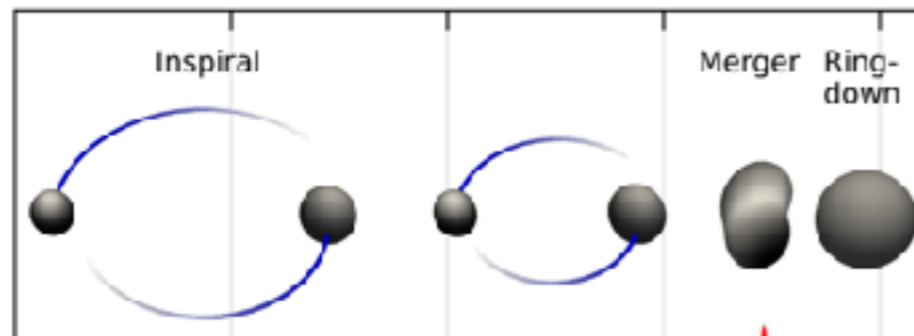
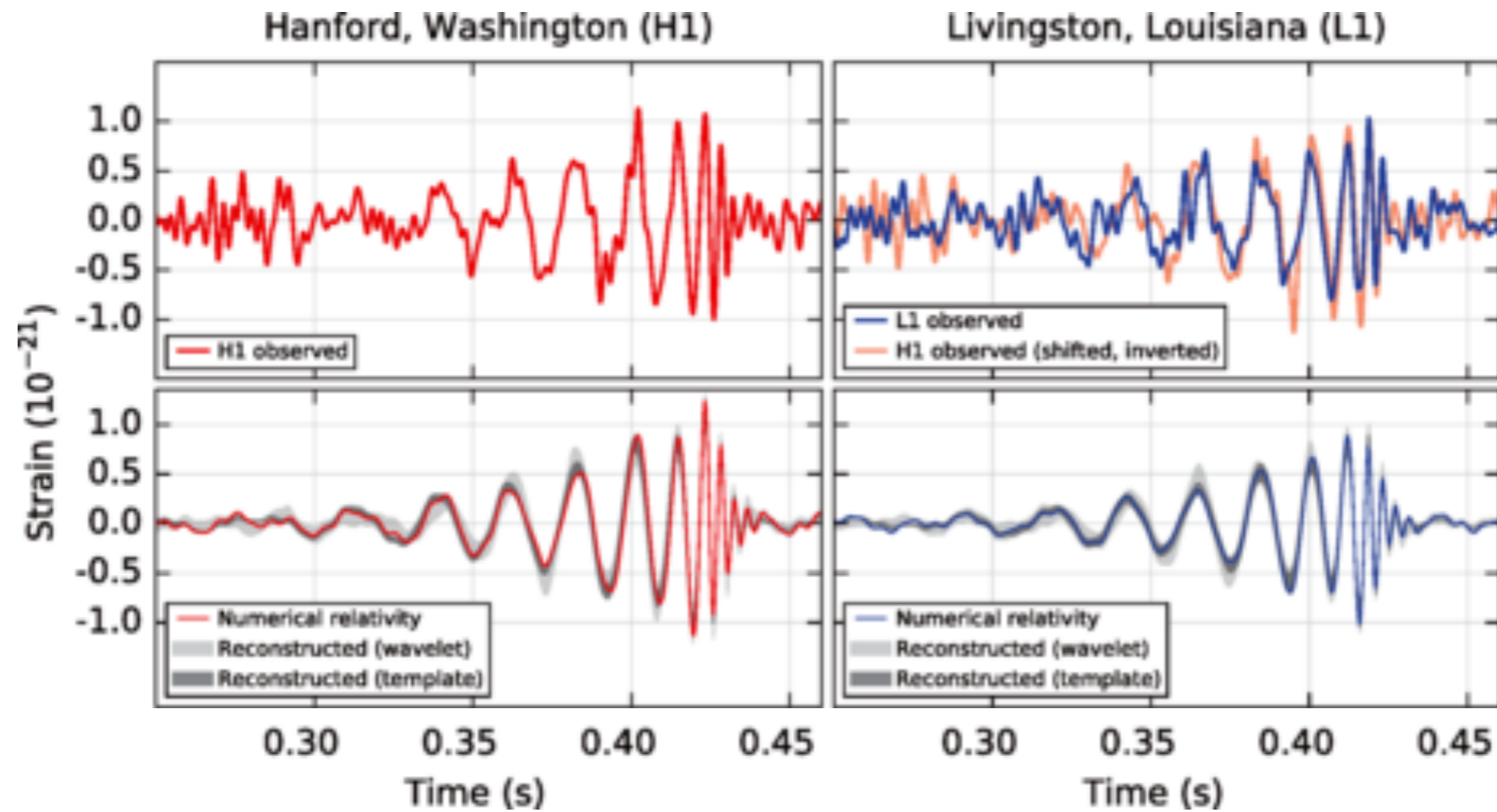
3. The Schwarzian theory

4. Connections to black holes
with AdS_2 horizons

Black holes

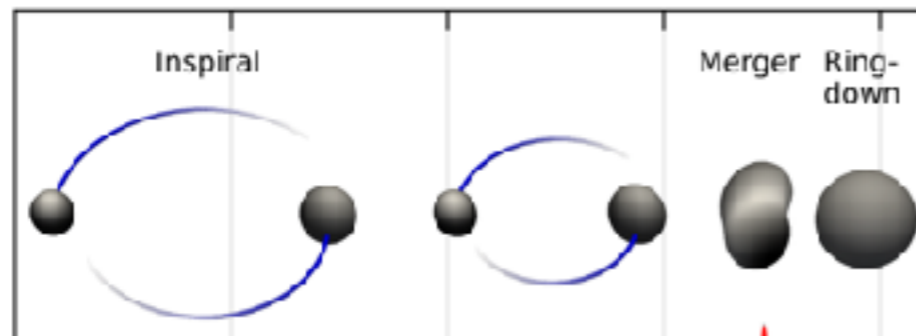
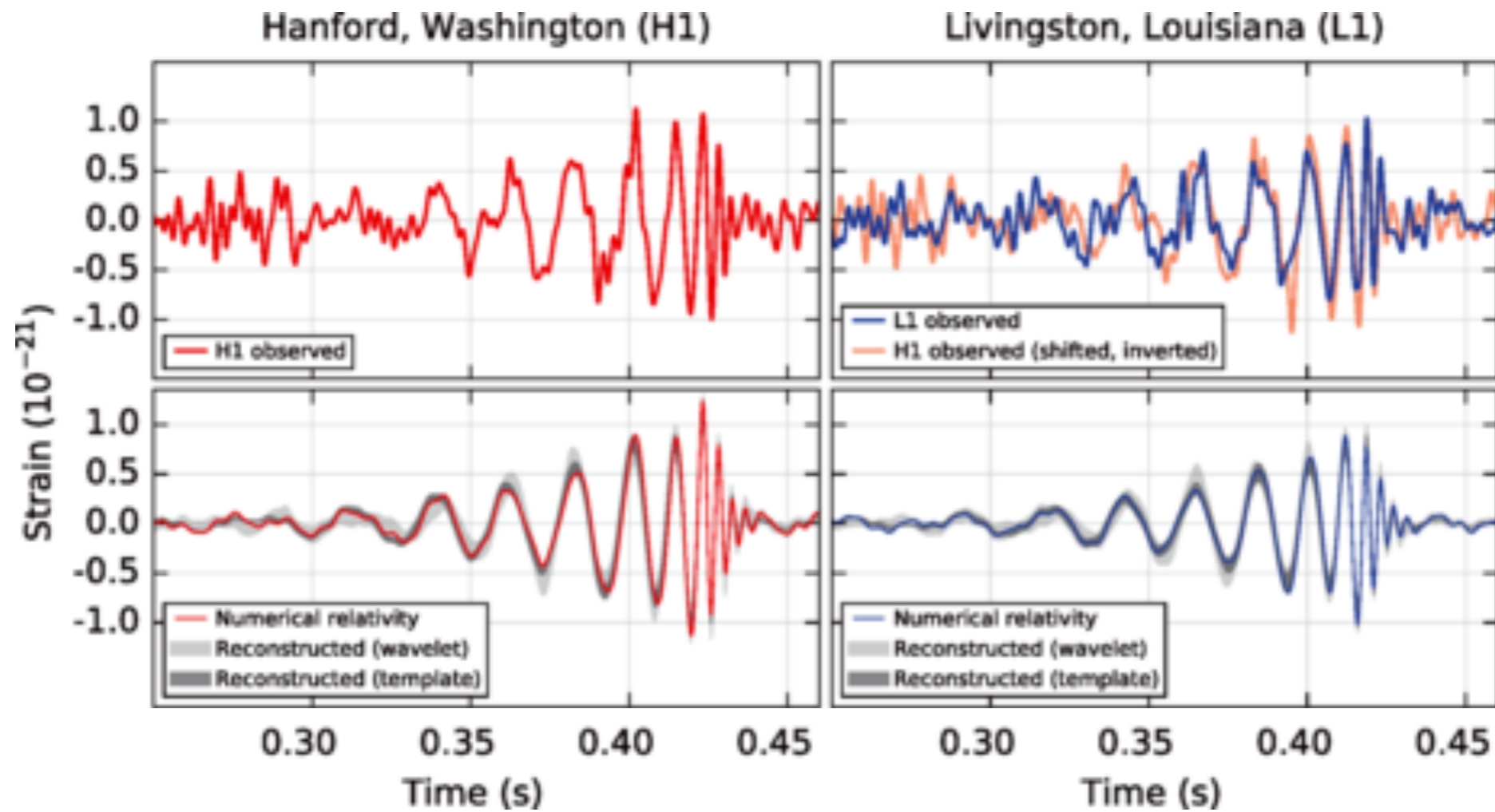
- Black holes have an entropy and a temperature, $T_H = \hbar c^3 / (8\pi G M k_B)$.
- The entropy is proportional to their surface area.





LIGO
September 14, 2015

- The ring-down is predicted by General Relativity to happen in a time $\frac{8\pi GM}{c^3} \sim 8$ milliseconds.



LIGO
September 14, 2015

- The ring-down is predicted by General Relativity to happen in a time $\frac{8\pi GM}{c^3} \sim 8$ milliseconds. Curiously this happens to equal $\frac{\hbar}{k_B T_H}$; so the ring down can also be viewed as the approach of a quantum system to thermal equilibrium at the fastest possible rate!

Black holes

- Black holes have an entropy and a temperature, $T_H = \hbar c^3 / (8\pi G M k_B)$.
- The entropy is proportional to their surface area.
- They relax to thermal equilibrium in a time $\sim \hbar / (k_B T_H)$.

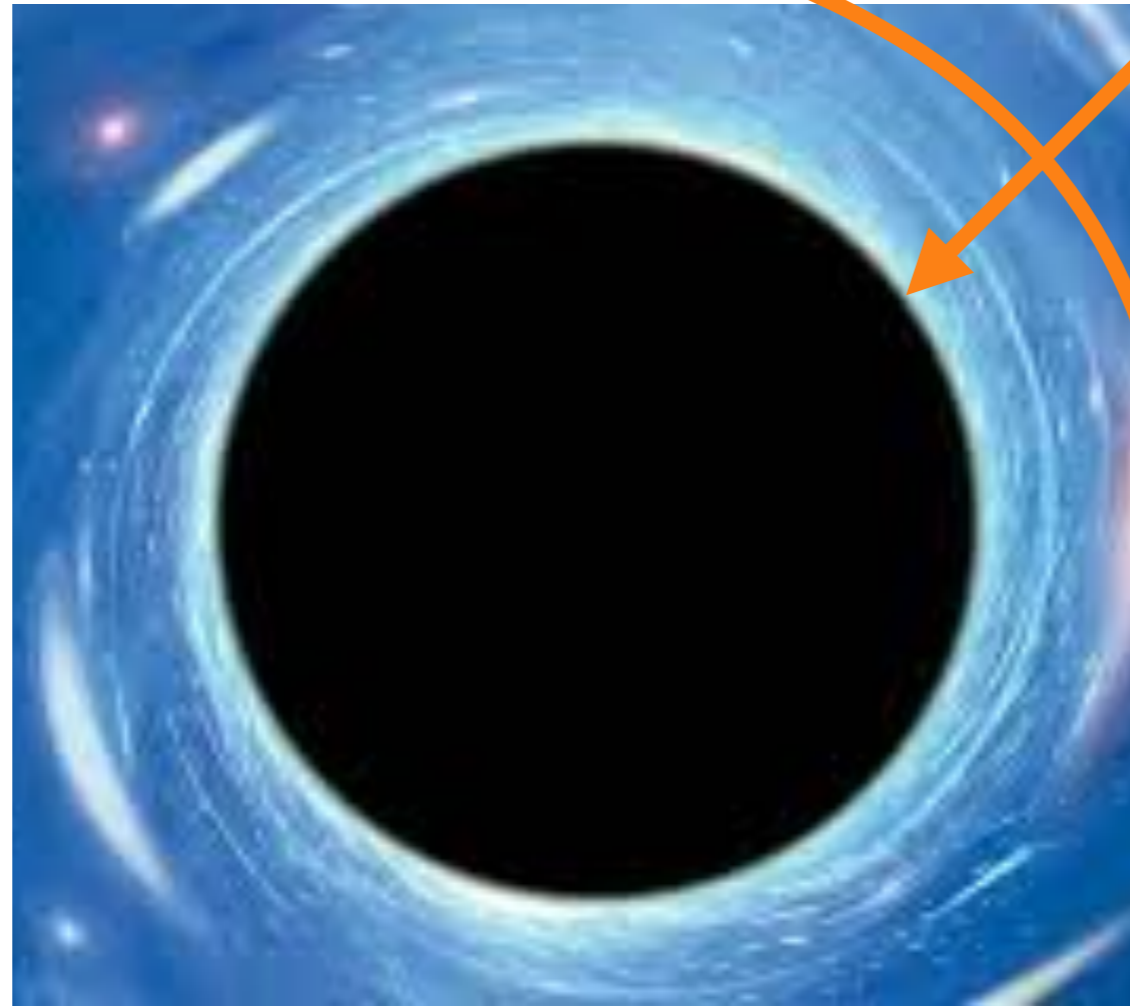


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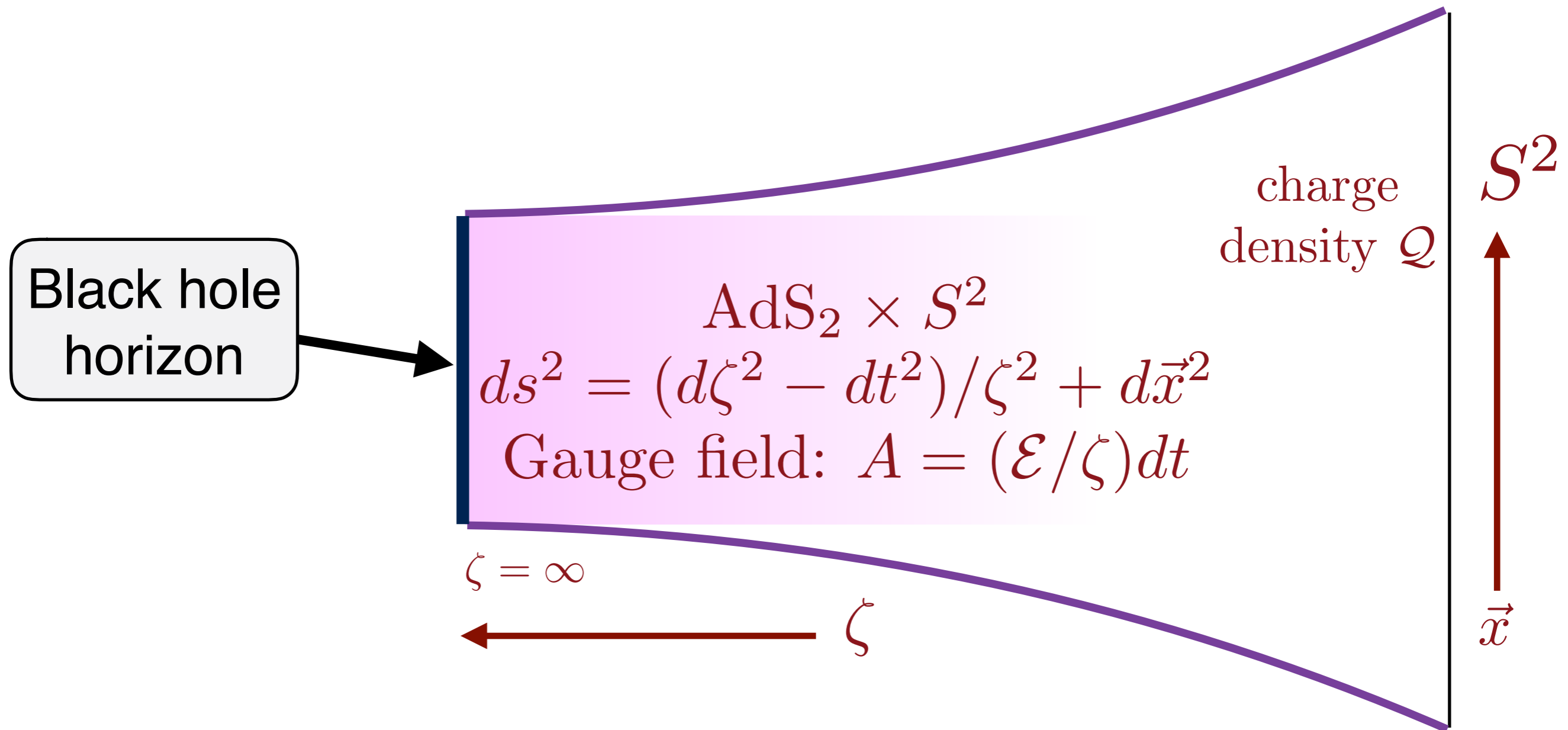
Holography:

Quantum black holes “look like” quantum many-particle systems without quasiparticle excitations, residing “on” the surface of the black hole



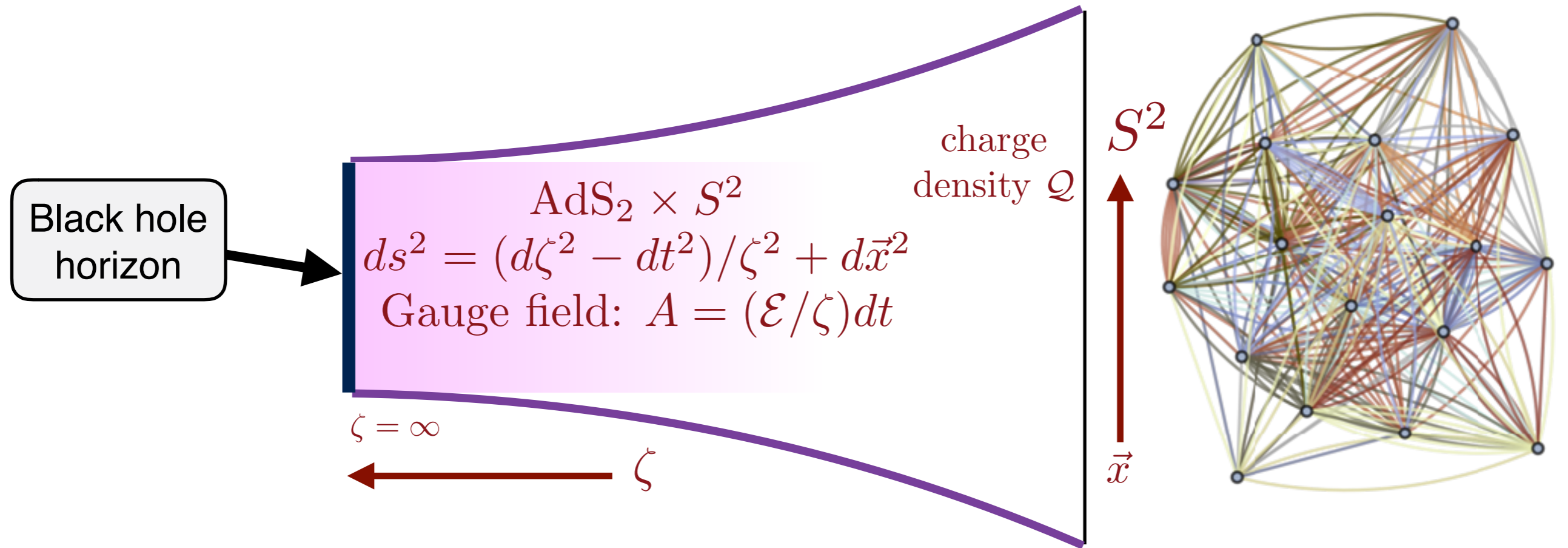
Consider a charged black hole with the smallest possible mass: the extremal limit. Zoom in to the near-horizon region at low energies. In this limit, the quantum theory lives in one space (ζ) and one time dimension

SYK models and black holes



The near-horizon region of an extremal charged black hole has the geometry of (1+1)-dimensional anti-de Sitter spacetime. By holography, this should map to a zero-dimensional quantum system: this turns out to be the SYK model

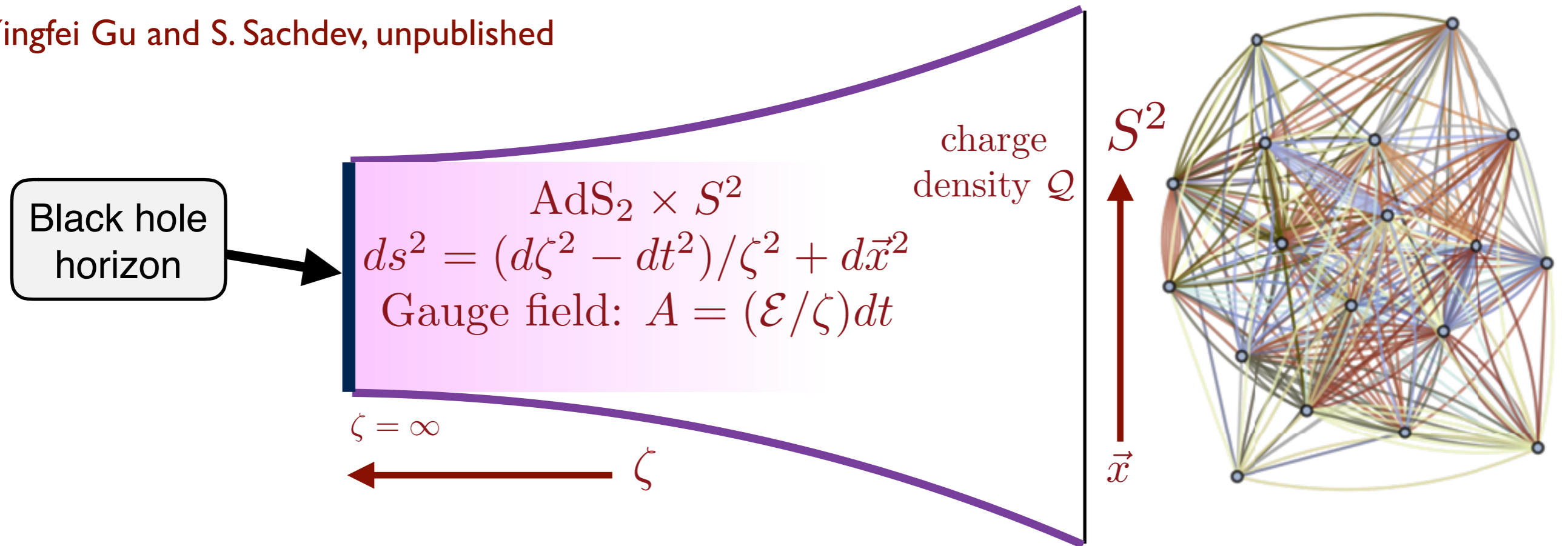
SYK models and black holes



Bekenstein-Hawking entropy of AdS_2 horizon
at $T = 0 \Leftrightarrow N s_0$ entropy of SYK model

SYK models and black holes

Yingfei Gu and S. Sachdev, unpublished



The correspondence between the complex SYK model and extremal black holes holds also for the low T thermodynamics and low energy density of states. Both obey

$$\Omega(T) - E_0 = N \left[-s_0 T - \frac{1}{2} (\gamma + 4\pi^2 \mathcal{E}^2 K) T^2 + \mathcal{O}(T^3) \right] + 2T \ln \left(\frac{U}{T} \right) \dots$$

for the grand potential, and for the density of states at a fixed charge Q

$$d(E) \sim \exp(Ns_0) \sinh \left(\sqrt{2N\gamma(E - E_0)} \right) \quad , \quad E > E_0 \quad , \quad e^{-cN} \ll \gamma(E - E_0) \ll N$$

with the relation $\frac{ds_0}{dQ} = 2\pi\mathcal{E}$ also obtained from Einstein's equations

SYK models and black holes

- Reparameterization invariance is a defining property of Einstein's theory of gravity
- In imaginary time, AdS_2 is the homogeneous hyperbolic space: two-dimensional surface of constant negative curvature. Its metric is invariant under $SL(2, \mathbb{R})$

$ds^2 = (d\tau^2 + d\zeta^2)/\zeta^2$ is invariant under

$$\tau' + i\zeta' = \frac{a(\tau + i\zeta) + b}{c(\tau + i\zeta) + d} \text{ with } ad - bc = 1.$$



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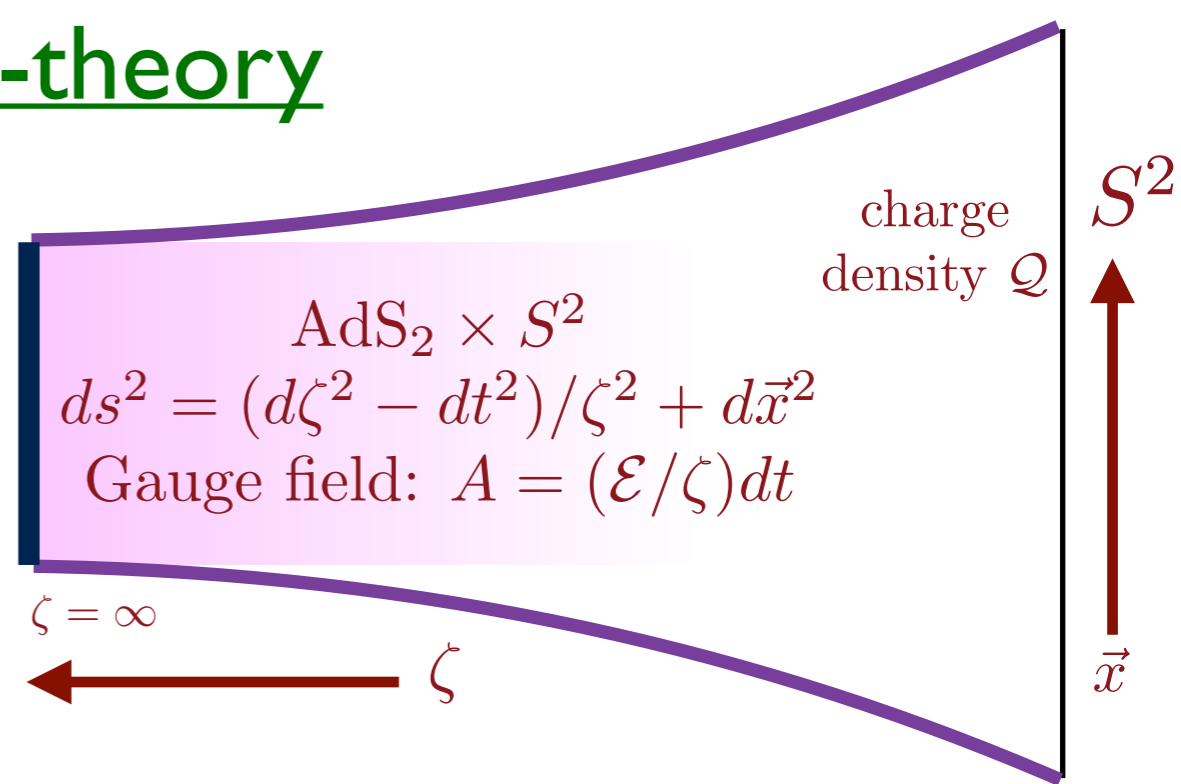
Their identical symmetries lead to the same low energy quantum theory for the SYK model and extremal charged black holes !





Einstein-Maxwell-theory

$$S_{4D} = \int d^4x \sqrt{-\hat{g}} \left(\hat{\mathcal{R}} + 6/L^2 - \frac{1}{4} \hat{F}_{\mu\nu} \hat{F}^{\mu\nu} \right),$$

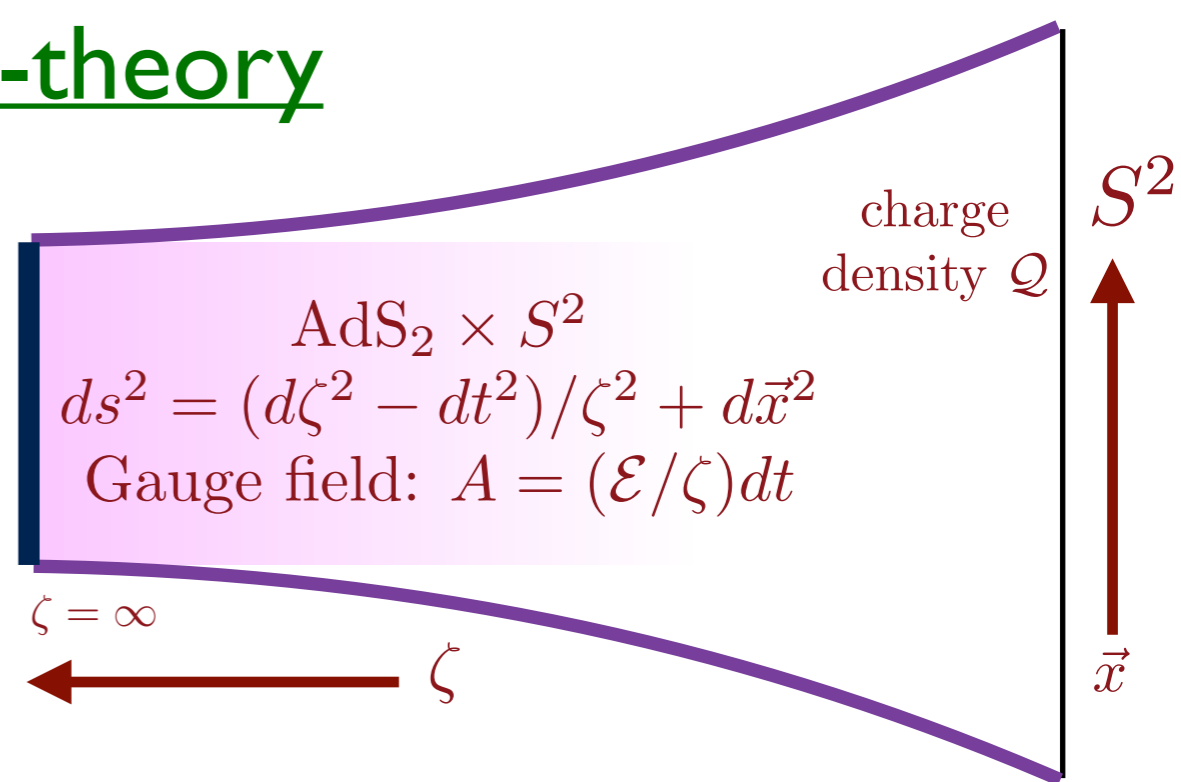


- Has Reissner-Nördstrom-AdS charged black hole solution, with charge density \mathcal{Q} , a near-horizon $\text{AdS}_2 \times S^2$ geometry, and surface electric field \mathcal{E} . (This analysis also applies in asymptotically Minkowski spacetime ($L \rightarrow \infty$) provided the black hole mass is extremal.)



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- Has Reissner-Nördstrom-AdS charged black hole solution, with charge density Q , a near-horizon $AdS_2 \times S^2$ geometry, and surface electric field \mathcal{E} . (This analysis also applies in asymptotically Minkowski spacetime ($L \rightarrow \infty$) provided the black hole mass is extremal.)
- From Einstein's equations, the Bekenstein-Hawking black hole entropy S_{4D} is found to obey the same relation as the entropy of the SYK model

$$\frac{\partial S_{4D}}{\partial Q} = 2\pi\mathcal{E},$$

A Sen, JHEP **0509**, 038 (2005)

where \mathcal{E} is identified from the spectral asymmetry of probe particle Green's functions in both cases. This establishes that the SYK entropy Ns_0 maps onto (Area of horizon)/(4G)

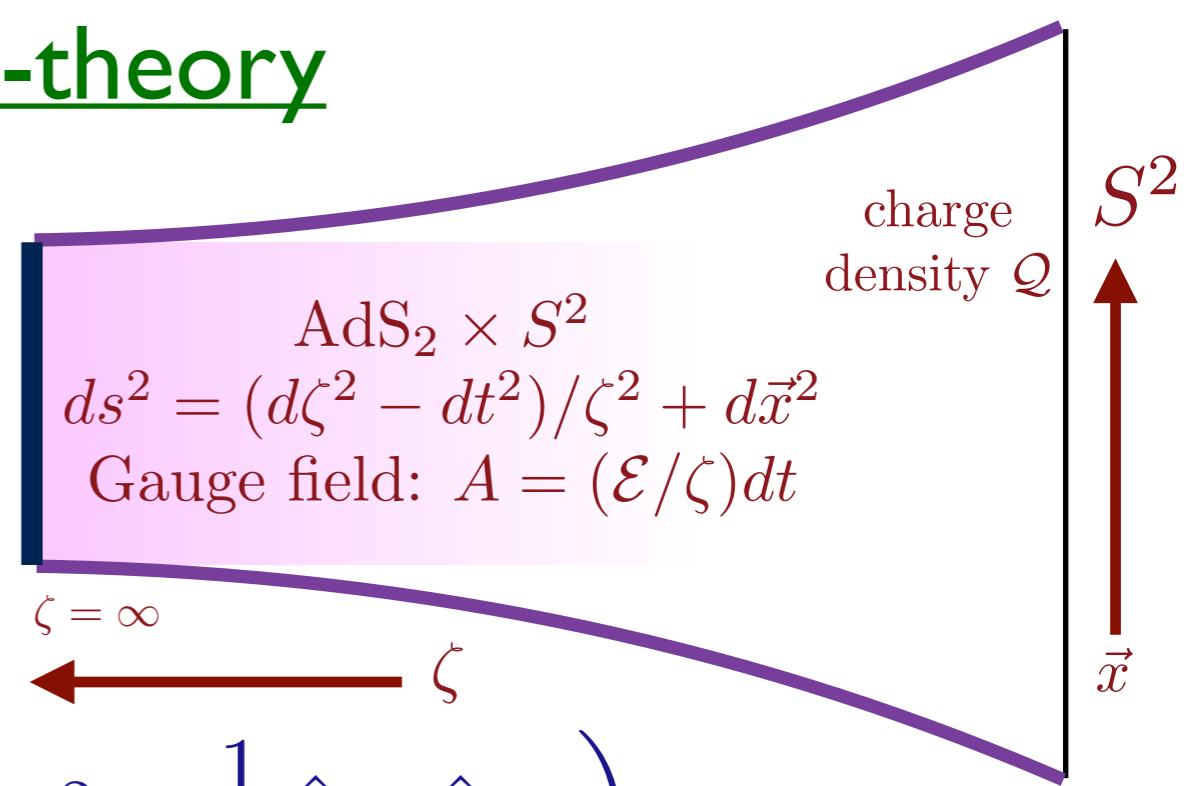
S. Sachdev, PRX **5**, 041025 (2015)



Einstein-Maxwell-theory

P. Nayak, A. Shukla, R.M. Soni, S.P. Trivedi, and V. Vishal,
arXiv:1802.09547;

A. Gaikwad, L.K. Joshi, G. Mandal, and S.R. Wadia,
arXiv:1802.07746



$$S_{4D} = \int d^4x \sqrt{-\hat{g}} \left(\hat{\mathcal{R}} + 6/L^2 - \frac{1}{4} \hat{F}_{\mu\nu} \hat{F}^{\mu\nu} \right),$$

In the small black hole size limit, $T \ll 1/R$, where R is the radius of the black hole, the theory dimensionally reduces to an Einstein-Maxwell-dilaton theory in two dimensions (the Jackiw-Teitelbaum model), along with Maxwell term

$$S_{2D} = N s_0 + \int d^2x \sqrt{-g} \left(\Phi (\mathcal{R} - \Lambda) - \frac{Z(\Phi)}{4} F_{ab} F^{ab} \right).$$

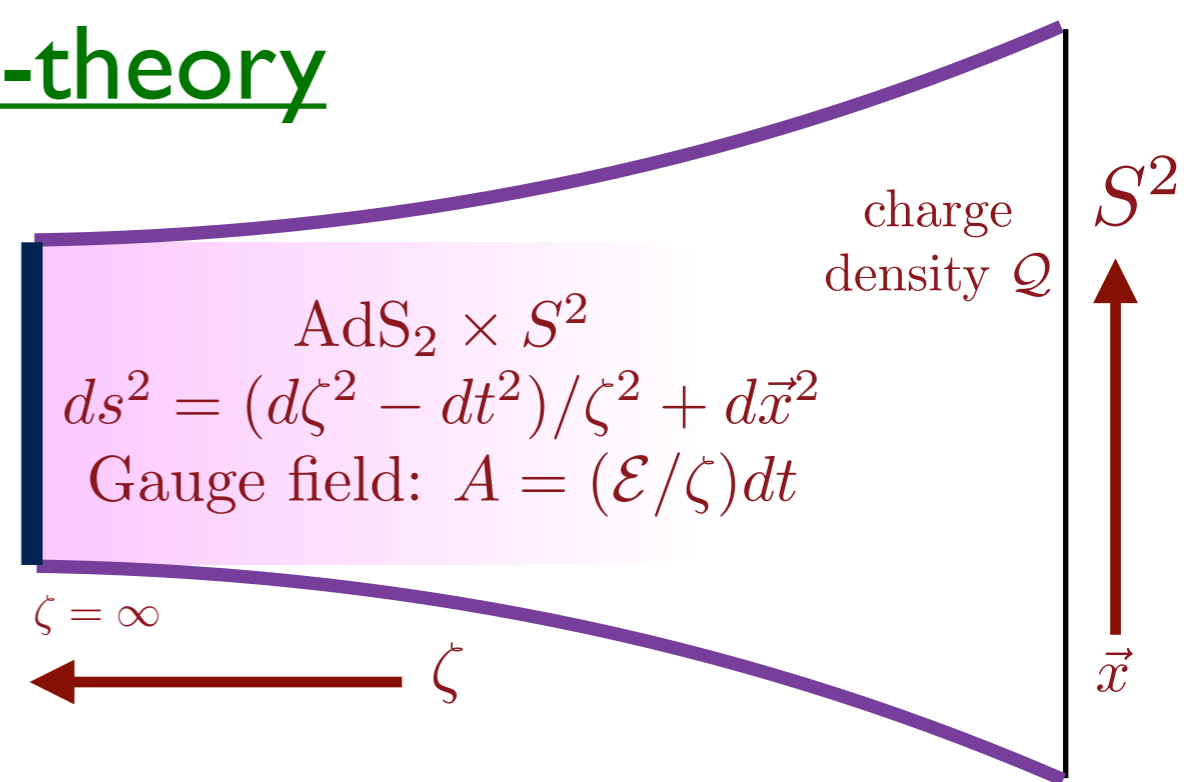
The dilaton Φ represents the radial oscillations of the small black hole.



Einstein-Maxwell-theory

P. Nayak, A. Shukla, R.M. Soni, S.P. Trivedi, and V. Vishal,
arXiv:1802.09547;

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$$S_{2D} = N s_0 + \int d^2 x \sqrt{-g} \left(\Phi (\mathcal{R} - \Lambda) - \frac{Z(\Phi)}{4} F_{ab} F^{ab} \right).$$

There are no bulk quantum fluctuations of the metric in two-dimensional gravity, and there a further dimensional reduction to a $0 + 1$ dimensional theory representing fluctuations of the AdS_2 boundary: this $0 + 1$ dimensional turns out to be *precisely the Schwarzian theory obtained for the SYK model.*

$$S_{\text{eff}}[f, \phi] = \frac{K}{2} \int_0^{1/T} d\tau (\partial_\tau \phi + i(2\pi \mathcal{E} T) \partial_\tau f)^2 - \frac{\gamma}{4\pi^2} \int_0^{1/T} d\tau \{ \tan(\pi T f(\tau)), \tau \},$$

J. Maldacena, D. Stanford, and Zhenbin Yang, arXiv:1606.01857;

K. Jensen, arXiv:1605.06098;

J. Engelsoy, T.G. Mertens, and H. Verlinde, arXiv:1606.03438

Quantum matter without quasiparticles

- Planckian dynamics is realized in the ‘solvable’ SYK models
- Black holes thermalize in a time $\sim \hbar/(k_B T_H)$, where T_H is the Hawking temperature.
- A Schwarzian theory of a time reparameterization mode, with $SL(2, \mathbb{R})$ symmetry, describes the quantum dynamics of
 - the SYK models
 - black holes with near-extremal AdS_2 horizons